

Multiple timescale simulations of global macroscopic dynamics of magnetized plasma

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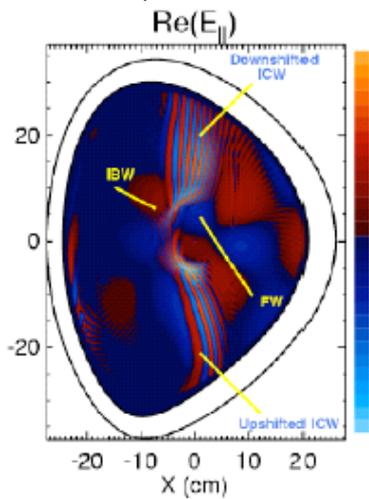
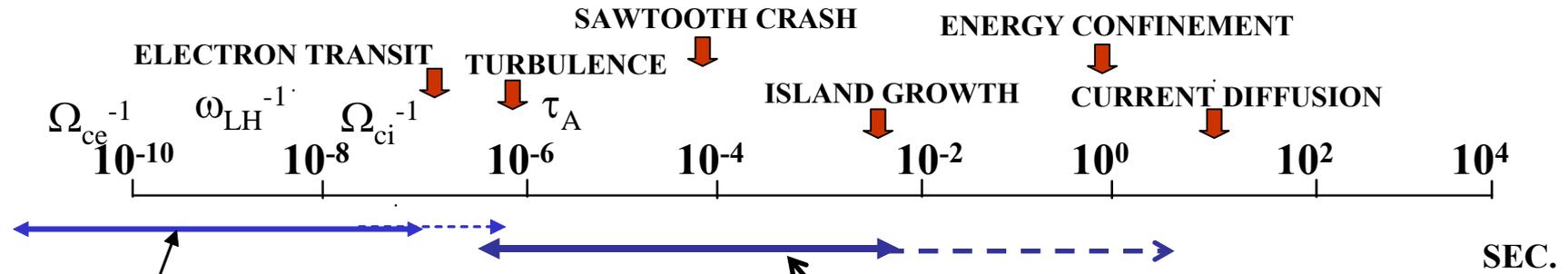
IPAM Workshop
Computational Challenges in Magnetized Plasma

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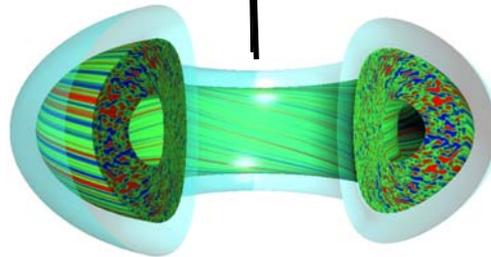
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Acknowledgements also to: G. Fu, S. Hudson, H. Strauss, L. Sugiyama

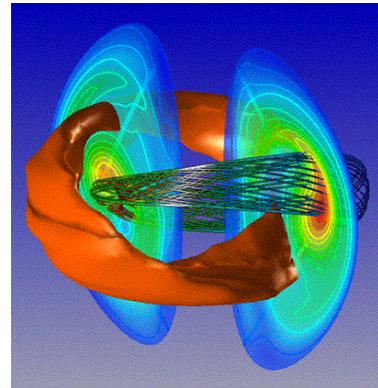
For magnetic confinement, there are 4 classes of major simulation codes, each addressing different phenomena



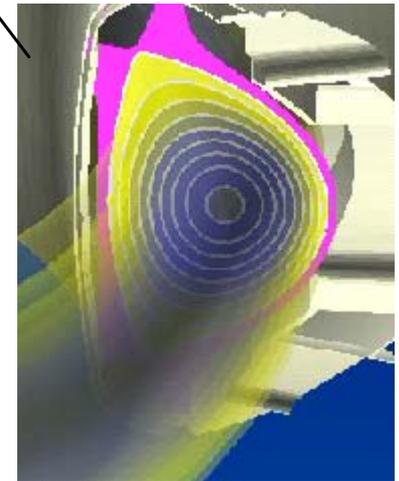
(a) RF codes



(b) Micro-turbulence codes



(c) Extended-MHD codes



(d) Transport Codes

Summary and Overview:

- 3D MHD equations are a mixed system: hyperbolic + parabolic
 - This leads to multiple timescales in a magnetized plasma
- The hyperbolic terms are associated with *ideal MHD* wave propagation and global instabilities.
 - These are the shortest timescales: typically micro-seconds
- The parabolic terms are associated with *diffusion and transport* of the magnetic field, current, pressures, and densities
 - These are the longest timescales: typically 100s of milliseconds
- To calculate both phenomena in a single simulation requires a highly *implicit* formulation so that the time step is determined by accuracy requirements only
 - not by numerical stability requirements such as Courant condition
- The implicit solution procedure is complicated by the fact that the multiple timescales present in the physics lead to a very *ill-conditioned* matrix equation that needs to be solved each time step.
 - Here we describe the techniques we use to deal with this in M3D-C¹

2-Fluid 3D MHD Equations:

$$\frac{\partial n}{\partial t} + \nabla \cdot (n\mathbf{V}) = 0 \quad \text{continuity}$$

$$\frac{\partial \mathbf{B}}{\partial t} = -\nabla \times \mathbf{E} \quad \nabla \cdot \mathbf{B} = 0 \quad \mu_0 \mathbf{J} = \nabla \times \mathbf{B} \quad \text{Maxwell}$$

$$nM_i \left(\frac{\partial \mathbf{V}}{\partial t} + \mathbf{V} \cdot \nabla \mathbf{V} \right) + \nabla p = \mathbf{J} \times \mathbf{B} - \nabla \cdot \mathbf{\Pi}_{GV} - \nabla \cdot \mathbf{\Pi}_{\mu} \quad \text{momentum}$$

$$\mathbf{E} + \mathbf{V} \times \mathbf{B} = \eta \mathbf{J} + \frac{1}{ne} (\mathbf{J} \times \mathbf{B} - \nabla p_e - \nabla \cdot \mathbf{\Pi}_e) \quad \text{Ohm's law}$$

$$\frac{3}{2} \frac{\partial p_e}{\partial t} + \nabla \cdot \left(\frac{3}{2} p_e \mathbf{V} \right) = -p_e \nabla \cdot \mathbf{V} + \eta J^2 - \nabla \cdot \mathbf{q}_e + Q_{\Delta} \quad \text{electron energy}$$

$$\frac{3}{2} \frac{\partial p_i}{\partial t} + \nabla \cdot \left(\frac{3}{2} p_i \mathbf{V} \right) = -p_i \nabla \cdot \mathbf{V} - \mathbf{\Pi}_{\mu} \cdot \nabla \mathbf{V} - \nabla \cdot \mathbf{q}_i - Q_{\Delta} \quad \text{ion energy}$$

Ideal MHD

Resistive MHD

2-fluid MHD

The objective of the **M3D-C'** project is to solve these equations as accurately as possible in 3D toroidal geometry with realistic B.C. and optimized for a low- β torus with a strong toroidal field.

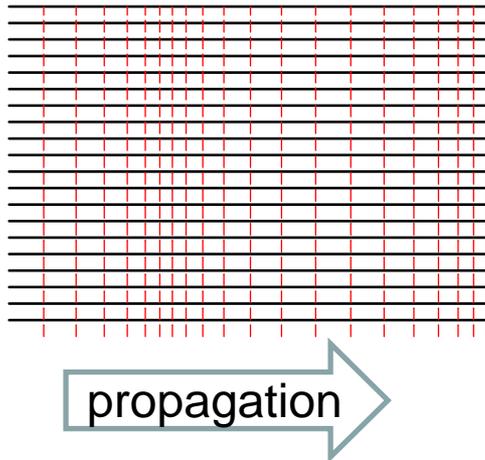
Contain ideal MHD, reconnection, and transport timescales

$$\tau_I \ll \tau_R \ll \tau_T$$

Three types of wave solutions in ideal MHD

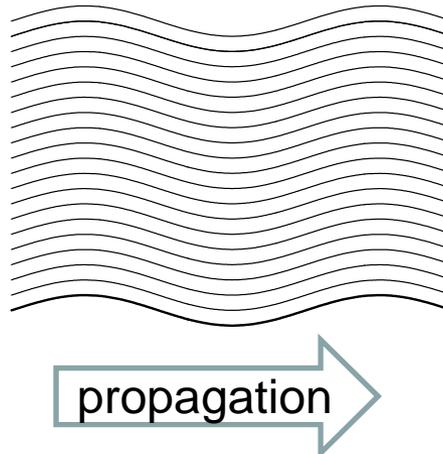


Slow Wave



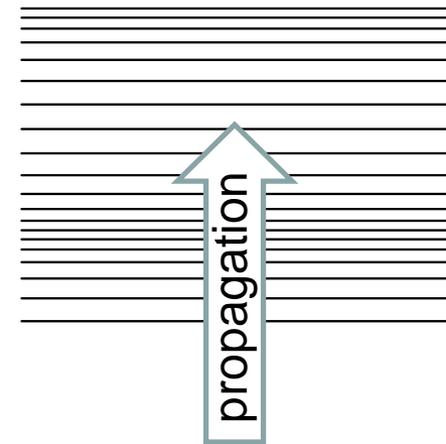
- only propagates parallel to \mathbf{B}
- only compresses fluid in parallel direction
- does not perturb magnetic field

Alfven Wave



- only propagates parallel to \mathbf{B}
- incompressible
- only bends the field, does not compress it

Fast Wave

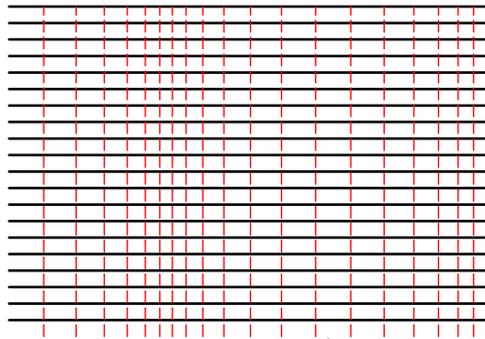


- can propagate perpendicular to \mathbf{B}
- only compresses fluid in \perp direction
- compresses the magnetic field

Three types of wave solutions in ideal MHD

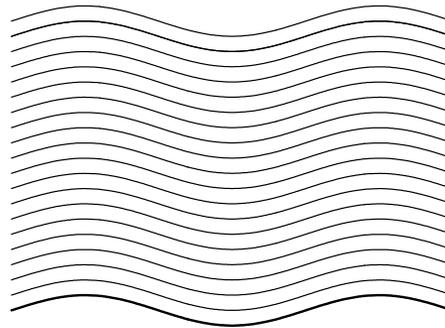


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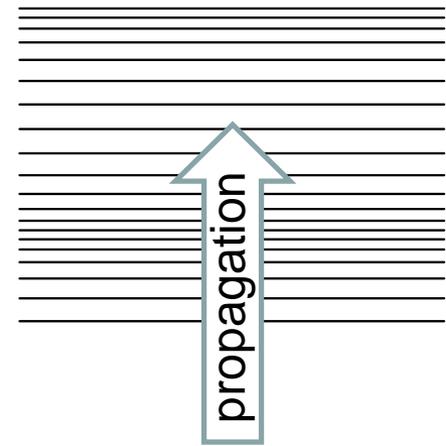
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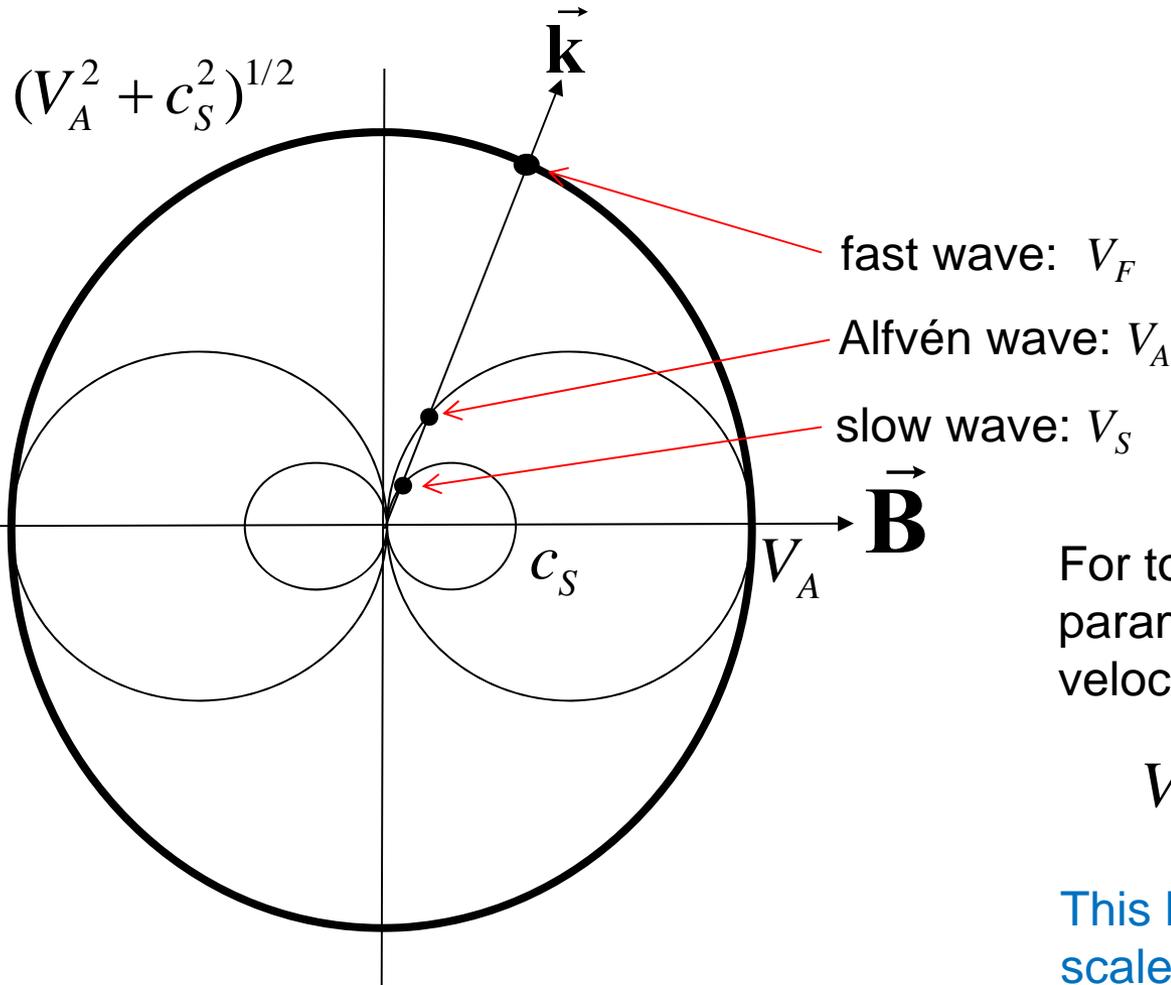


- can propagate perpendicular to \mathbf{B}
- only compresses fluid in \perp direction
- compresses the magnetic field

Plasma instabilities grow out of these two waves

•This is the wave that makes equations stiff!

The three ideal MHD waves have widely separate velocities for propagation with $\mathbf{k} \cdot \mathbf{B} \sim 0$



For tokamak geometry and parameters, the three wave velocities satisfy the inequalities:

$$V_F \gg V_A \gg V_S$$

This leads to multiple time-scales, even within ideal MHD

Wave speed diagram for ideal MHD. Intersection points show wave velocity for given propagation direction.

Implicit solution requires evaluating the spatial derivatives at the new time level.

The advantage of an implicit solution is that the time step can be very large and still be numerically stable (no Courant condition)

If we discretize in space (finite difference, finite element, or spectral) and linearize the equations about the present time level, the implicit equations take the form:

$$\begin{bmatrix} \mathbf{A}_{11} & \mathbf{A}_{12} & \mathbf{A}_{13} \\ \mathbf{A}_{21} & \mathbf{A}_{22} & \mathbf{A}_{23} \\ \mathbf{A}_{31} & \mathbf{A}_{32} & \mathbf{A}_{33} \end{bmatrix} \cdot \begin{bmatrix} \mathbf{V} \\ \mathbf{B} \\ p \end{bmatrix}^{n+1} = \begin{bmatrix} \mathbf{r}_1 \\ \mathbf{r}_2 \\ \mathbf{r}_3 \end{bmatrix}^n$$

Very large, $\sim (10^7 \times 10^7)$
non-diagonally dominant,
non-symmetric, ill-conditioned sparse
matrix (contains all MHD waves)

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How best to solve this?



Preconditioned iterative method

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matrix (contains all MHD waves)

Preconditioning

...*preconditioning* is a procedure of an application of a transformation, called the *preconditioner*, that conditions a given problem into a form that is more suitable for numerical solution.*Wikipedia*

$$\mathbf{A} \bullet \mathbf{X} = \mathbf{b}$$

Left preconditioning multiplies by a matrix from the left:

$$\mathbf{P} \bullet \mathbf{A} \bullet \mathbf{X} = \mathbf{P} \bullet \mathbf{b}$$

Right preconditioning multiplies by a matrix from the right:

$$\mathbf{A} \bullet \mathbf{P} \bullet \mathbf{Y} = \mathbf{b}$$

$$\mathbf{Y} = \mathbf{P}^{-1} \bullet \mathbf{X} \quad (\text{or } \mathbf{X} = \mathbf{P} \bullet \mathbf{Y})$$

The preconditioner \mathbf{P} is chosen so that $\mathbf{P} \mathbf{A}$ or $\mathbf{A} \mathbf{P}$ has *better properties* than the original matrix \mathbf{A} . Most of the differences between the different 3D MHD codes is due to a difference in the preconditioning techniques.

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Here we use a 3-level preconditioner that is motivated by the physics of MHD phenomena in tokamaks.

More on Preconditioning

From:

L. N. Trefethen and D. Bau, III, Numerical Linear Algebra (SIAM) 1997

“In ending this book with the subject of preconditioners, we find ourselves at the philosophical center of the scientific computing of the future...

Nothing will be more central to computational science in the next century than the art of transforming a problem that appears intractable into another whose solution can be approximated rapidly.

For Krylov subspace matrix iterations, this is *preconditioning*.”

Still more on Preconditioning

“ Direct solvers are often the best option for 2-dimensional problems, but not for 3-dimensional problems.

Generally speaking, preconditioning attempts to *improve the spectral properties of the coefficient matrix*. For symmetric positive definite problems, the rate of convergence of the CG method depends on the spectral radius.

For non-symmetric problems, the situation is more complicated and the eigenvalues may not describe the convergence properties. Nevertheless, a clustered spectrum (away from 0) often results in rapid convergence, particularly when the preconditioned matrix is close to normal. “

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Make the matrix
(close to) symmetric

+

Reduce the spectral radius (using
2D direct solves if needed)

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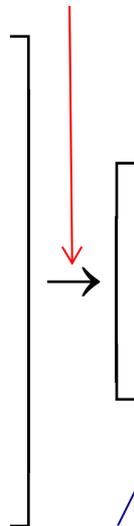
$$\rho(A) \equiv \frac{|\lambda|_{\max}}{|\lambda|_{\min}}$$

3 step physics-based preconditioner greatly improves iterative solve

(1) Split implicit formulation

Original matrix multiplying $\mathbf{V}^{n+1}, \mathbf{B}^{n+1}, \mathbf{p}^{n+1}$

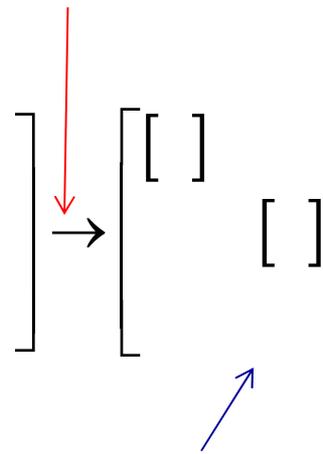
- non-symmetric,
- non-diagonally dominant &
- large range of eigenvalues



Smaller matrix multiplying \mathbf{V}^{n+1} only,

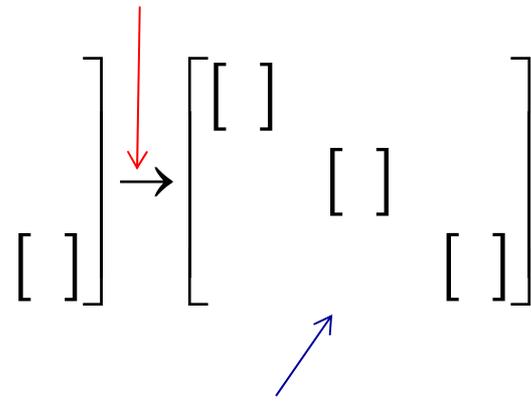
- nearly symmetric
- closer to diagonal
- still with large range of eigenvalues

(2) Apply annihilation operators



Matrix now consists of 3 dominant diagonal blocks, each with narrower range of eigenvalues.

(3) Apply block-Jacobi preconditioner by using SuperLU_dist on each poloidal plane independently



Now, range of eigenvalues in each block is greatly reduced.

Preconditioned system converges in 10's of iterations

GMRES

3 step physics-based preconditioner greatly improves iterative solve

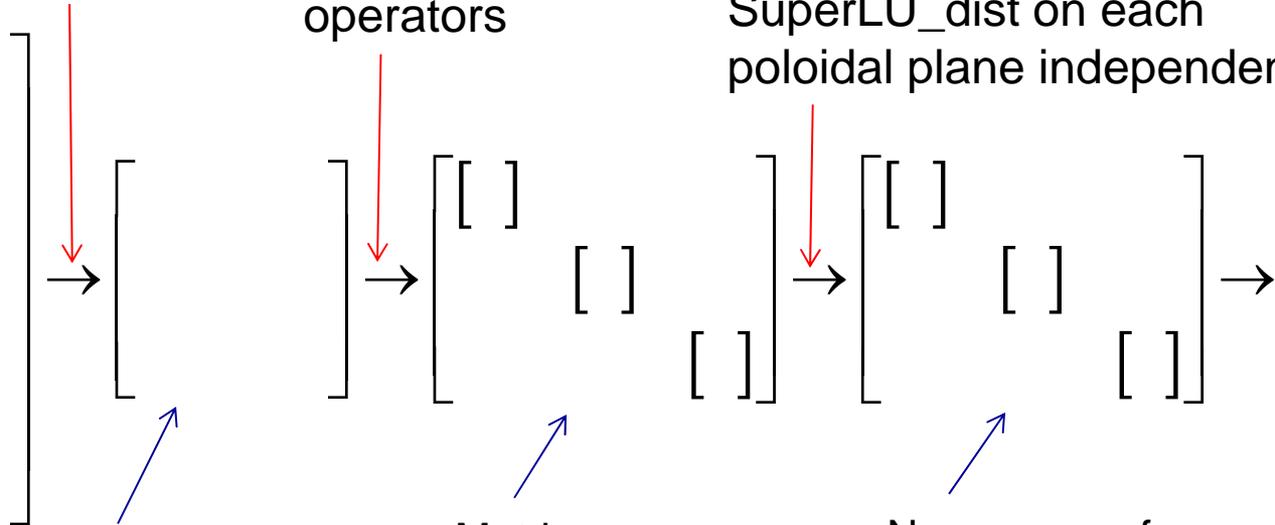
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(1) Split implicit formulation eliminates \mathbf{B}^{n+1} and p^{n+1} in favor of \mathbf{V}^{n+1}

As an example, consider the simple 1D wave equation for velocity V and pressure p

$$\left. \begin{aligned} \frac{\partial V}{\partial t} &= c \frac{\partial p}{\partial x} \\ \frac{\partial p}{\partial t} &= c \frac{\partial V}{\partial x} \end{aligned} \right\} \frac{\partial^2 V}{\partial t^2} - c^2 \frac{\partial^2 V}{\partial x^2} = 0$$

Implicit FD time advance evaluates spatial derivatives at the new time level

$$\frac{V_j^{n+1} - V_j^n}{\delta t} = c \left(\frac{p_{j+1/2}^{n+1} - p_{j-1/2}^{n+1}}{\delta x} \right)$$
$$\frac{p_{j+1/2}^{n+1} - p_{j+1/2}^n}{\delta t} = c \left(\frac{V_{j+1}^{n+1} - V_j^{n+1}}{\delta x} \right)$$

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$$p_{j+1/2}^{n+1} - p_{j+1/2}^n = c \left(\frac{V_{j+1}^{n+1} - V_j^{n+1}}{\delta x} \right)$$

Now, algebraically eliminate new time pressure in favor of velocity

$$V_j^{n+1} = V_j^n + (\delta t c)^2 \left(\frac{V_{j+1}^{n+1} - 2V_j^{n+1} + V_{j-1}^{n+1}}{\delta x^2} \right) + \delta t c \left(\frac{p_{j+1/2}^n - p_{j-1/2}^n}{\delta x} \right)$$

← Symmetric & diagonally dominant!

$$p_{j+1/2}^{n+1} = p_{j+1/2}^n + \frac{\delta t c}{\delta x} (V_{j+1}^{n+1} - V_j^{n+1})$$

These equations will give exactly the same answers, but can be solved sequentially!

Relation of split implicit method to Schur Complement

The original system can be written as:

$$\begin{bmatrix} \mathbf{A} & \mathbf{B} \\ \mathbf{C} & \mathbf{D} \end{bmatrix} \cdot \begin{bmatrix} \mathbf{V} \\ \mathbf{P} \end{bmatrix}^{n+1} = \begin{bmatrix} \mathbf{R} \\ \mathbf{Q} \end{bmatrix}^n \quad \leftarrow \text{known RHS}$$

Here, \mathbf{A} and \mathbf{D} are diagonal matrices, and

$$\mathbf{B} = \begin{bmatrix} \dots & & & & & \\ s & -s & & & & \\ & s & -s & & & \\ & & s & -s & & \\ & & & s & -s & \\ & & & & s & \dots \end{bmatrix}, \quad \mathbf{C} = -\mathbf{B}^T = \begin{bmatrix} \dots & & & & & \\ & s & -s & & & \\ & & s & -s & & \\ & & & s & -s & \\ & & & & s & -s \\ & & & & & \dots \end{bmatrix}, \quad \mathbf{V} = \begin{bmatrix} \dots \\ V_{j-1} \\ V_j \\ V_{j+1} \\ \dots \end{bmatrix}, \quad \mathbf{P} = \begin{bmatrix} \dots \\ p_{j-1/2} \\ p_{j+1/2} \\ p_{j+3/2} \\ \dots \end{bmatrix}.$$

Solve first for \mathbf{P}^{n+1} in terms of \mathbf{V}^{n+1} : $\mathbf{P}^{n+1} = -\mathbf{D}^{-1} \mathbf{C} \mathbf{V}^{n+1} + \mathbf{D}^{-1} \mathbf{Q}^n$

Next, eliminate \mathbf{P}^{n+1} : $\mathbf{A}' \mathbf{V}^{n+1} = \mathbf{R}'^n \quad \mathbf{A}' = \mathbf{A} - \mathbf{B} \mathbf{D}^{-1} \mathbf{C}$

$$\mathbf{R}'^n = \mathbf{R}^n - \mathbf{B} \mathbf{D}^{-1} \mathbf{Q}^n$$

In this example, \mathbf{A}' is symmetric

Now apply this technique to the basic 3D MHD equations:

$$\rho_0 \dot{\mathbf{V}} = \frac{1}{\mu_0} [\nabla \times \mathbf{B}] \times \mathbf{B} - \nabla p$$

$$\dot{\mathbf{B}} = \nabla \times [\mathbf{V} \times \mathbf{B}]$$

$$\dot{p} = -\mathbf{V} \cdot \nabla p - \gamma p \nabla \cdot \mathbf{V}$$

Ideal MHD Equations for velocity, magnetic field, and pressure:

Symmetric Hyperbolic System

7-waves

$$\rho_0 \dot{\mathbf{V}} = \frac{1}{\mu_0} [\nabla \times (\mathbf{B} + \theta \delta t \dot{\mathbf{B}})] \times (\mathbf{B} + \theta \delta t \dot{\mathbf{B}}) - \nabla (p + \theta \delta t \dot{p})$$

$$\dot{\mathbf{B}} = \nabla \times [(\mathbf{V} + \theta \delta t \dot{\mathbf{V}}) \times \mathbf{B}]$$

Taylor Expand in Time

$$\dot{p} = -(\mathbf{V} + \theta \delta t \dot{\mathbf{V}}) \cdot \nabla p - \gamma p \nabla \cdot (\mathbf{V} + \theta \delta t \dot{\mathbf{V}})$$

Substitute from 2nd and 3rd equation into first, finite difference in time:

$$\{\rho - \theta^2 (\delta t)^2 L\} \mathbf{V}^{n+1} = \{\rho - \theta^2 (\delta t)^2 L\} \mathbf{V}^n + \delta t \left\{ -\nabla p + \frac{1}{\mu_0} (\nabla \times \mathbf{B}) \times \mathbf{B} \right\}^n$$

MHD Operator: \longrightarrow

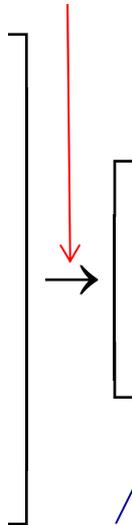
$$L\{\mathbf{V}\} = \frac{1}{\mu_0} \left\{ \nabla \times [\nabla \times (\mathbf{V} \times \mathbf{B})] \right\} \times \mathbf{B} + \frac{1}{\mu_0} (\nabla \times \mathbf{B}) \times [\nabla \times (\mathbf{V} \times \mathbf{B})] + \nabla (\mathbf{V} \cdot \nabla p + \gamma p \nabla \cdot \mathbf{V})$$

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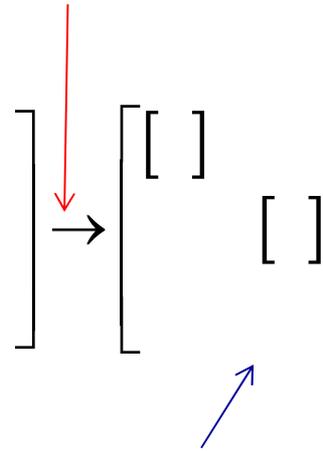
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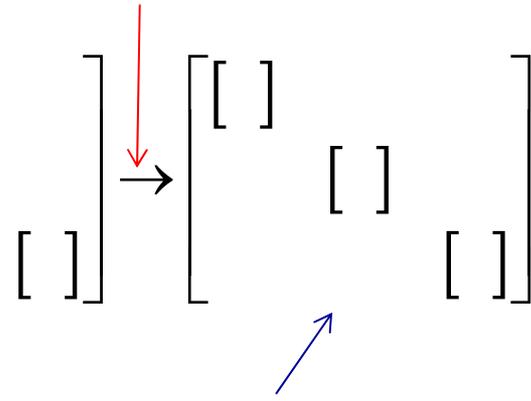
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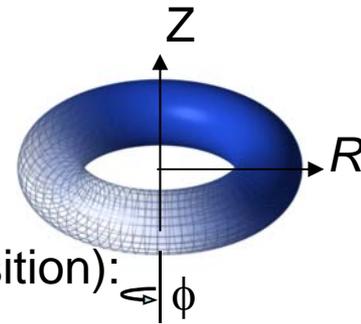


Now, range of eigenvalues in each block is greatly reduced.

Preconditioned system converges in 10's of iterations

GMRES

(2) Apply annihilation operators to separate eigenvalues into diagonal blocks



Velocity vector written in terms of 3 scalar fields (Helmholtz decomposition):

$$\mathbf{V} = R^2 \nabla U \times \nabla \phi + R^2 \omega \nabla \phi + \frac{1}{R^2} \nabla_{\perp} \chi$$

$$\nabla_{\perp} \equiv \hat{R} \frac{\partial}{\partial R} + \hat{Z} \frac{\partial}{\partial Z}$$

Associated mainly with the **shear Alfvén** wave: **does not** compress the toroidal field

Associated mainly with the **slow** wave: also **does not** compress the toroidal field

Associated mainly with the **fast** wave: **does** compress the toroidal field

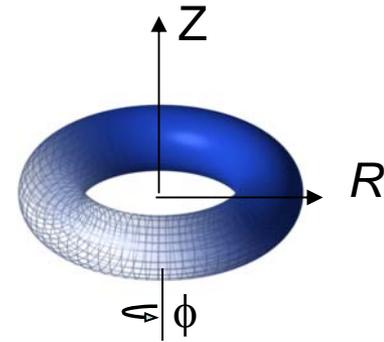
To obtain scalar equations, we apply annihilation projections to isolate the physics associated with the different wave types in different blocks in the matrix

Alfvén wave:	$\nabla \phi \cdot \nabla_{\perp} \times R^2$	$\left. \begin{aligned} & (1 - \theta^2(\delta t)L) \frac{\partial \mathbf{V}}{\partial t} + \mathbf{V} \cdot \nabla \mathbf{V} \\ & = \frac{1}{nM_i} \left[-\nabla p + \mathbf{J} \times \mathbf{B} - \nabla \cdot \mathbf{\Pi}_{GV} - \nabla \cdot \mathbf{\Pi}_{\mu} \right] \end{aligned} \right\}$
slow wave:	$R^2 \nabla \phi \cdot$	
fast wave:	$-\nabla_{\perp} \cdot R^{-2}$	

Code can be run with 1,2 (reduced MHD) or 3 (full MHD) velocity variables

Aside on the form of the Vector Fields

Because the externally imposed toroidal field in a tokamak is very strong, any plasma instability will slip through this field and not compress it. We need to be able to model this motion very accurately because of the weak forces causing the instability.



In **M3D-C1**, we express the velocity field as shown

$$\mathbf{V} = R^2 \nabla U \times \nabla \phi + R^2 \omega \nabla \phi + \frac{1}{R^2} \nabla_{\perp} \chi$$

All components orthogonal!

$$\int |\mathbf{V}|^2 d\tau = \int \left[R^2 |\nabla_{\perp} U|^2 + R^2 \omega^2 + \frac{1}{R^4} |\nabla_{\perp} \chi|^2 \right] d\tau$$

Consider now the action of the first term in \mathbf{V} on the external toroidal field:

$$\mathbf{B} = F_0 \nabla \phi$$

$$\mathbf{V} = R^2 \nabla U \times \nabla \phi$$

$$\nabla \phi \cdot \left[\frac{\partial \mathbf{B}}{\partial t} = \nabla \times (\mathbf{V} \times \mathbf{B}) \right]$$

$$= \nabla \phi \cdot \nabla \times \left[\left(R^2 \nabla U \times \nabla \phi \right) \times F_0 \nabla \phi \right] \quad \rightarrow$$

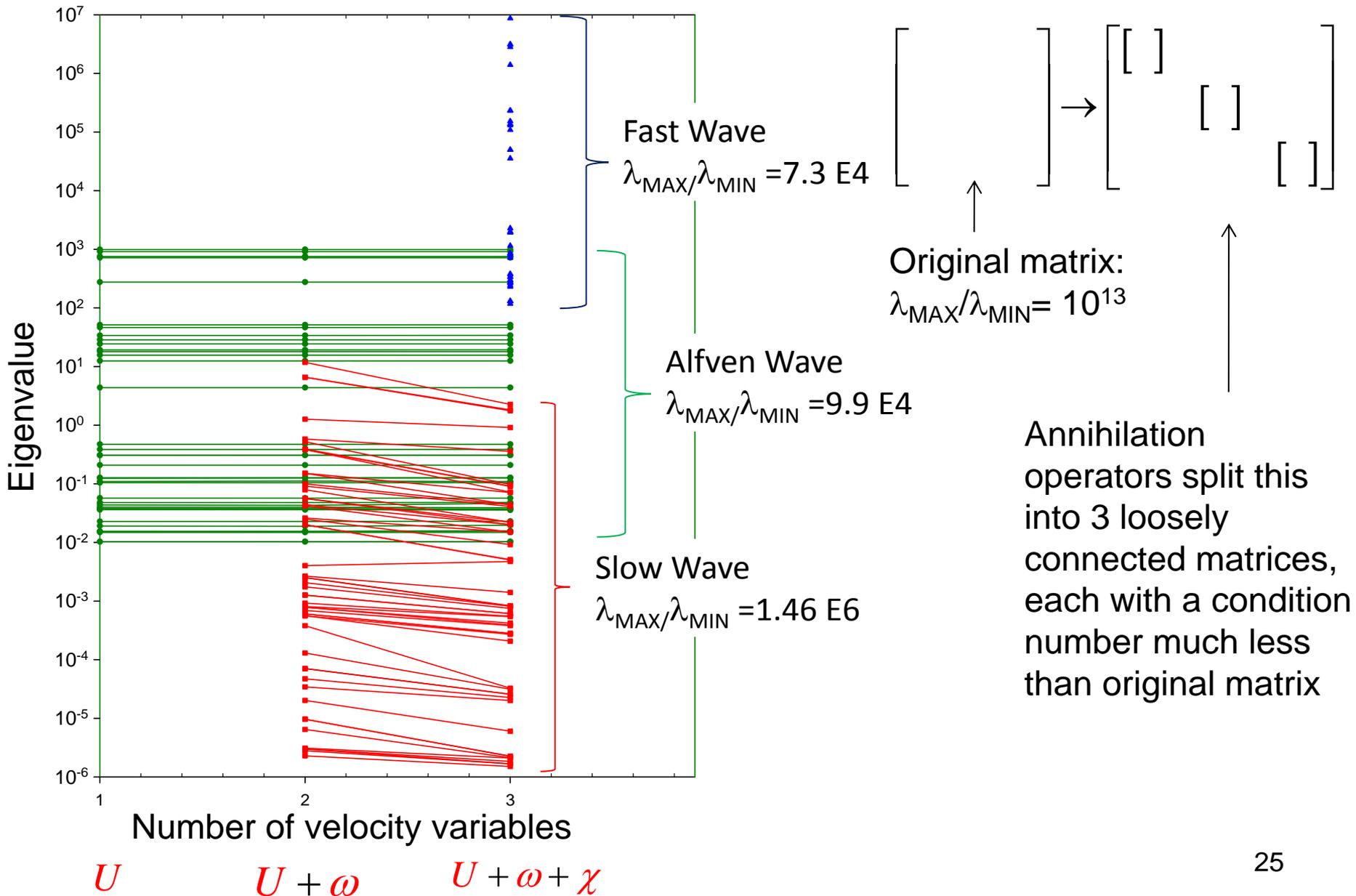
$$= -F_0 \nabla \cdot \left[\nabla U \times \nabla \phi \right]$$

$= 0$ The velocity field U does not compress the external toroidal field!

- Any unstable motion will mostly consist of the velocity component U

- Analytic elimination of this potentially stabilizing term greatly increases accuracy. 24

M3D-C¹ can be run with 1, 2, or 3 velocity variables. Tracking the eigenvalues shows how they separate into 3 groups in a cylinder

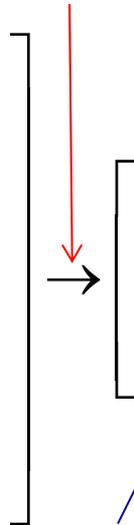


3 step physics-based preconditioner greatly improves iterative solve

(1) Split implicit formulation

Original matrix multiplying \mathbf{V}^{n+1} , \mathbf{B}^{n+1} , \mathbf{p}^{n+1}

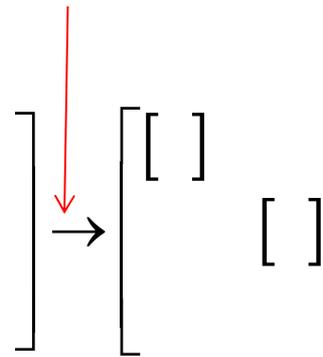
- non-symmetric,
- non-diagonally dominant &
- large range of eigenvalues



Smaller matrix multiplying \mathbf{V}^{n+1} only,

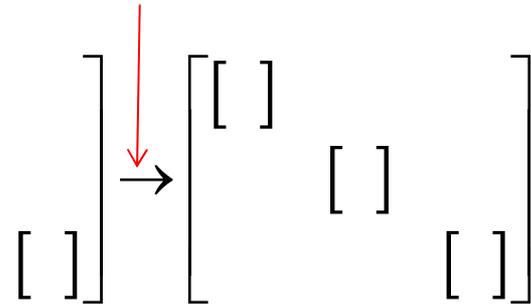
- nearly symmetric
- closer to diagonal
- still with large range of eigenvalues

(2) Apply annihilation operators



Matrix now consists of 3 dominant diagonal blocks, each with narrower range of eigenvalues.

(3) Apply block-Jacobi preconditioner by using SuperLU_dist on each poloidal plane independently



Now, range of eigenvalues in each block is greatly reduced.

Preconditioned system converges in 10's of iterations

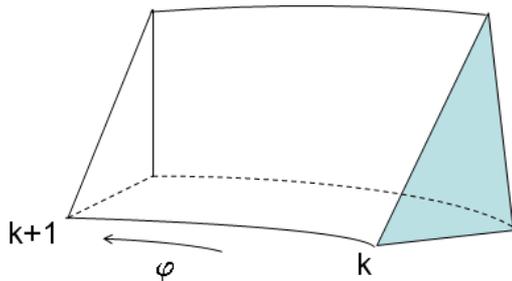
GMRES

(3) Apply block-Jacobi preconditioner by using SuperLU_dist on each poloidal plane independently

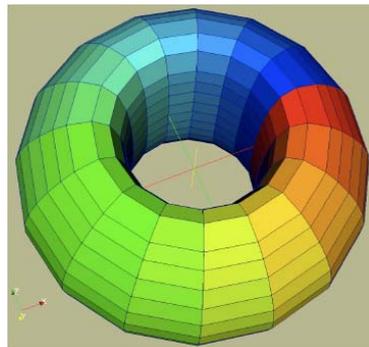
M3D-C¹ uses a triangular wedge high order finite element

- Continuous 1st derivatives in all directions ... *C¹ continuity*
- Unstructured triangles in (R,Z) plane
- Structured in toroidal direction (φ)

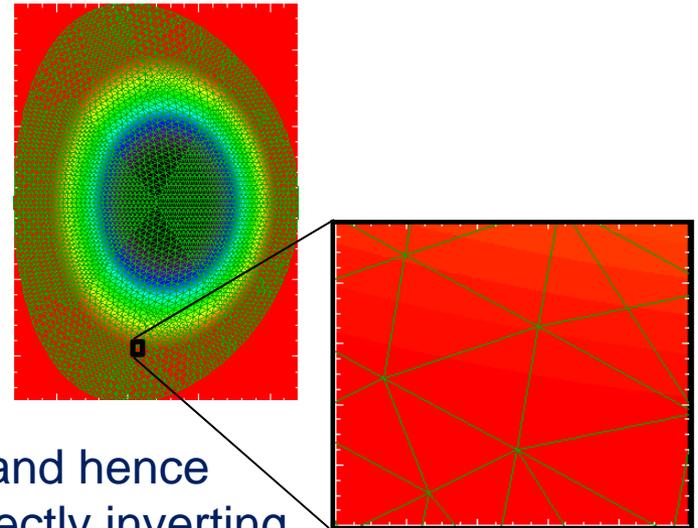
Triangular wedge integration volume



Top view: 16-32 toroidal prisms



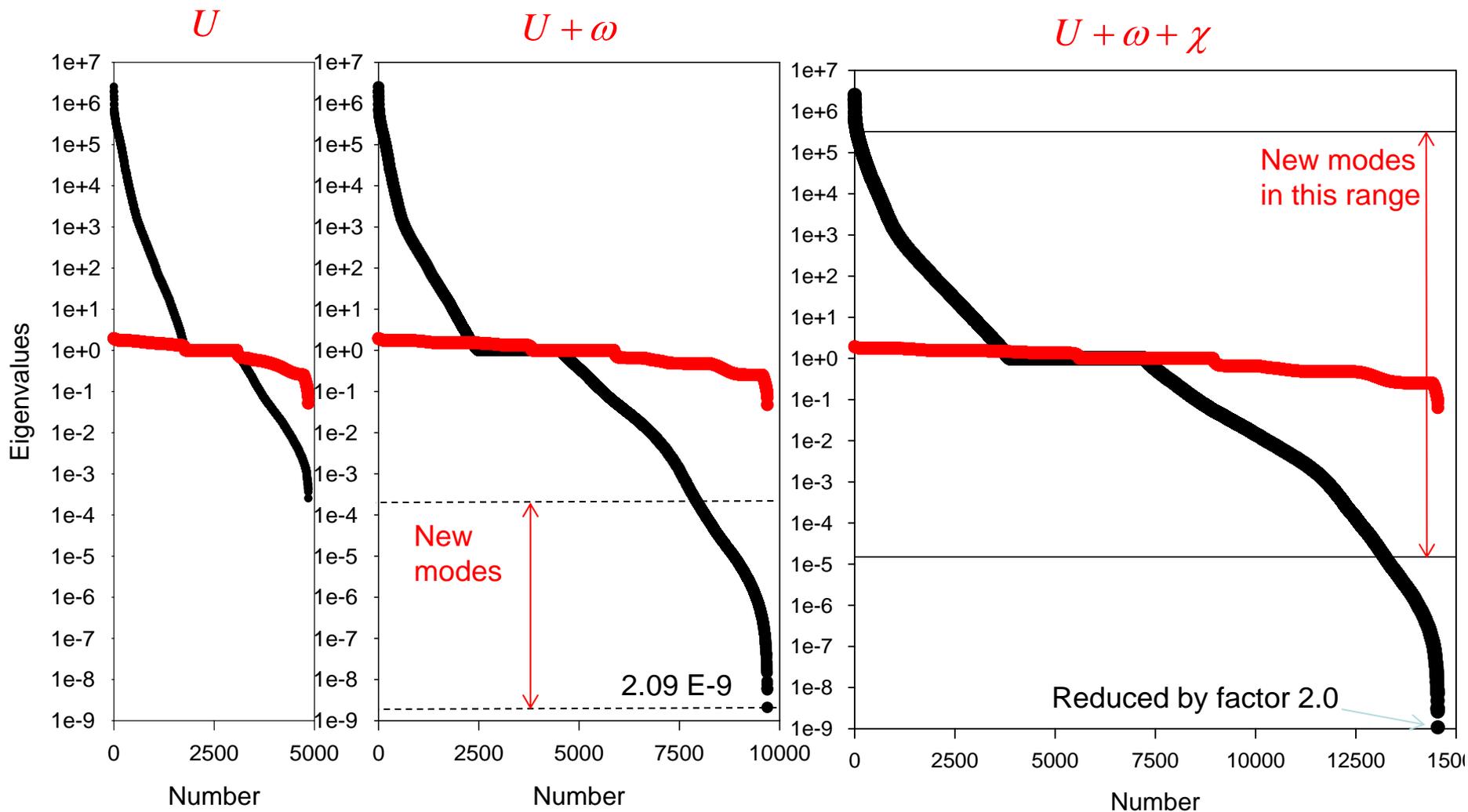
Slice view: ~ 10⁴ nodes/plane



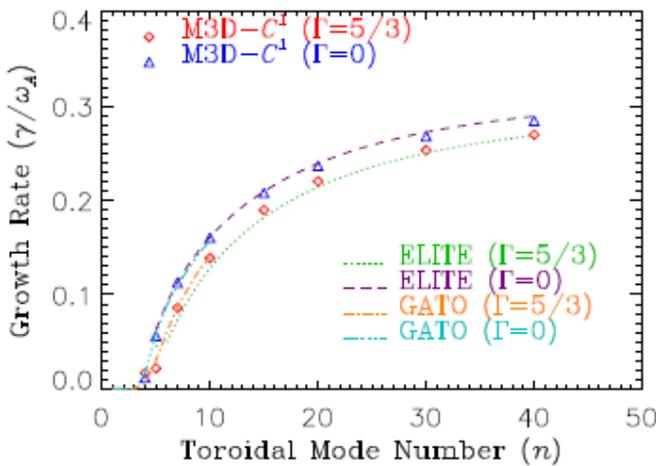
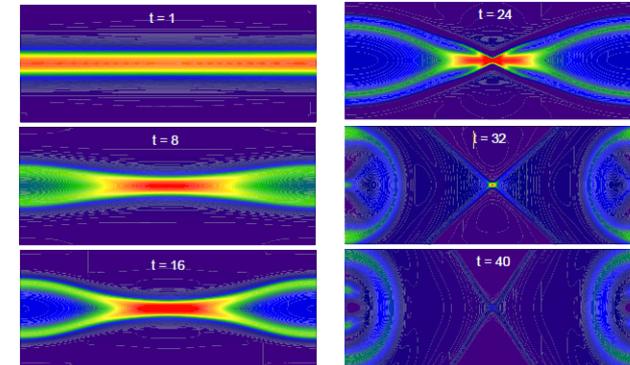
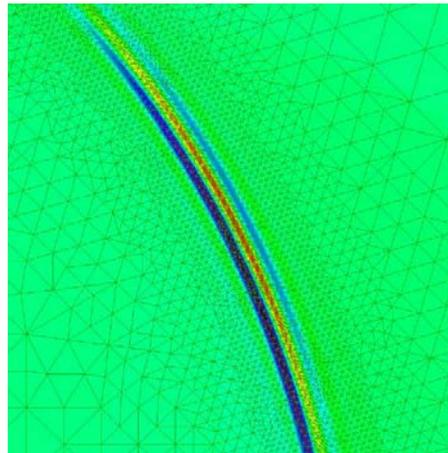
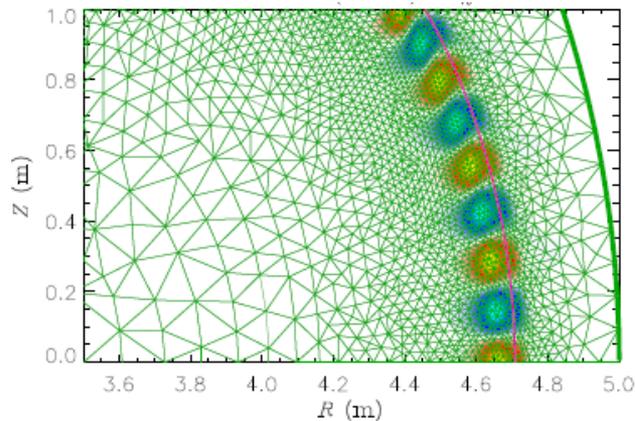
Because of the small zone size within the plane, and hence strong coupling, we precondition the matrix by directly inverting the components within each poloidal plane simultaneously.

Block Jacobi Preconditioner: greatly reduces condition number

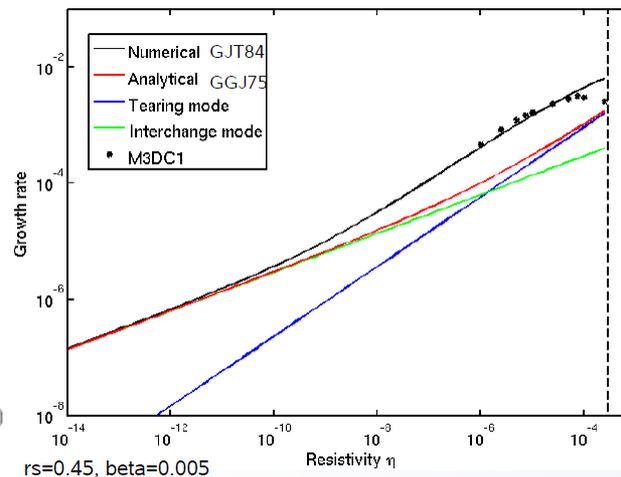
Eigenvalues of A=3 3D Matrix Before and **After** Preconditioning



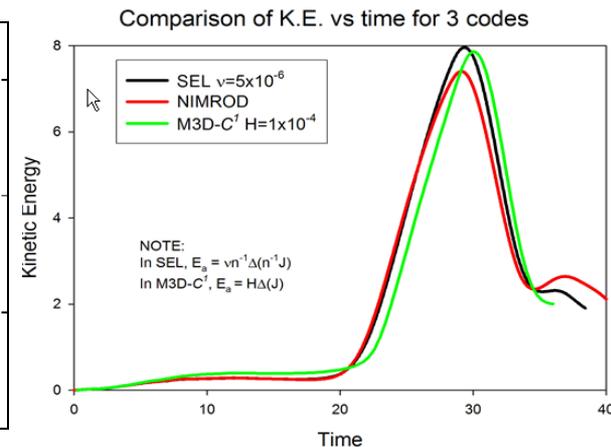
Extensive benchmarking for ideal, resistive, and two fluid modes



Ideal MHD

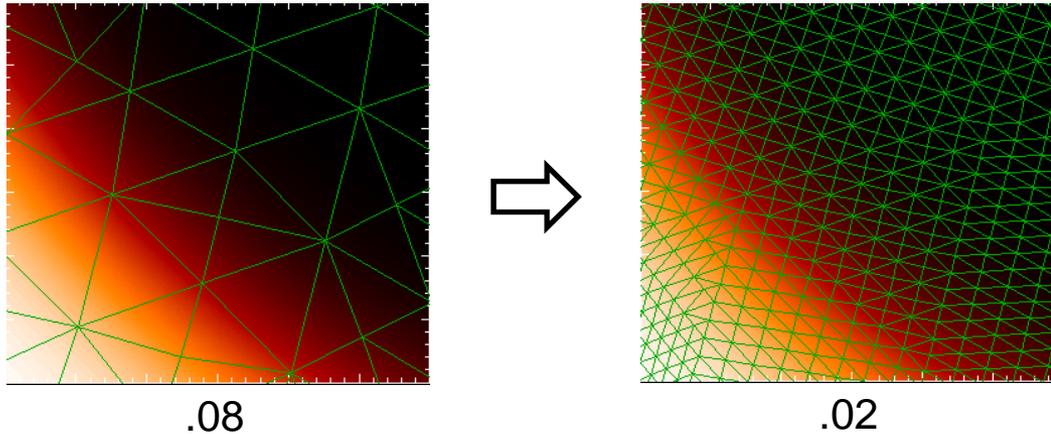


Resistive MHD

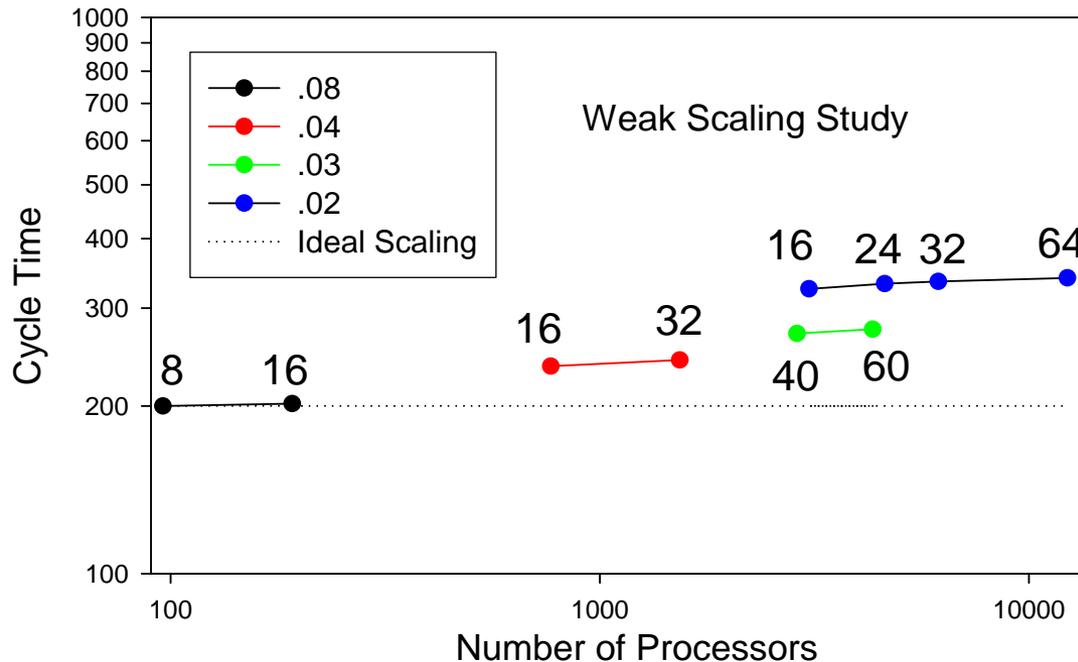


2-Fluid
Reconnection

Parallel Scaling Studies have been performed from 96 to 12288 p



Triangle linear dimension varied by factor of 4



Number of toroidal planes varied from 8 to 64

Time increased by 1.7 as # of zones increased by 130

Transport Timescale simulations in which stability is important:
with $\Delta t = 40 \tau_A$

Specify a transport model:

Resistivity: $\eta = n^{3/2} p^{-3/2}$

Thermal Conductivity: $\kappa_{\perp} = n^{3/2} p^{-1/2}$

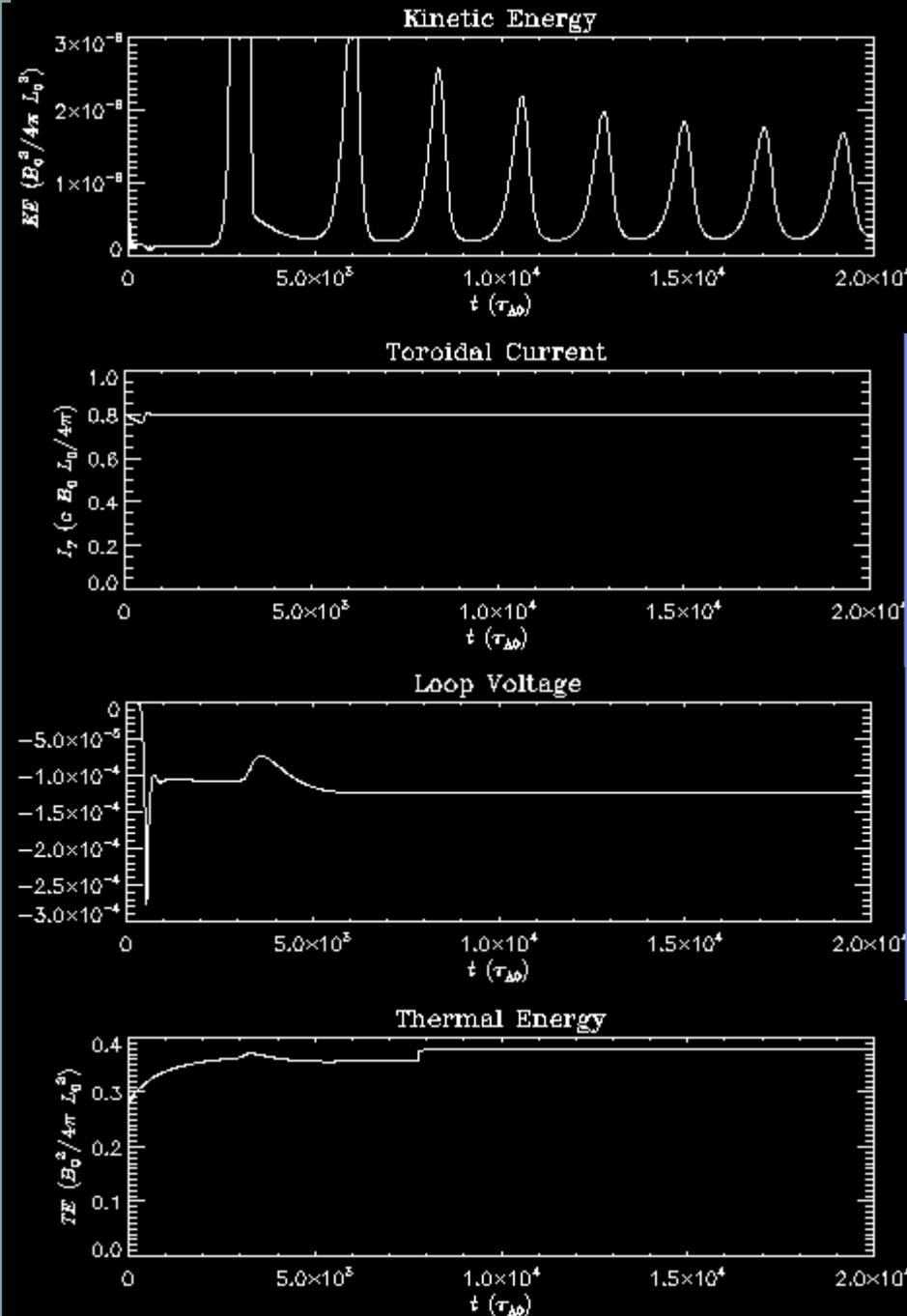
$$\kappa_{\parallel} = 10^6 \kappa_{\perp}$$

Viscosity: uniform ($\sim \eta$)

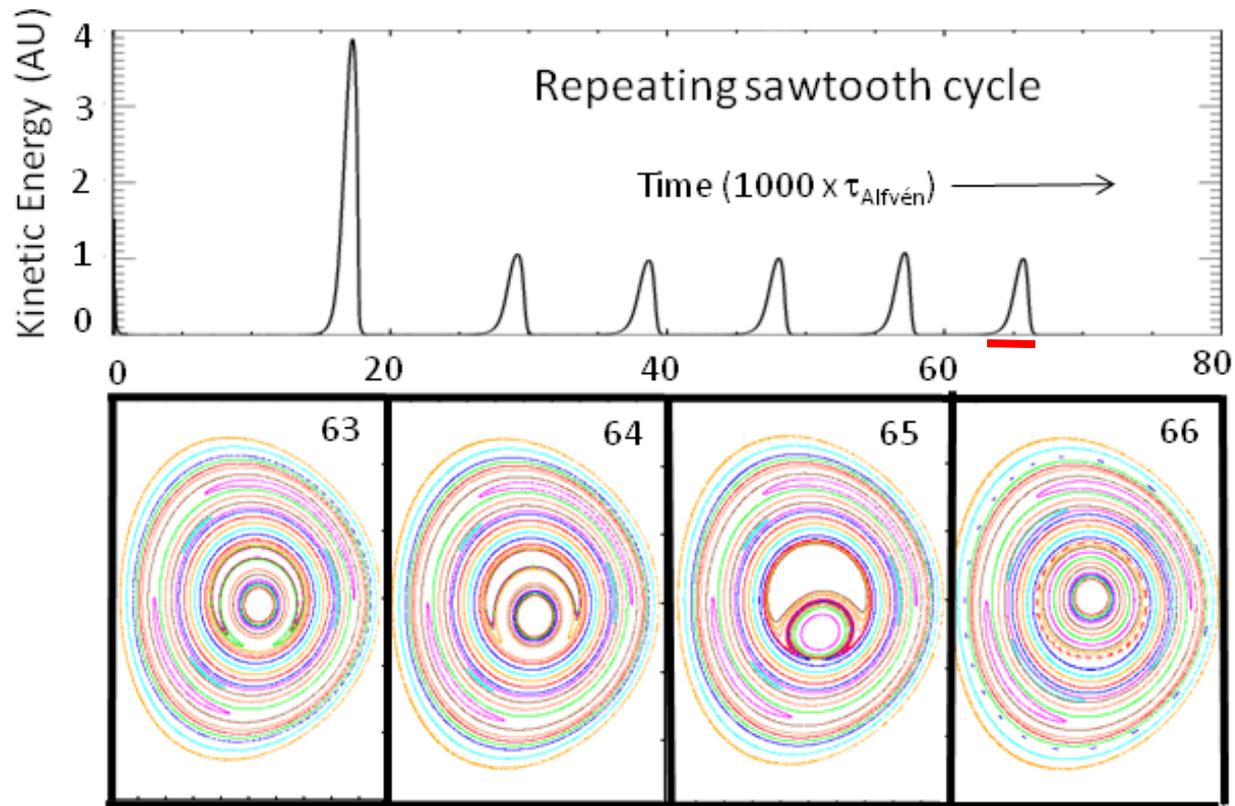
Current controller provides loop voltage to maintain plasma current at initial value.

Loop voltage provides thermal energy through Ohmic heating

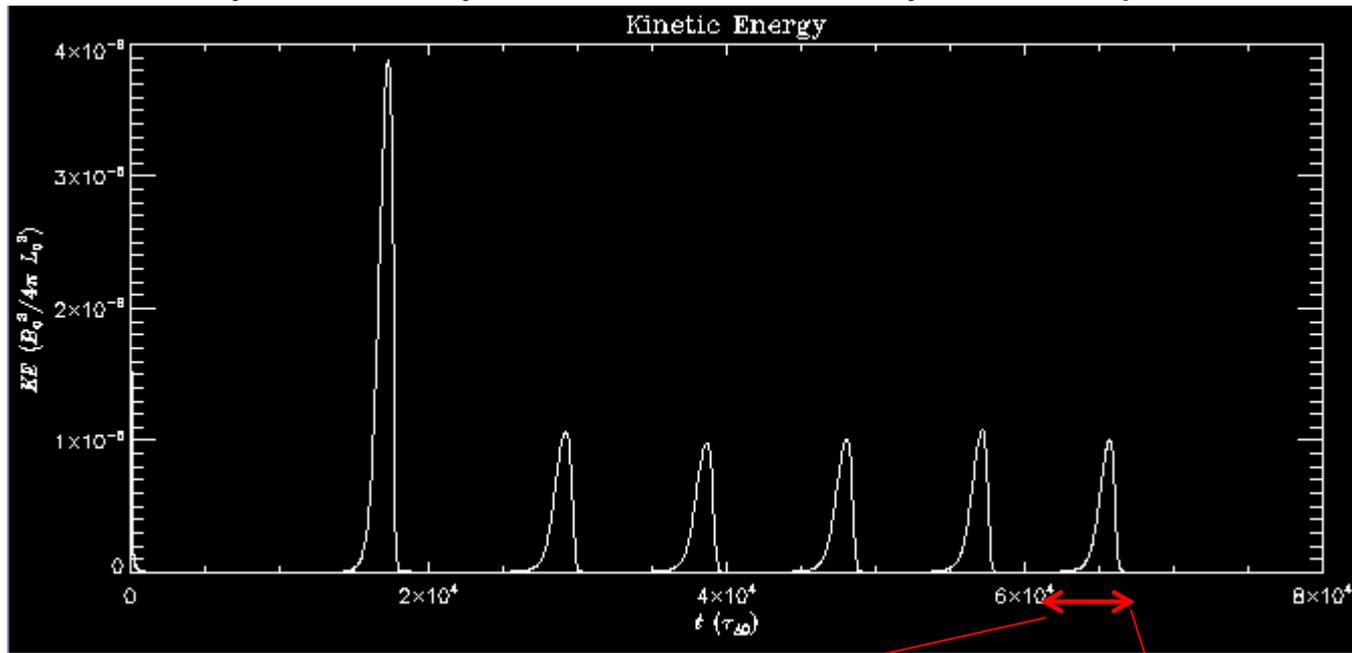
Current density periodically peaks, becomes unstable, reconnects, and broadens...periodic cycle (sawtooth)



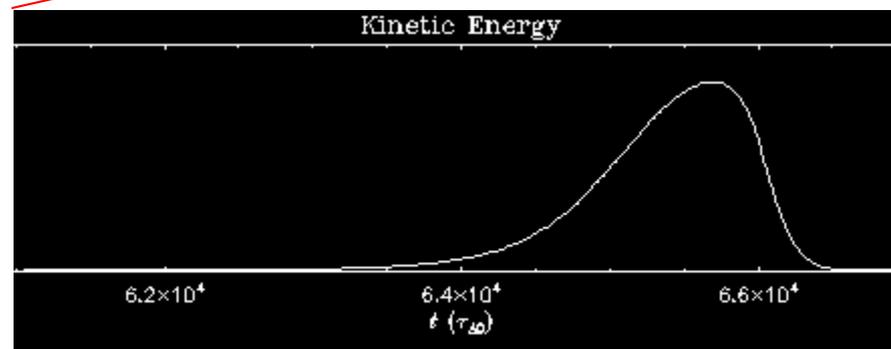
Generic sawtooth studies are now underway



Typical result: 1st sawtooth event depends on initial conditions.
After many events, system reaches steady-state or periodic behavior

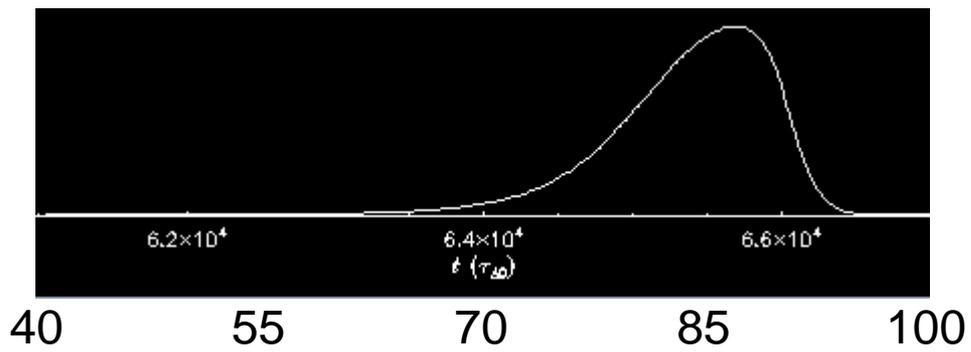
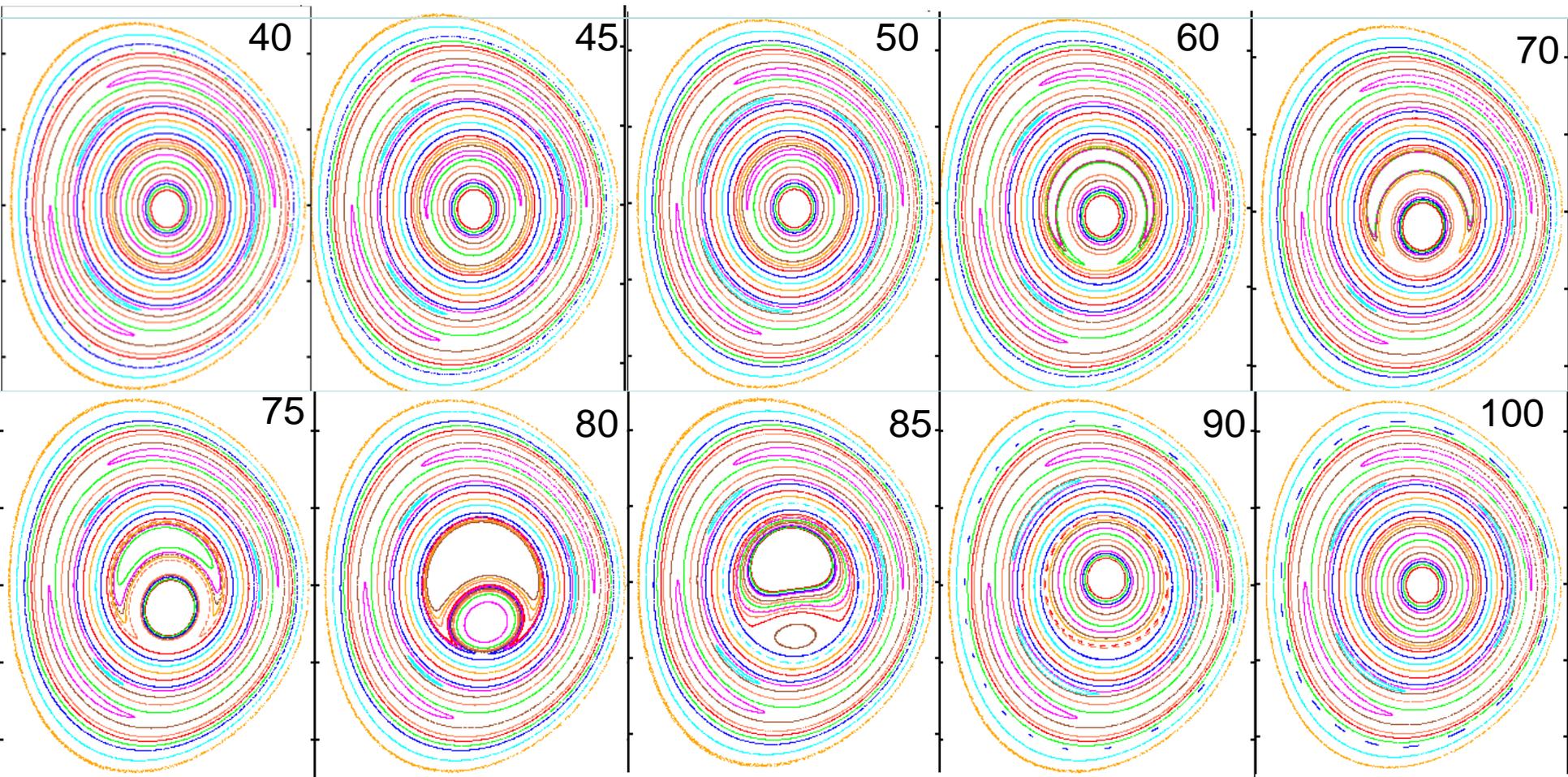


Repeating
sawtooth cycle



← Precursor phase → ← Crash phase →

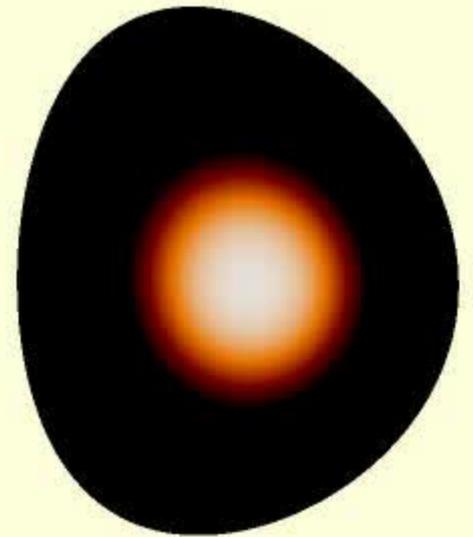
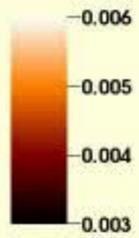
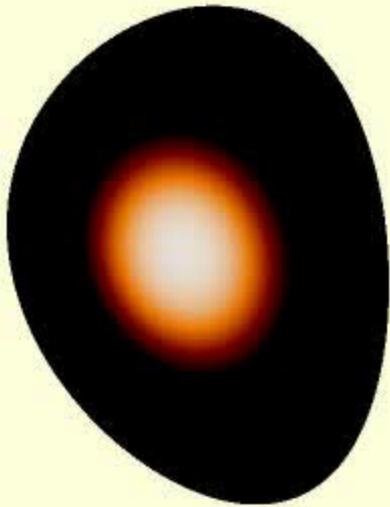
Poincare plots during a single sawtooth cycle



DB: C1.h5
Cycle: 30

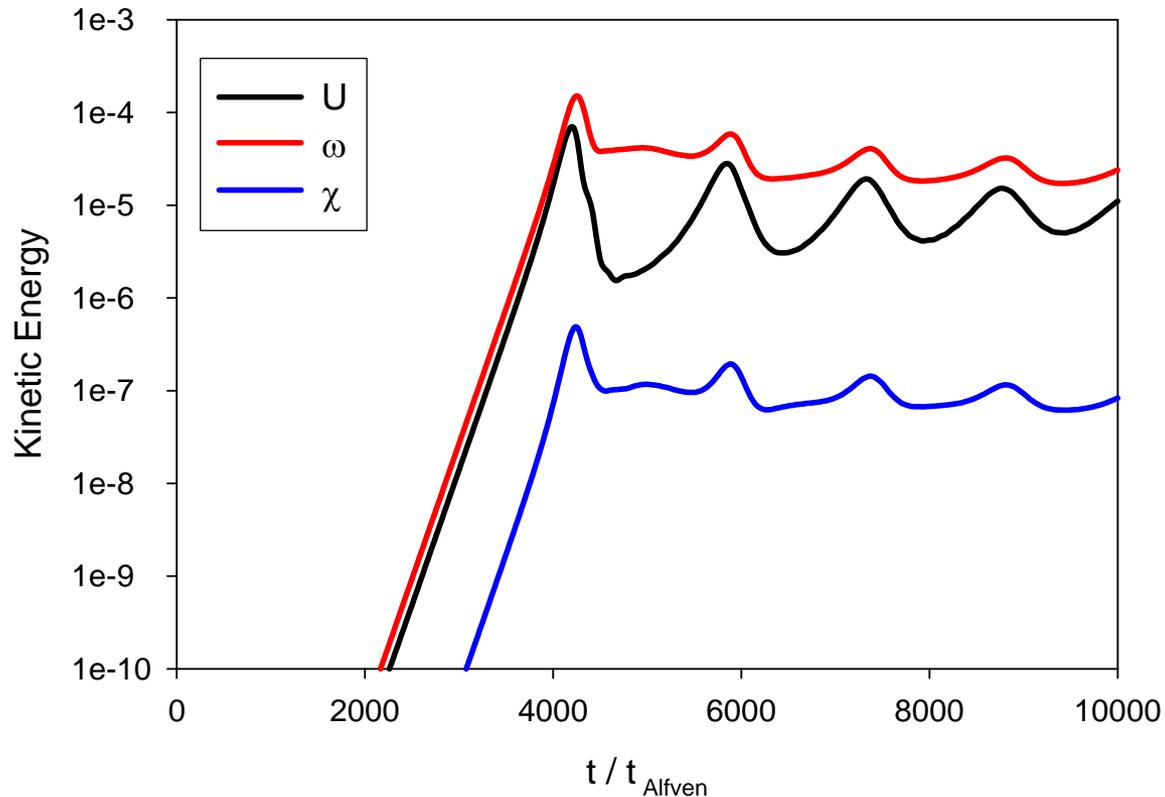
Time:60000

Isosurface of $p = .006$



$$\mathbf{V} = R^2 \nabla U \times \nabla \phi + R^2 \boldsymbol{\omega} \nabla \phi + \frac{1}{R^2} \nabla_{\perp} \chi$$

Kinetic Energy in the three velocity components

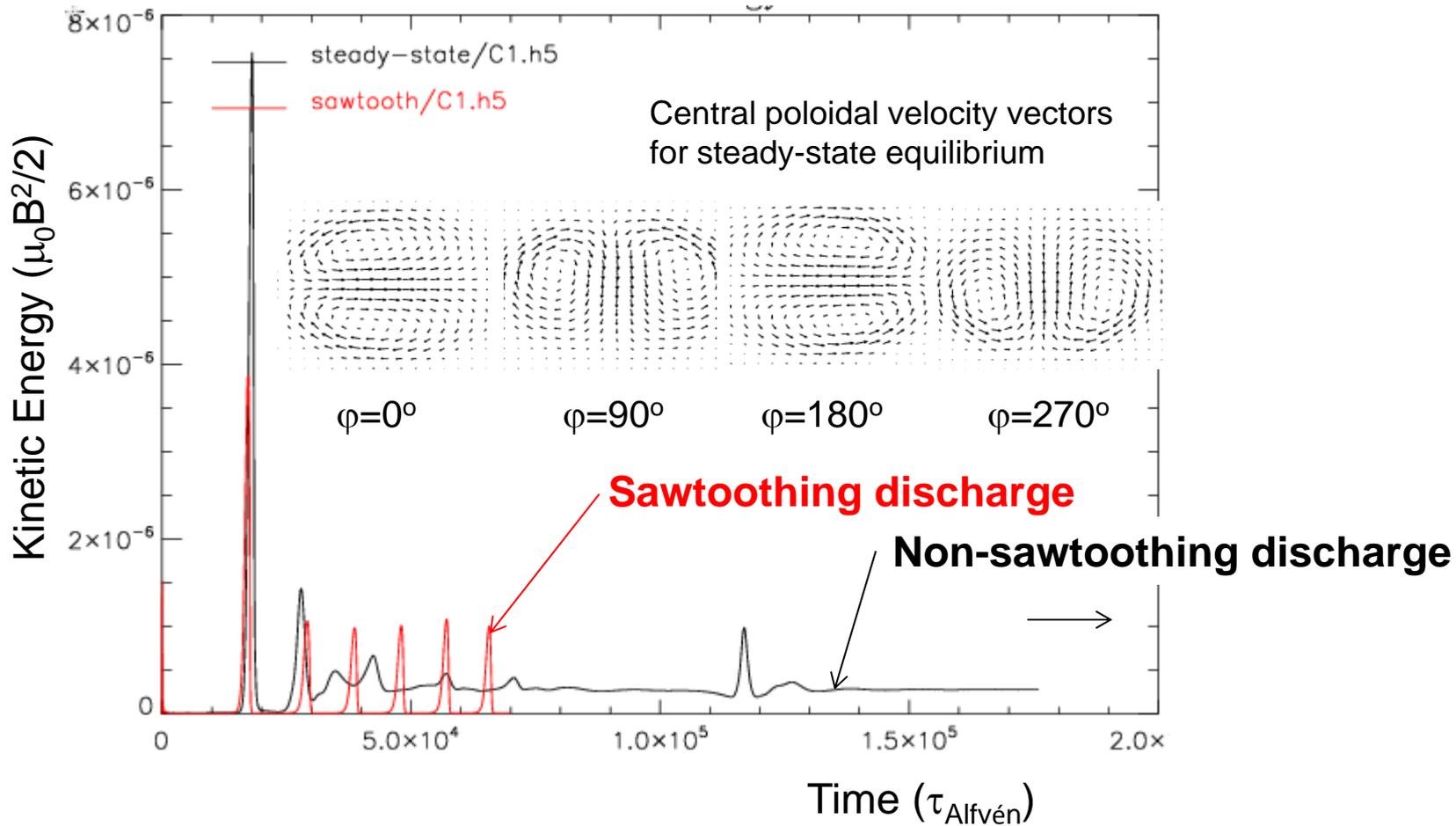


The poloidal velocity decomposition used in M3D-C¹ is very effective in capturing most of the poloidal flow in U .

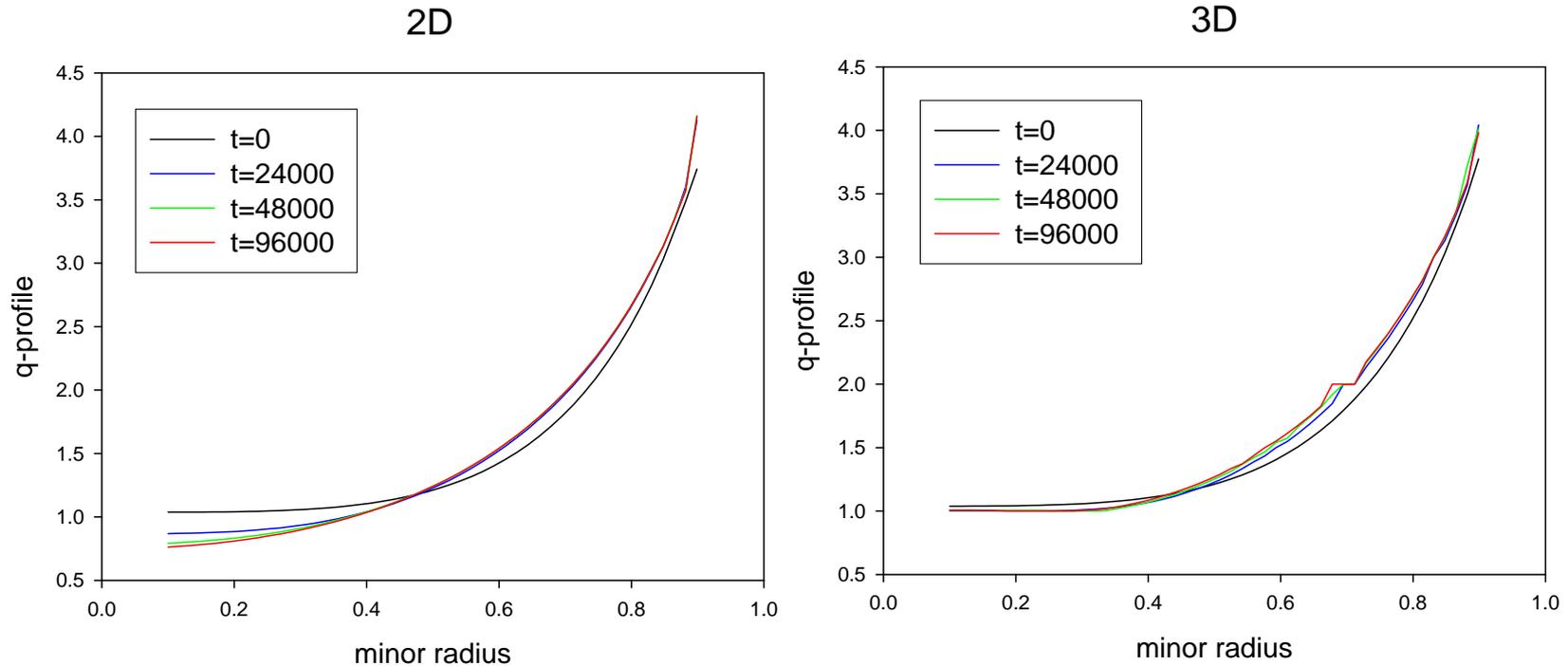
Stationary States with Flow

In lower viscosity cases, the sawtooth behavior stops after a few cycles, and a central helical (1,1) structure forms with flow.

This flow is such as to flatten the central pressure and temperature. This flattening causes the current density to also flatten near the center, keeping $q_0 \sim 1$ in the central region.



Comparison of 3D Sawtooth Free Helical Stationary State (SFHSS) with 2D configuration with same transport parameters

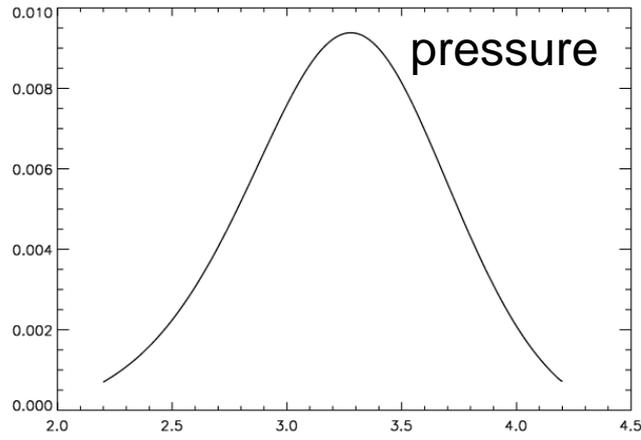


Exact same case was run with M3D-C¹ in 2D and in 3D.

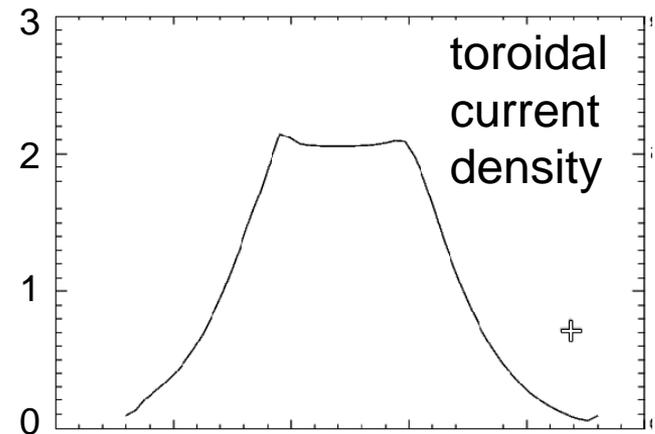
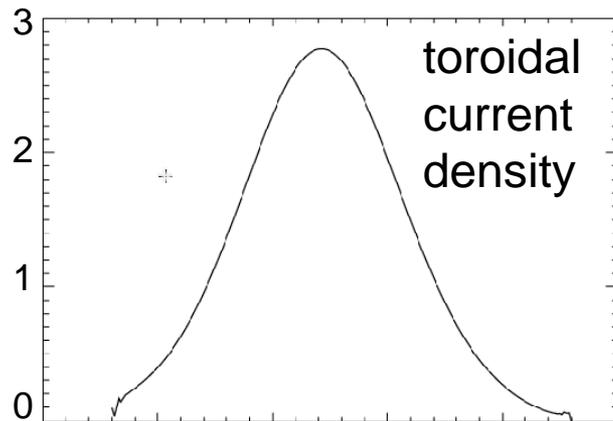
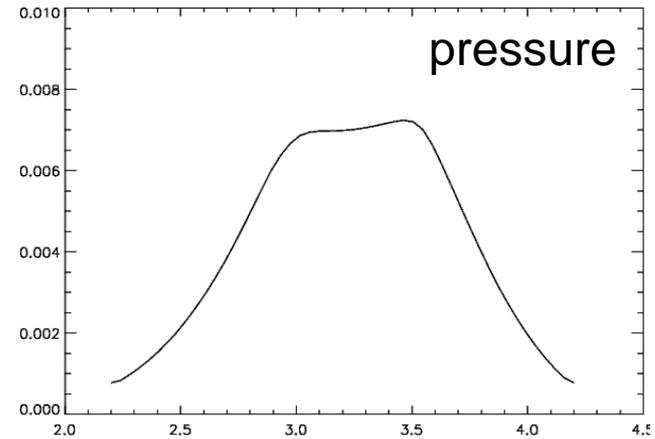
- In 2D, q_0 drops to about 0.7.
- In 3D, it is clamped at 1.0

Comparison of 3D Sawtooth Free Helical Stationary State (SFHSS) with 2D configuration with same transport parameters

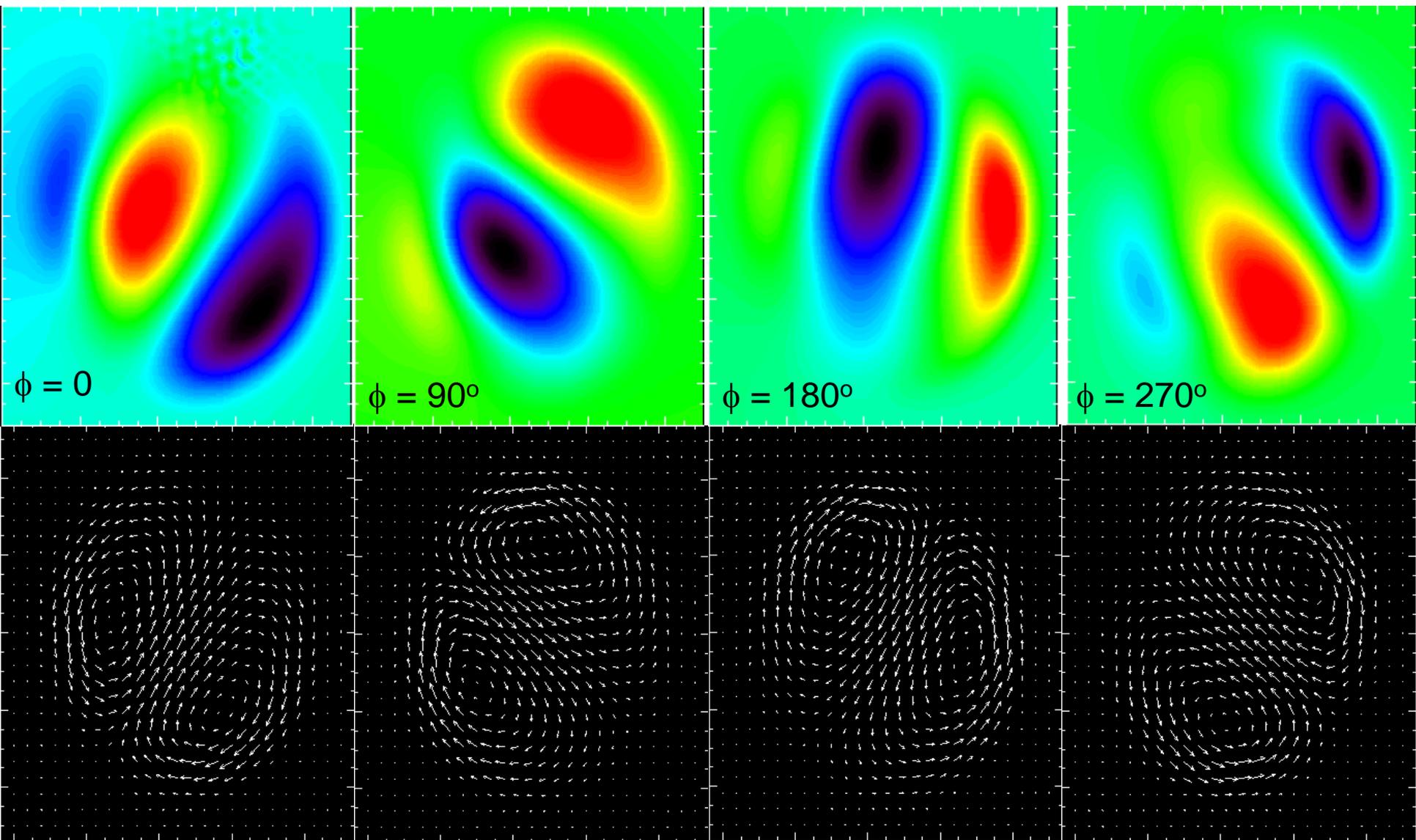
2D



3D

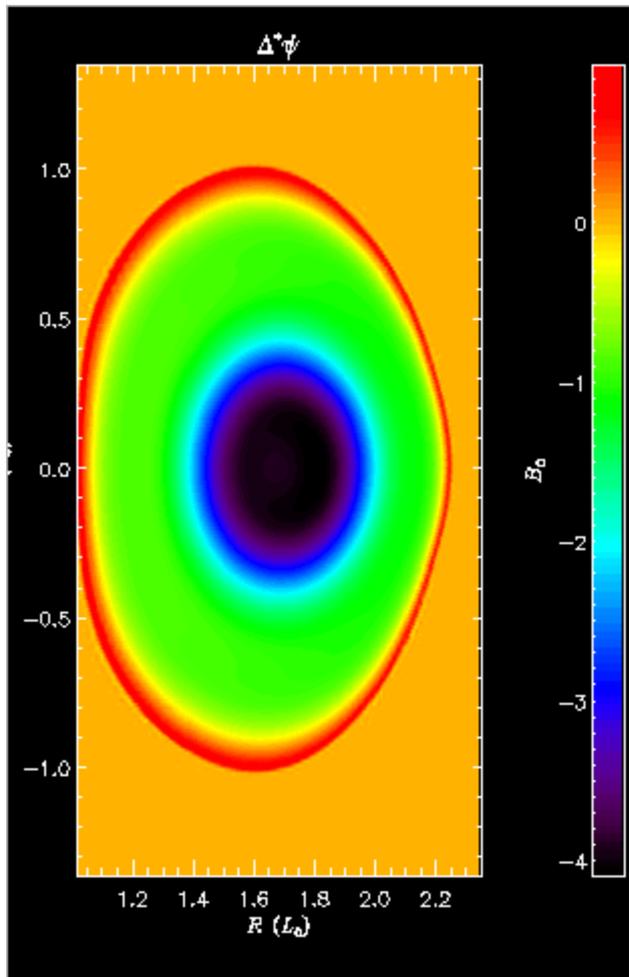


In 3D, pressure and toroidal current density are much less peaked.

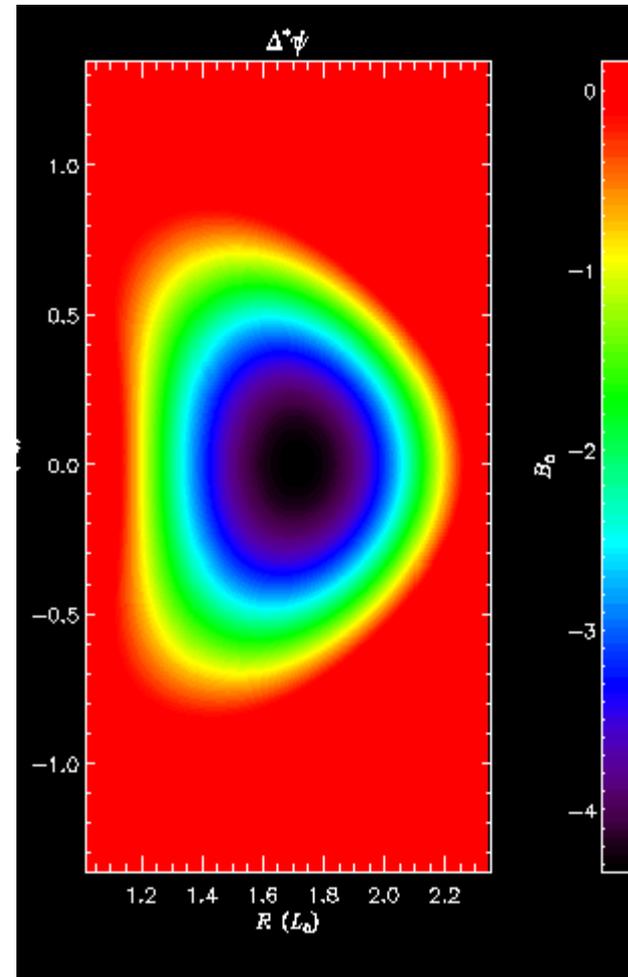


Close up of central toroidal velocity contours (top) and poloidal velocity vectors for stationary state at different toroidal angles: $V_T(\text{max}) \sim 0.0004$, $V_P(\text{max}) \sim 0.0002$

Differences in sawtooth behavior for bean-shaped and elliptical-shaped plasmas has been well documented experimentally (Lazarus, Tobias, ...)

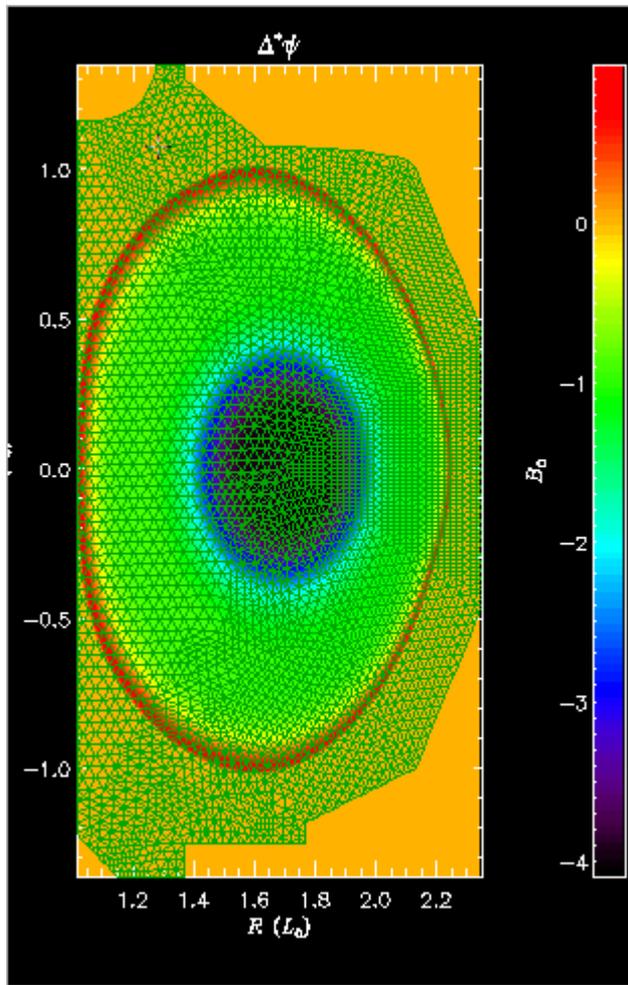


DIII-D shot 118164

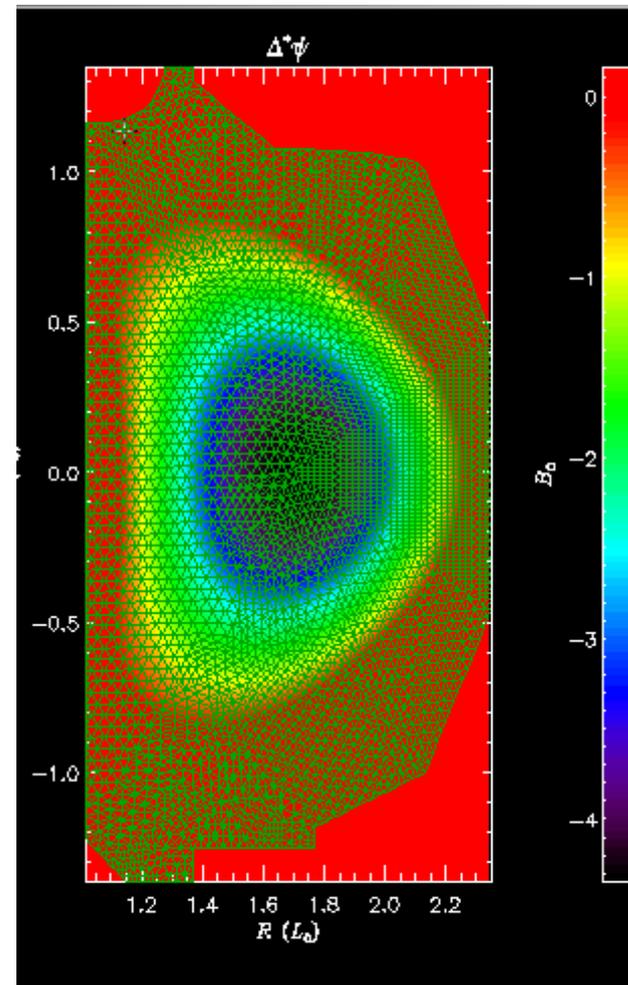


DIII-D shot 118162

We have imported these equilibria from geqdsk files, and inferred the transport properties from the plasma properties...simulations in progress

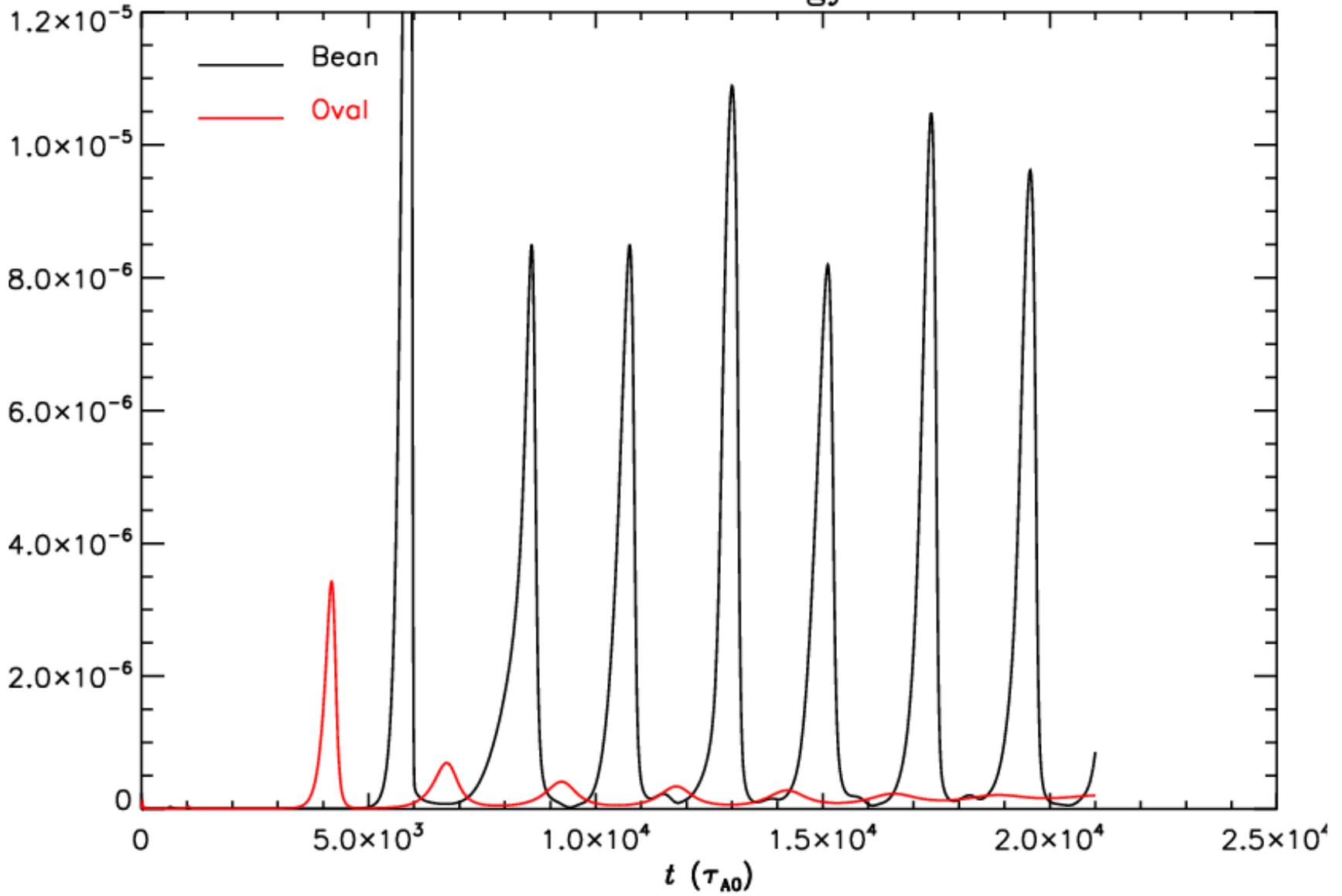


DIII-D shot 118164

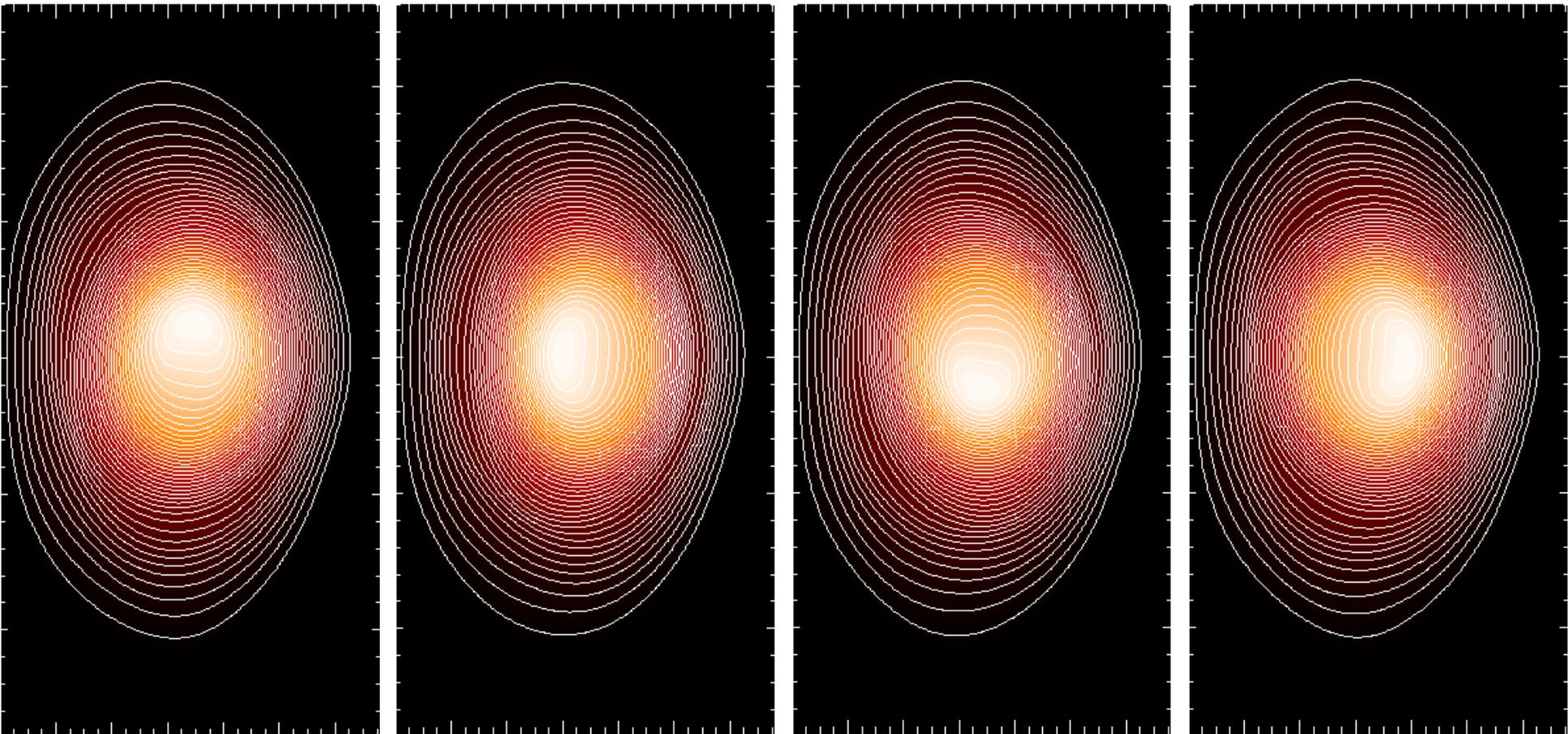


DIII-D shot 118162

Kinetic Energy



Stationary Pressure for DIII-D Oval (shot 118164)



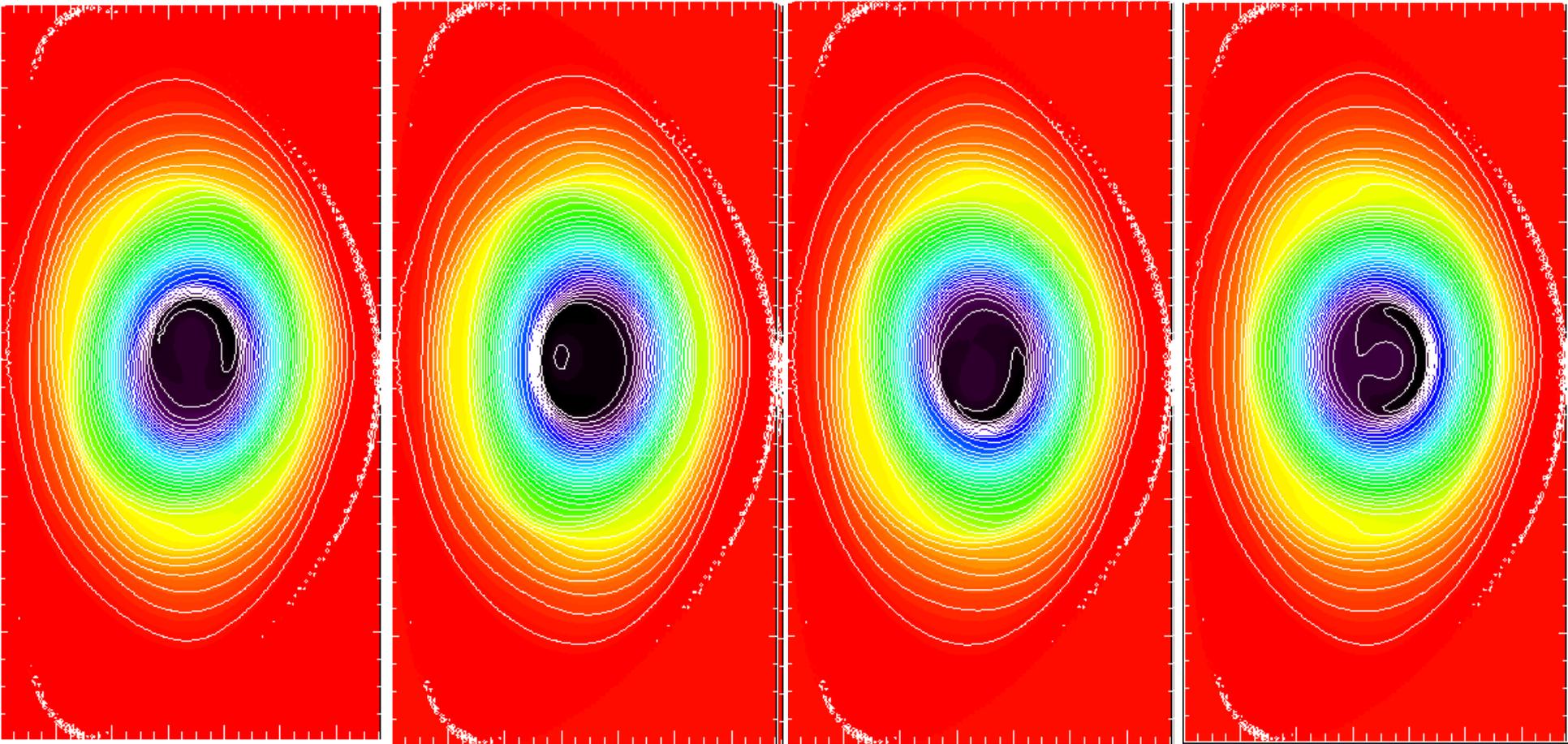
$\varphi = 0^\circ$

$\varphi = 90^\circ$

$\varphi = 180^\circ$

$\varphi = 270^\circ$

Stationary Current Density for DIII-D Oval (shot 118164)



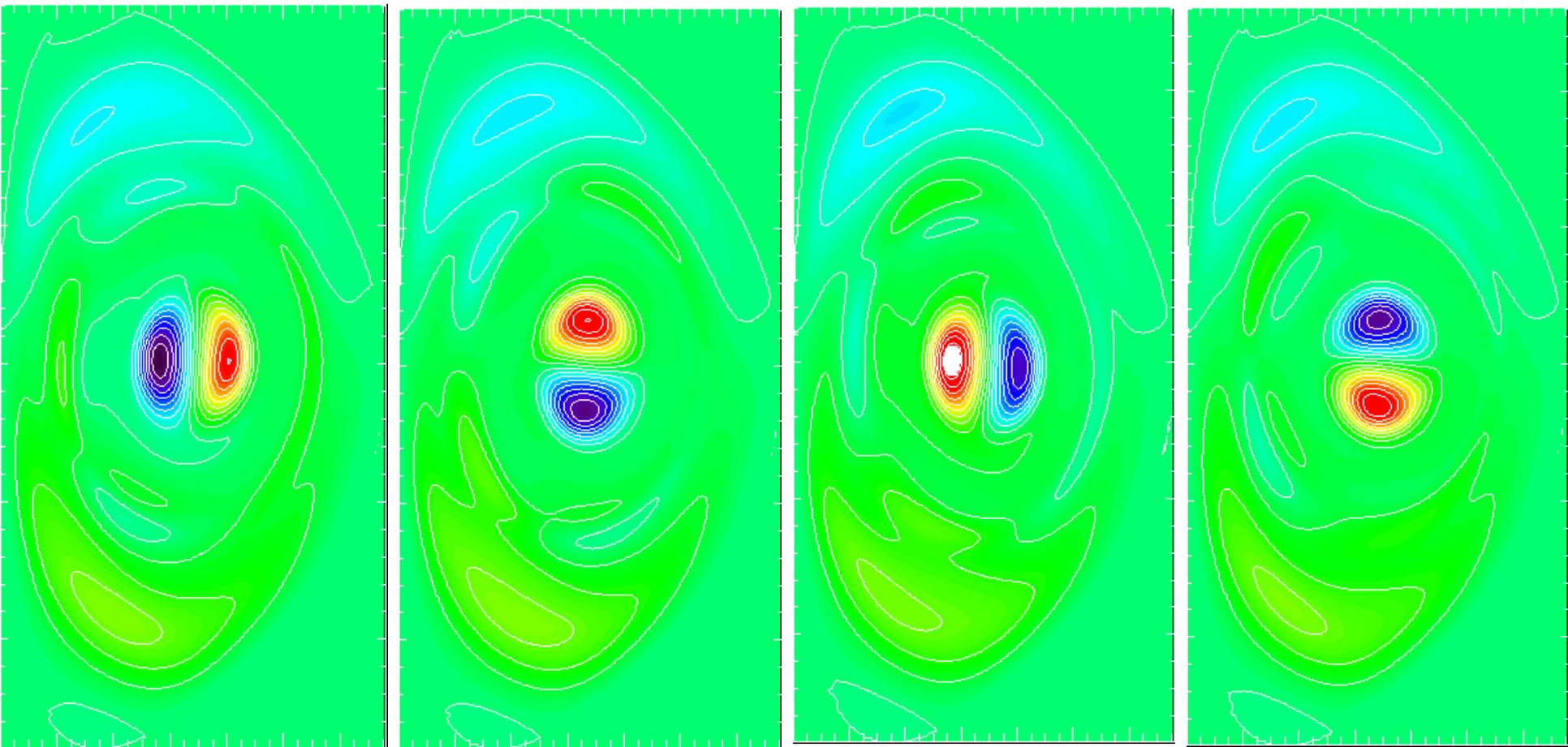
$\varphi = 0^\circ$

$\varphi = 90^\circ$

$\varphi = 180^\circ$

$\varphi = 270^\circ$

Stationary velocity stream function for DIII-D Oval (shot 118164)



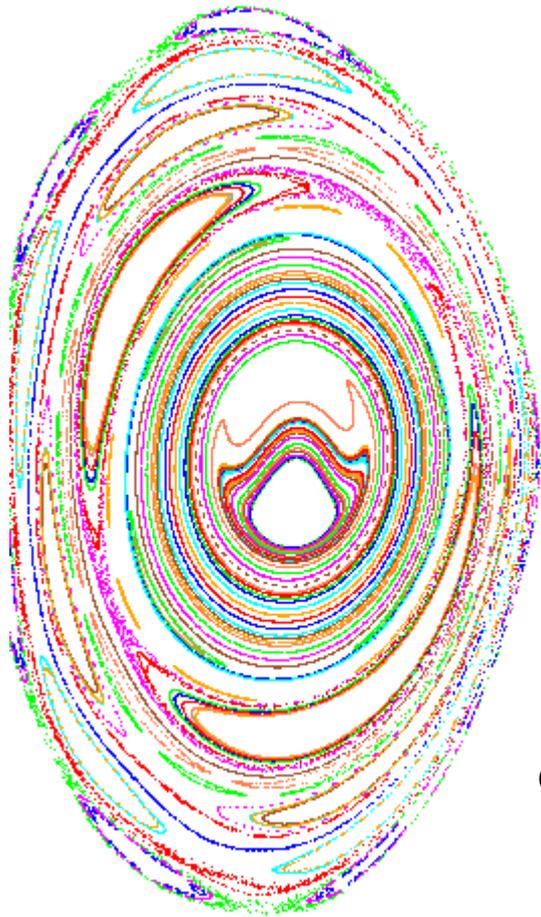
$\varphi = 0^\circ$

$\varphi = 90^\circ$

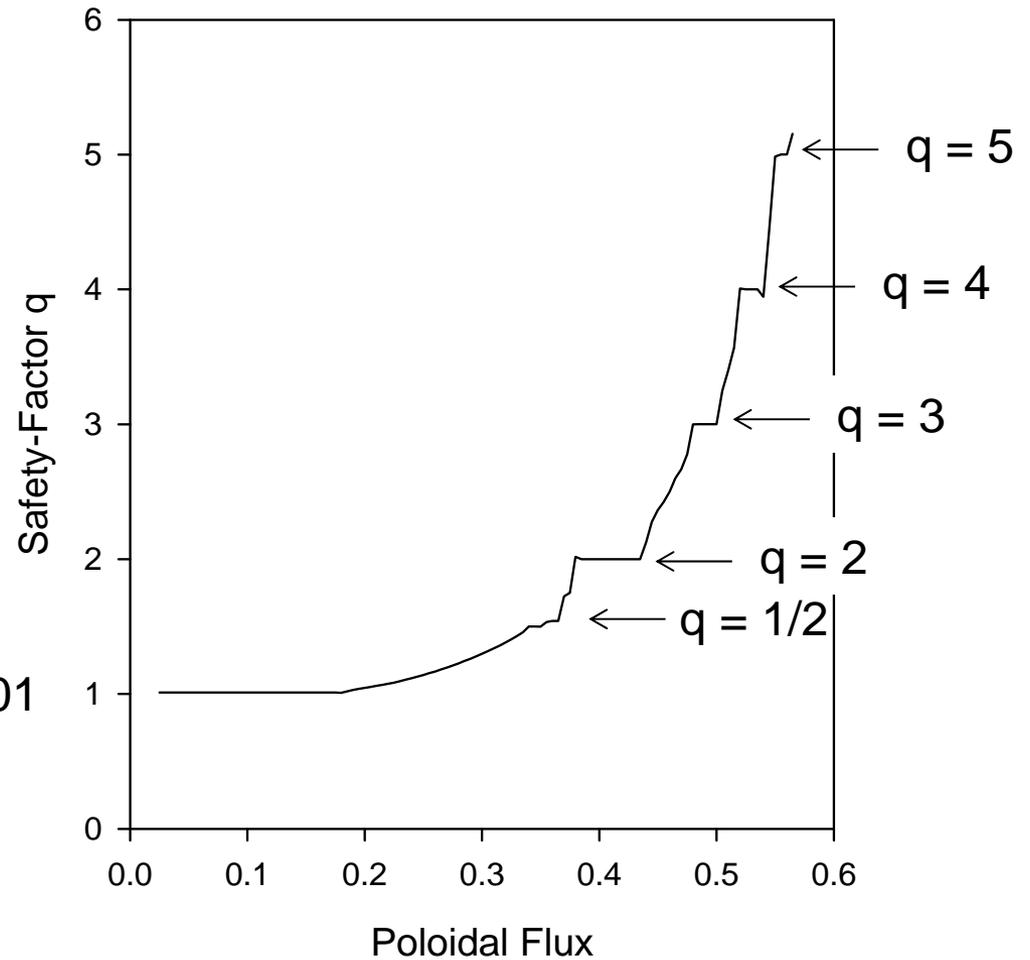
$\varphi = 180^\circ$

$\varphi = 270^\circ$

Oval stationary state has many islands



$q_{\min} = 1.01$

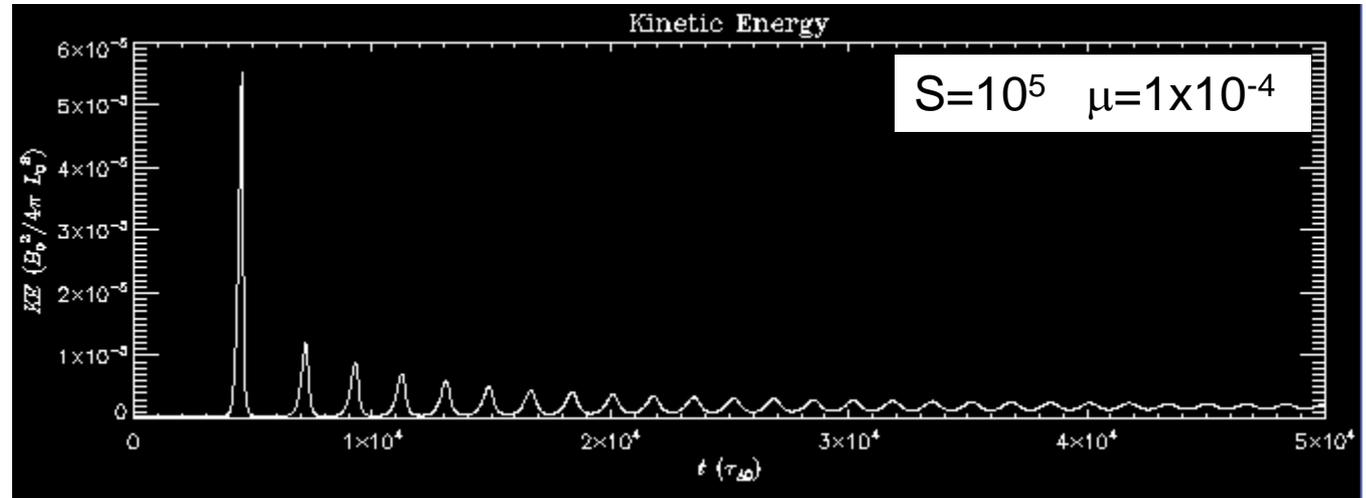


Summary

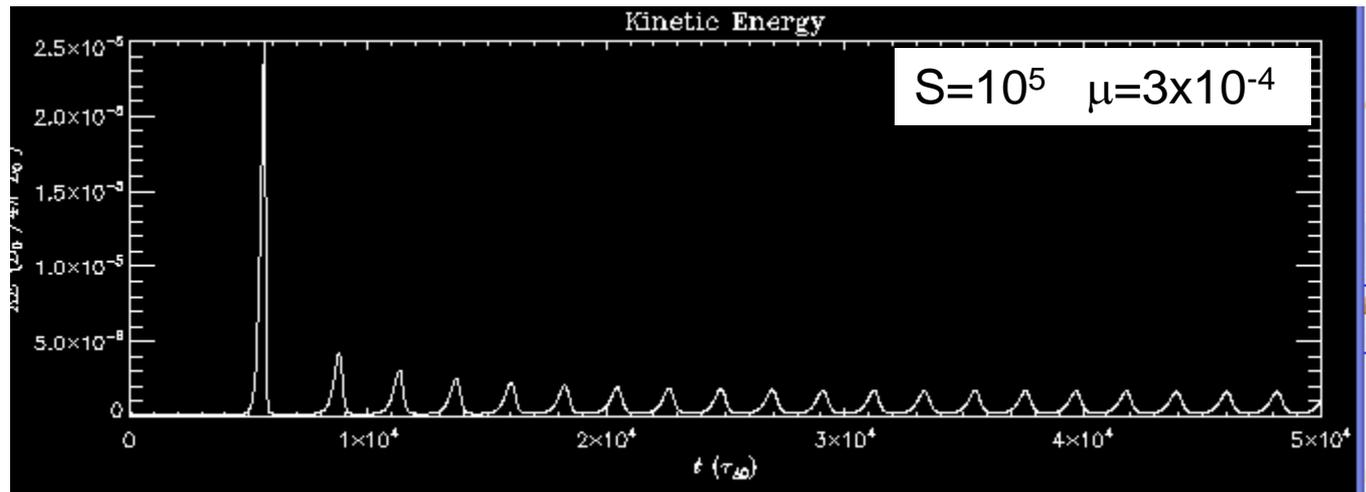
- 3D MHD in a highly magnetized high temperature plasma
 - Multiple timescales (ideal, reconnection, transport) demand implicit time advance
 - Implicit matrix contains large range of eigenvalues associated with the 3 different MHD wave types
- 3-step physics based preconditioner employed
 - Split implicit method reduces matrix size by 2 and makes matrix near symmetric and diagonally dominant
 - Annihilation operators approximately split matrix into 3 diagonal blocks, each with a greatly reduced condition number
 - Block Jacobi preconditioner dramatically reduces the condition number of each of the diagonal blocks
 - Final preconditioned matrix given to GMRES converges in 10s of iterations for fine zoned problem
- Recent Results
 - Repeating sawtooth demonstrate multiple timescale calculations
 - Stationary helical state can exist for some transport parameters

Extra Viewgraphs

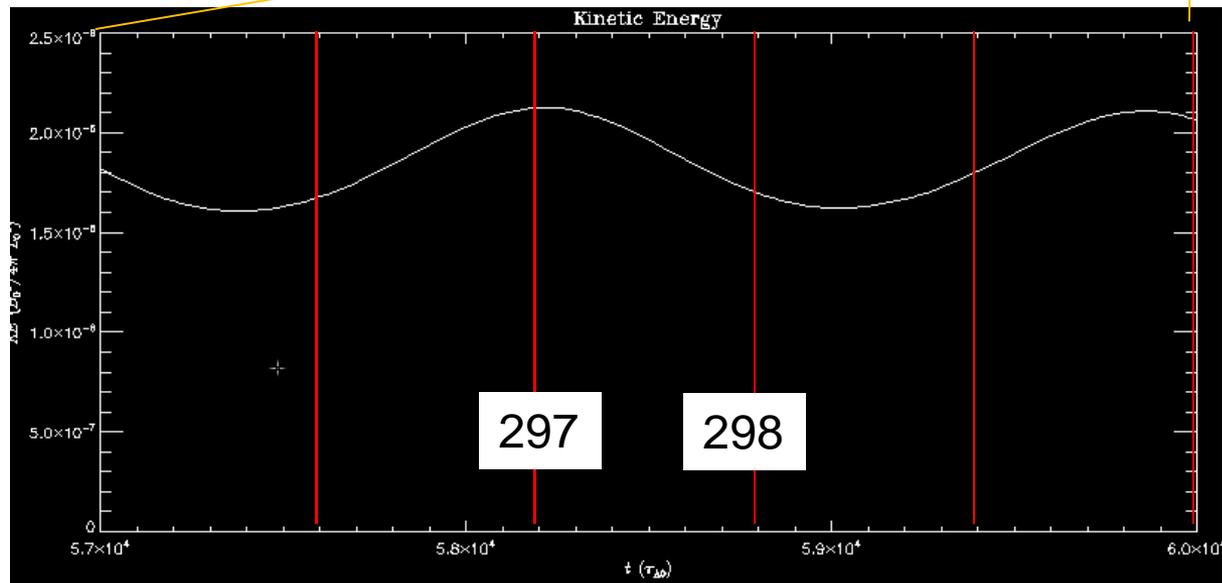
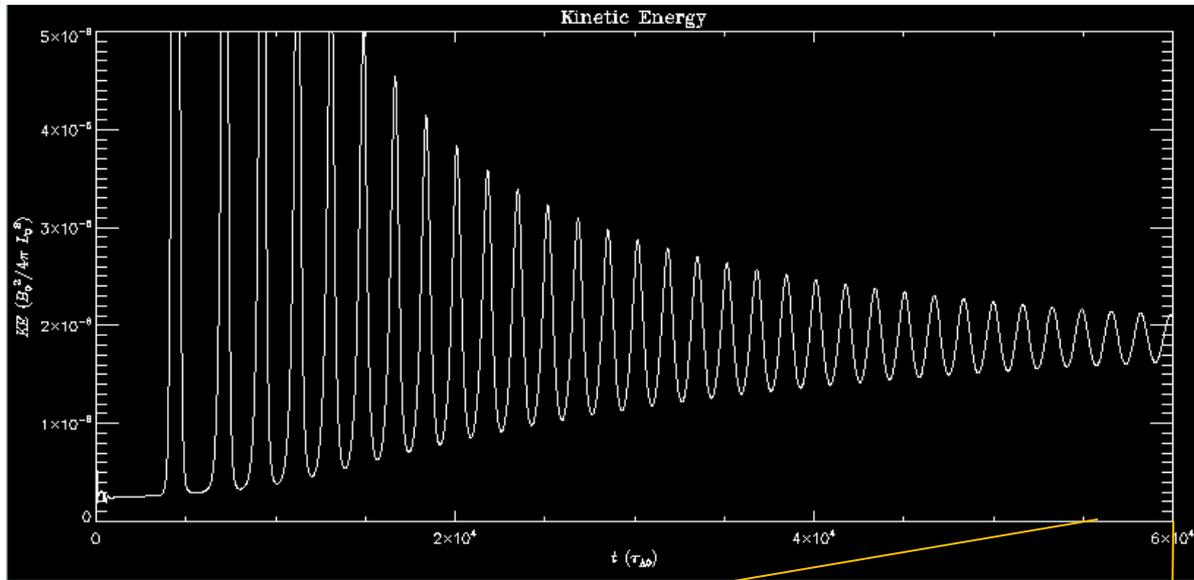
In low viscosity runs, the periodic sawtooth dies out and a helical steady state is formed



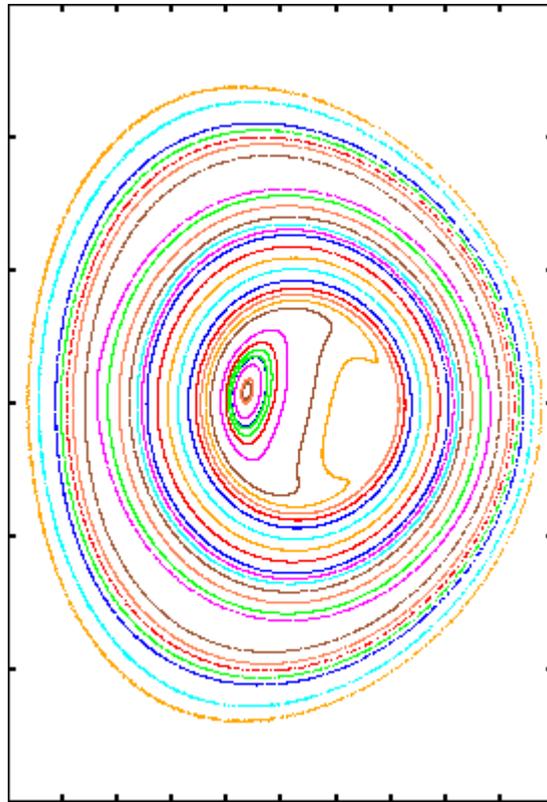
In higher viscosity runs, the periodic sawtooth remains



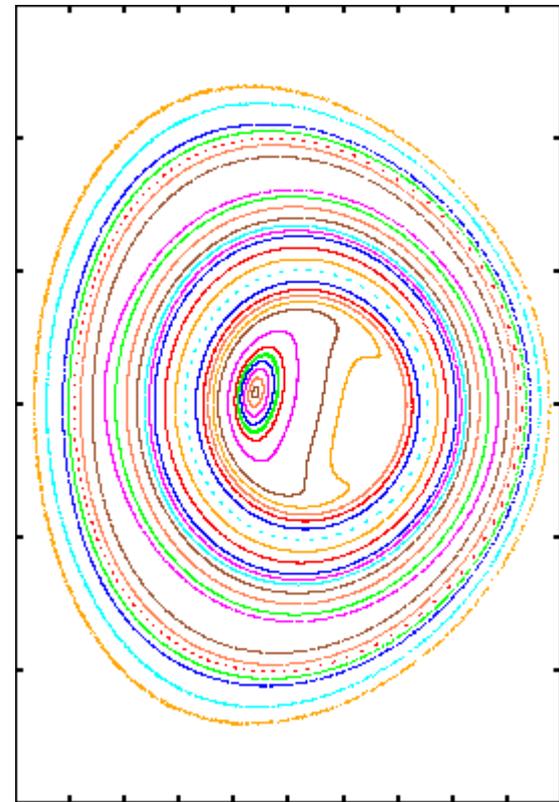
Low viscosity run settles down to stationary state where KE is constant in time



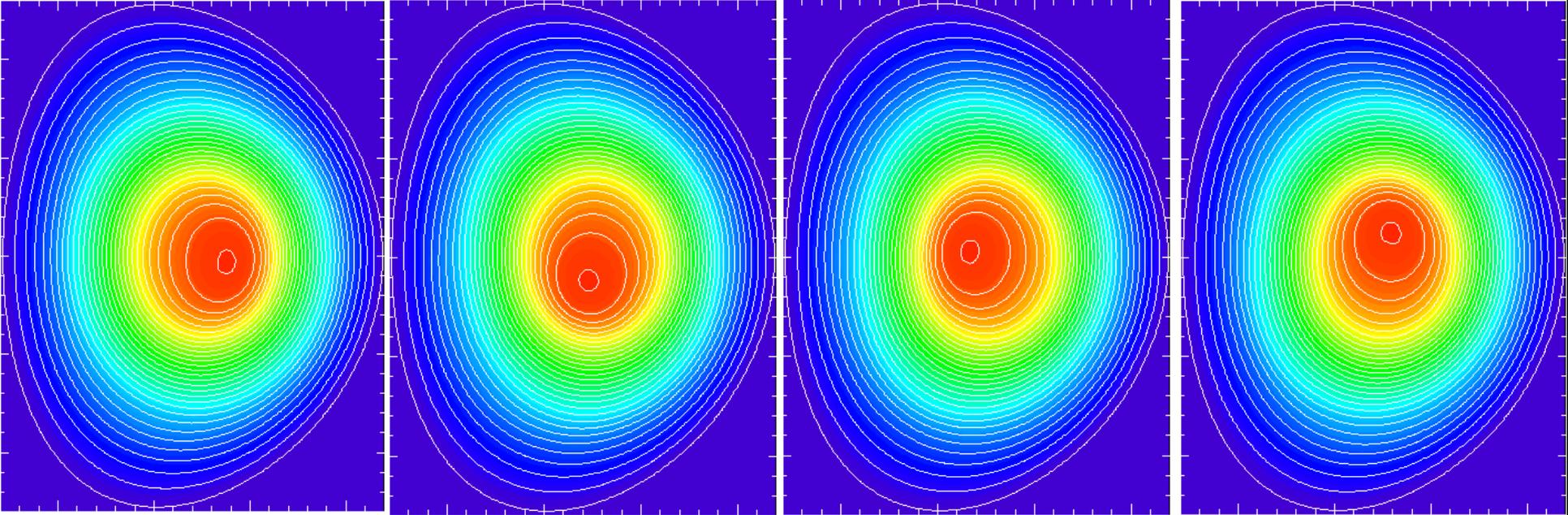
Poincare plots are very similar at peak and valley of kinetic energy.
Two magnetic axes. Large region in center where $q \sim 1$



297



298

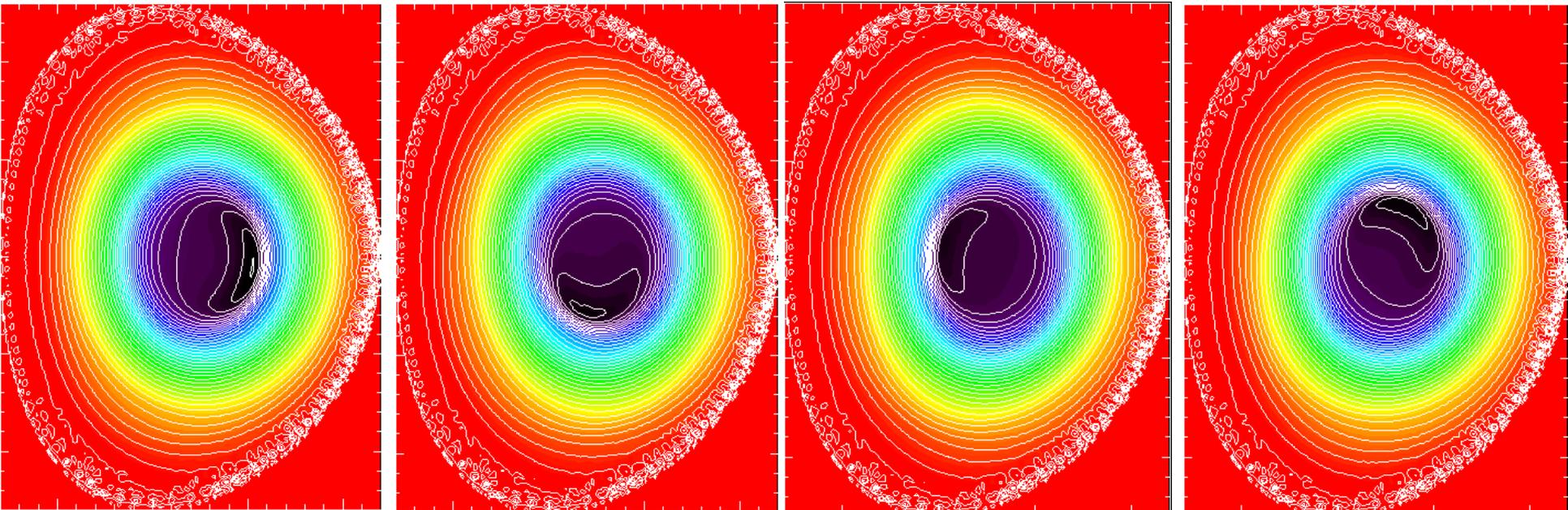


$\varphi=0$

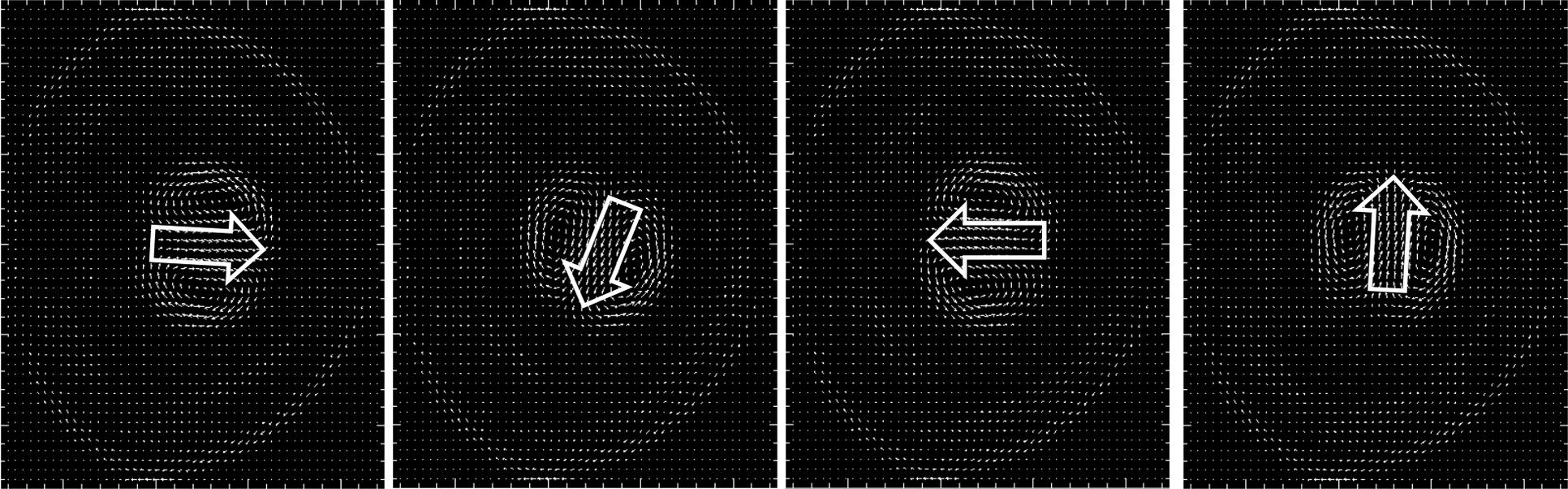
$\varphi=\pi/2$

$\varphi=\pi$

$\varphi=3\pi/2$



Helical distortion in central pressure (top) and toroidal current density

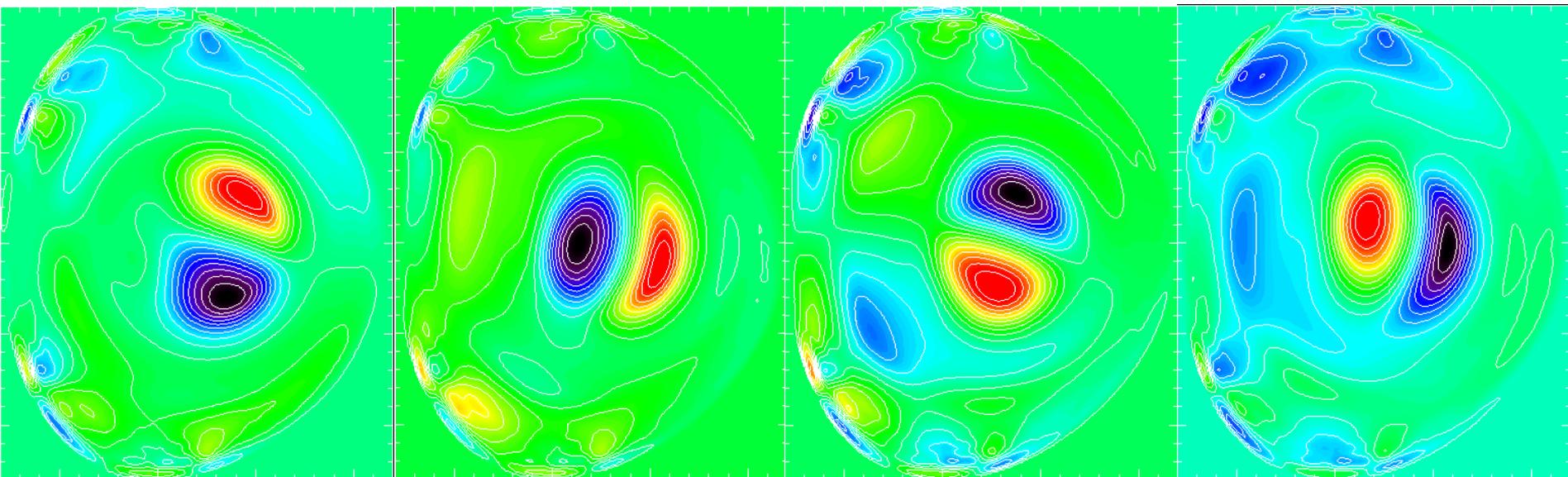


$\varphi=0$

$\varphi=\pi/2$

$\varphi=\pi$

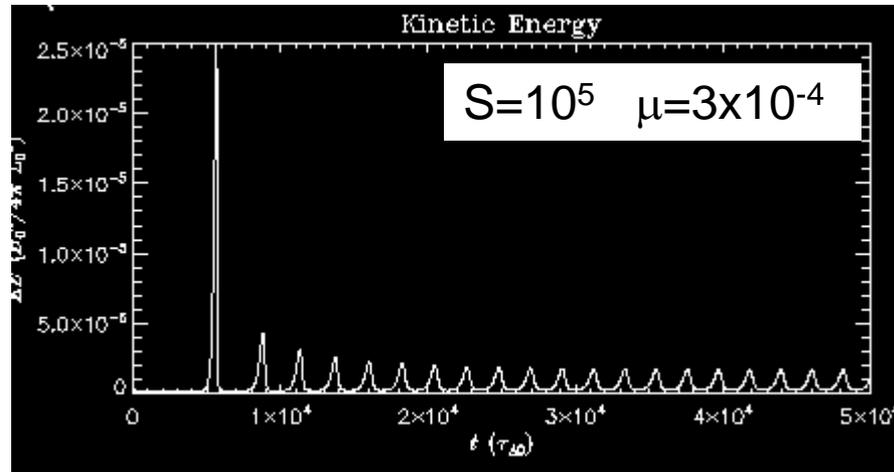
$\varphi=3\pi/2$



Poloidal velocity (top) and toroidal velocity (bottom) form stationary helical structures

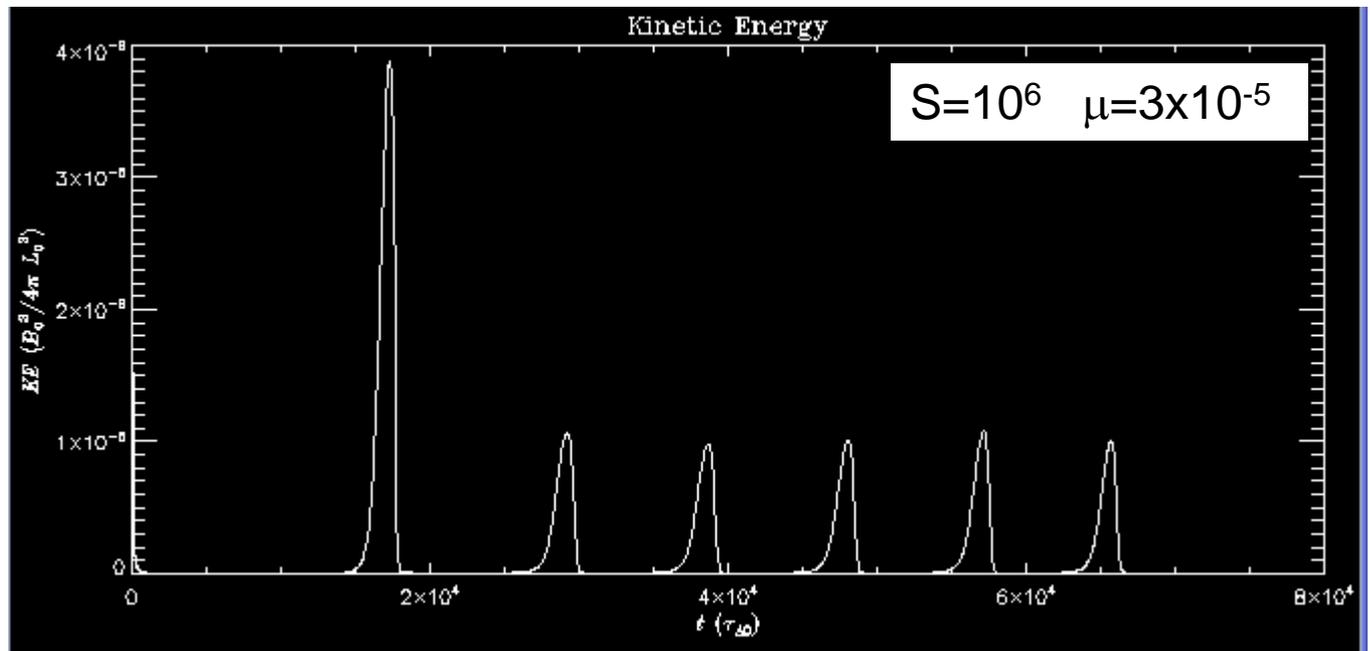
Comparison of sawtooth cycles for $S=10^5$ and $S=10^6$

Period:
 $\sim 2000\tau_A$



η/μ constant

Period:
 $\sim 9000\tau_A$



Form of the Vector Fields

We express magnetic field in terms of two scalar fields such that $\nabla \cdot \mathbf{B} \equiv 0$ and it reduces to the standard form in 2D

$$\mathbf{A} = R^2 \nabla \phi \times \nabla f + \psi \nabla \phi - F_0 \ln R \hat{Z}$$

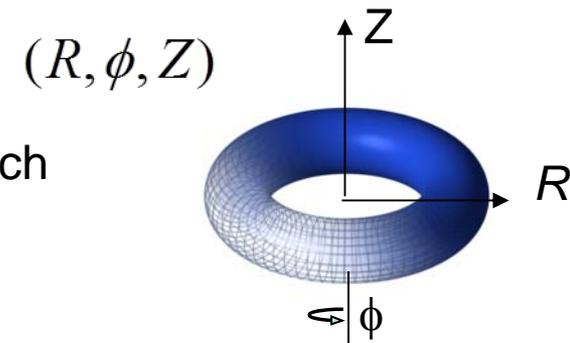
$$\begin{aligned} \mathbf{B} &= \nabla \psi \times \nabla \phi - \nabla_{\perp} \frac{\partial f}{\partial \phi} + (F_0 + R^2 \nabla_{\perp}^2 f) \nabla \phi \\ &= \nabla \psi \times \nabla \phi - \nabla \frac{\partial f}{\partial \phi} + (F_0 + R^2 \nabla^2 f) \nabla \phi \end{aligned}$$

$$\mathbf{J} = \nabla (R^2 \nabla^2 f) \times \nabla \phi + \frac{1}{R^2} \nabla_{\perp} \frac{\partial \psi}{\partial \phi} - \Delta^* \psi \nabla \phi$$

In a standard tokamak, these satisfy the orderings:

$$f \ll \psi \ll F_0$$

This ordering is used, not to neglect terms, but to devise efficient and accurate numerical methods.



Field from TF coils:

$$F_0 = \frac{\mu_0 I_{TF}}{2\pi}$$

Gauge condition:

$$R^2 \nabla_{\perp} \cdot \frac{1}{R^2} \mathbf{A} = 0$$

$$\nabla_{\perp} \equiv \nabla - \nabla \phi \frac{\partial}{\partial \phi}$$

Momentum Equation Projections

Recall the form of the momentum equation we are using:

$$(ME) \quad \left\{ \mathbf{I} - \theta^2 (\delta t)^2 L \right\} \mathbf{V}^{n+1} = \left\{ \mathbf{I} - \theta^2 (\delta t)^2 L \right\} \mathbf{V}^n + \frac{\delta t}{\rho} \left\{ -\nabla p + \frac{1}{\mu_0} (\nabla \times \mathbf{B}) \times \mathbf{B} \right\}^{n+1/2} + \dots$$

To apply the Galerkin method we take the following annihilation projections, multiply by v_i , the i^{th} finite element basis function, and integrate over the domain

$$\iint d^3\tau v_i \nabla \phi \cdot \nabla_{\perp} \times R^2 (ME) \quad \rightarrow \quad \iint d^3\tau R^2 \nabla_{\perp} v_i \times \nabla \phi \cdot (ME)$$

$$\iint d^3\tau v_i R^2 \nabla \phi \cdot (ME) \quad \rightarrow \quad \iint d^3\tau v_i R^2 \nabla \phi \cdot (ME)$$

$$-\iint d^3\tau v_i \nabla_{\perp} \cdot R^{-2} (ME) \quad \rightarrow \quad \iint d^3\tau R^{-2} \nabla_{\perp} v_i \cdot (ME)$$

Comparing expressions on right with the velocity field, we see that this is equivalent to multiplying by each term in the velocity.

$$\mathbf{V} = R^2 \nabla \mathbf{U} \times \nabla \phi + R^2 \boldsymbol{\omega} \nabla \phi + \frac{1}{R^2} \nabla_{\perp} \chi \quad \int \mathbf{V} \cdot \mathbf{L} = \delta W$$

Leads to (1) Discrete energy conserving equations, (2) Two energy conserving subsets (reduced MHD). Also, orthogonality property leads to (3) Separation of ⁵⁹ incompressible physics and (4) Well conditioned, diagonally dominant matrices.