On the Distribution of Subsidence in the Hurricane Eye

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> IPAM: The Hurricane I2 February 2007

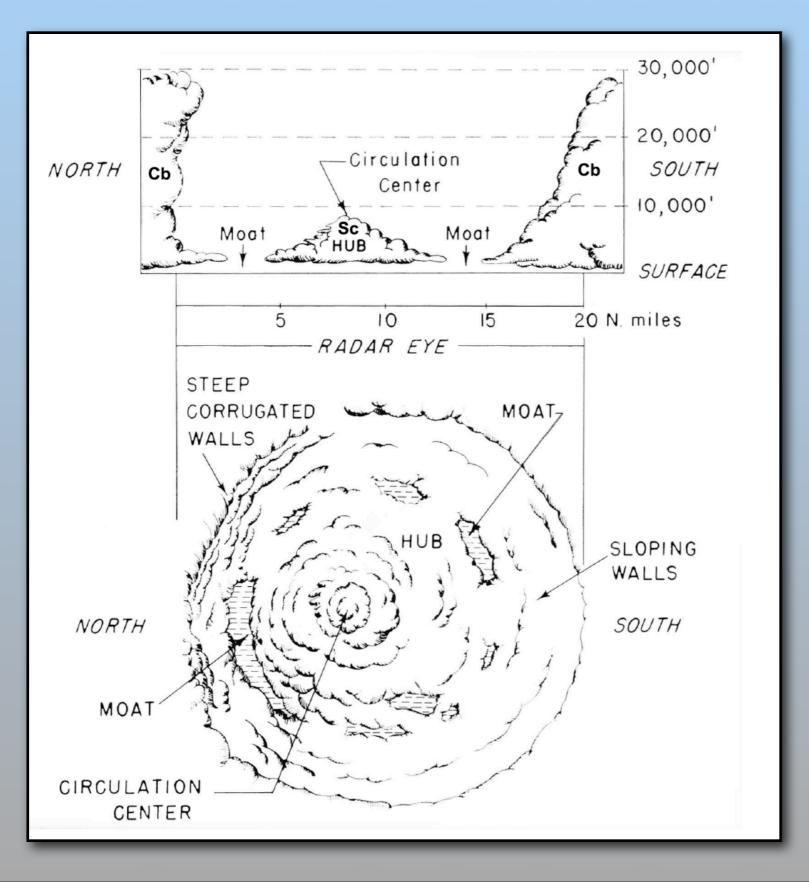
Hub Clouds, Moats, and Warm Rings

Simpson, R.H., and L.G. Starrett, 1955: Bull.Amer. Meteor. Soc., 36, 459-468.

Fletcher, R.D., J.R. Smith, and R.C. Bundgaard, 1961: Weatherwise, 14, 102-109.

LaSeur, N.E., and H.F. Hawkins, 1963: Mon. Wea. Rev., 91, 694-709.

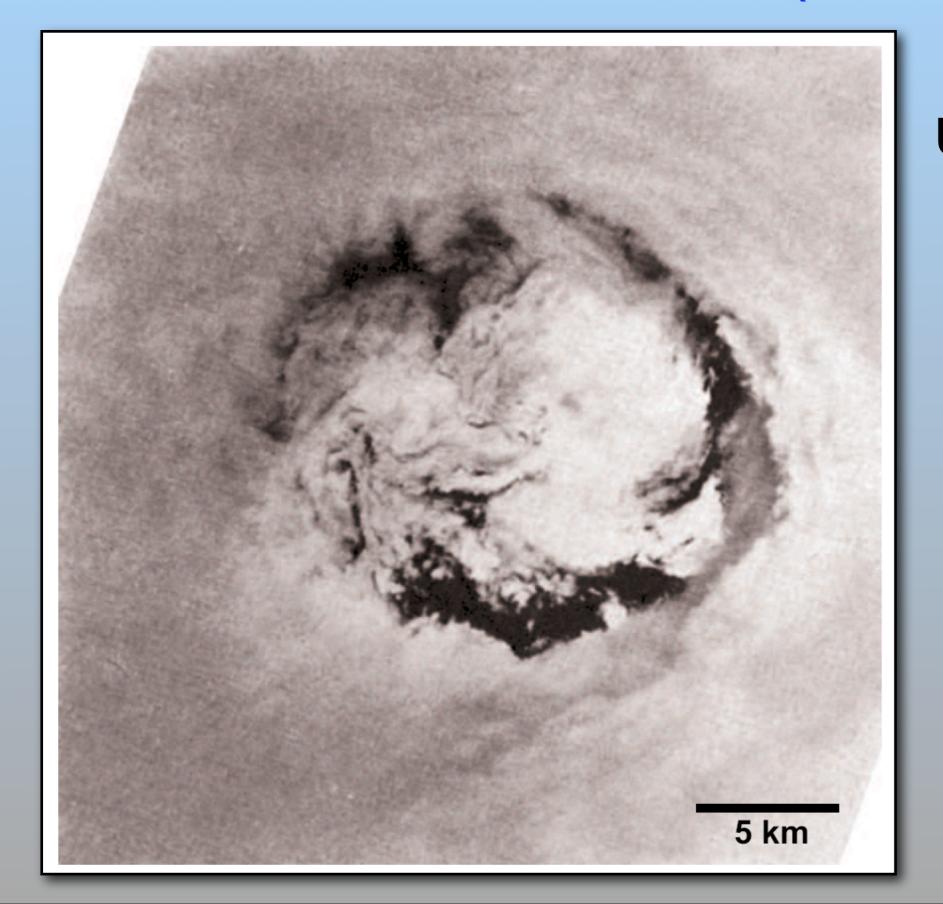
Simpson and Starrett (1955)



Schematic Diagram of the Eye of Hurricane Edna

9-10 Sept 1954

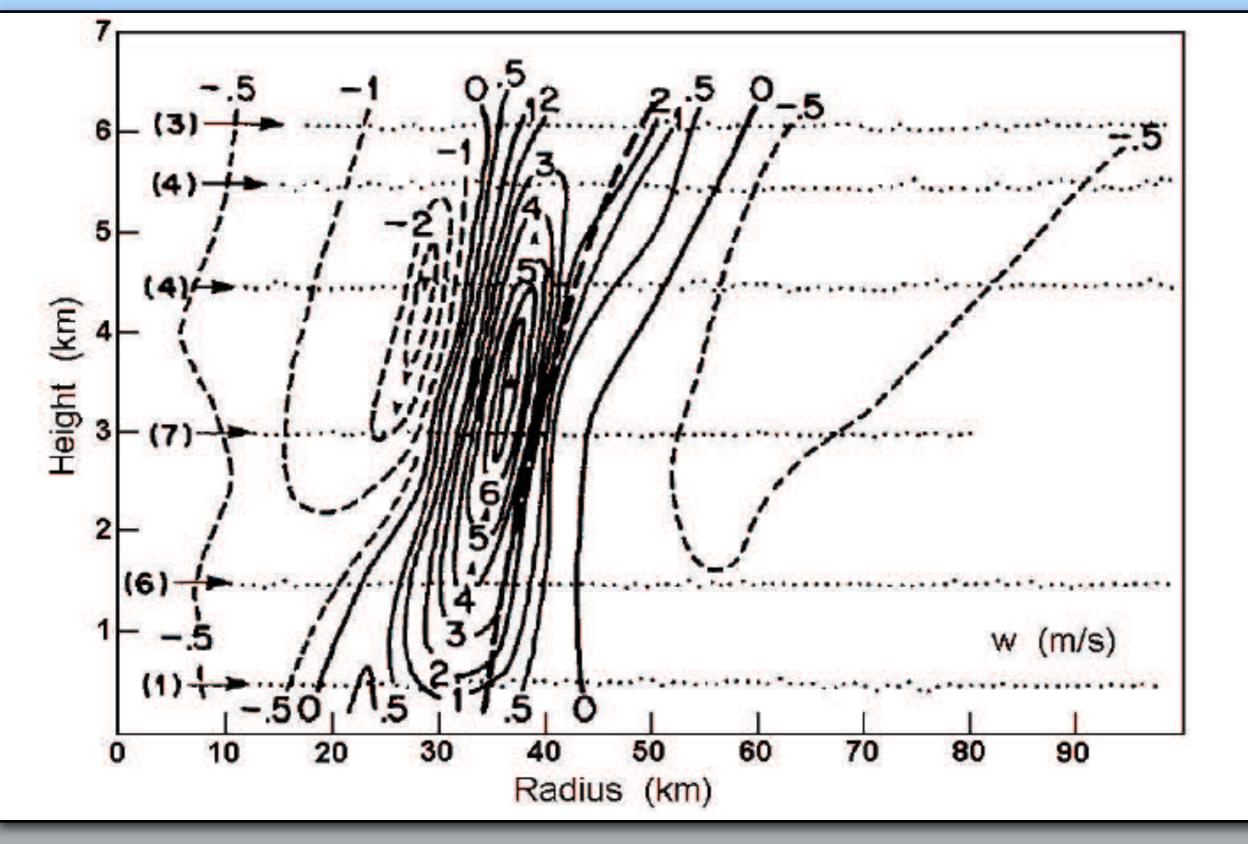
Fletcher et al. (1961)



U-2 Photograph of Typhoon Ida

25 Sept 1958

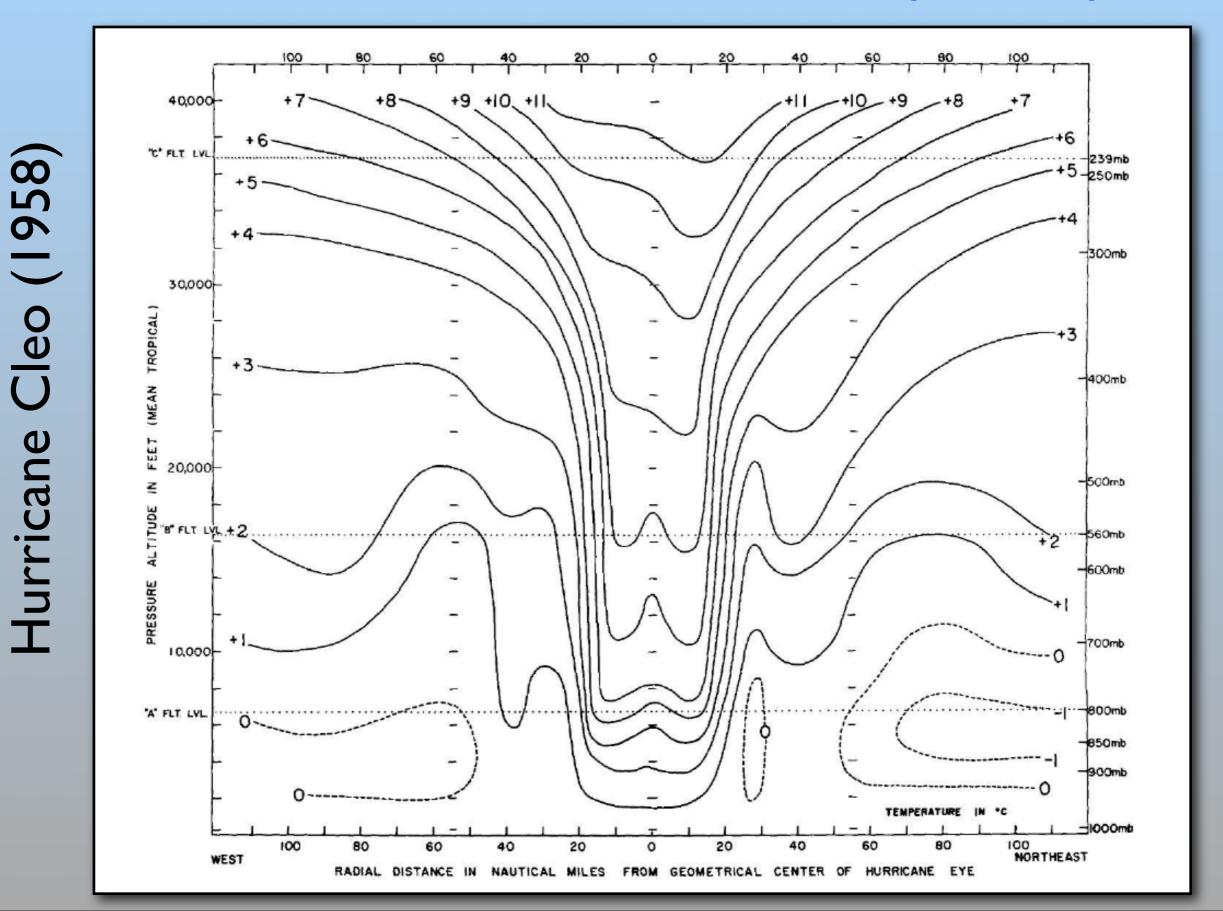
Hurricane Allen Vertical Velocities



5 Aug 1980

Jorgensen (1984)

LaSeur and Hawkins (1963)



Balanced Vortex Model

Assumptions:

Axisymmetric
Inviscid
Hydrostatic
Gradient Balanced

Vertical Coordinate:

$$z = H \ln\left(\frac{p_0}{p}\right)$$

where

$$H = \frac{RT_0}{g}$$

constant scale height

Balanced Vortex Model

$$\left(f + \frac{v}{r}\right)v = \frac{\partial\phi}{\partial r}$$
$$\frac{Dv}{Dt} + \left(f + \frac{v}{r}\right)u = 0$$
$$\frac{\partial\phi}{\partial z} = \frac{g}{T_0}T$$
$$\frac{\partial(ru)}{r\partial r} + \frac{\partial w}{\partial z} - \frac{w}{H} = 0$$
$$c_p\frac{DT}{Dt} + \frac{RT}{H}w = Q$$

Gradient Balance Momentum Eq. Hydrostatic Eq. Continuity Eq. Thermodynamic Eq. Streamfunction for the Transverse Circulation

The continuity equation

$$\frac{\partial(ru)}{r\partial r} + \frac{\partial w}{\partial z} - \frac{w}{H} = 0$$

is satisfied by

$$e^{-z/H}u = -\frac{\partial \psi}{\partial z}$$
 and $e^{-z/H}w = \frac{\partial (r\psi)}{r\partial r}$

Thermal Wind Equation

 Derived from the gradient balance and hydrostatic equations:

$$\left(f + \frac{2v}{r}\right)\frac{\partial v}{\partial z} = \frac{g}{T_0}\frac{\partial T}{\partial r}$$

Taking the local time derivative:

$$\frac{\partial}{\partial z} \left\{ \left(f + \frac{2v}{r} \right) \frac{\partial v}{\partial t} \right\} = \frac{g}{T_0} \frac{\partial}{\partial r} \left(\frac{\partial T}{\partial t} \right)$$

Deriving the Transverse Circulation Equation

 Thermodynamic equation yields:

$$\frac{g}{T_0}\frac{\partial T}{\partial t} + A\frac{\partial(r\psi)}{r\partial r} + B\frac{\partial\psi}{\partial z} = \frac{g}{c_pT_0}Q$$

Tangential wind equation yields:

$$-\left(f+\frac{2v}{r}\right)\frac{\partial v}{\partial t} + B\frac{\partial(r\psi)}{r\partial r} + C\frac{\partial\psi}{\partial z} = 0$$

Variable coefficients:

$$\begin{split} A &= e^{z/H} \frac{g}{T_0} \left(\frac{\partial T}{\partial z} + \frac{\kappa T}{H} \right) \\ B &= -e^{z/H} \left(f + \frac{2v}{r} \right) \frac{\partial v}{\partial z} = -e^{z/H} \frac{g}{T_0} \frac{\partial T}{\partial r} \\ C &= e^{z/H} \left(f + \frac{2v}{r} \right) \left(f + \frac{\partial (rv)}{r \partial r} \right) \end{split}$$

Static Stability

Baroclinity

Inertial Stability

Sawyer-Eliassen Transverse Circulation Equation

$$\frac{\partial}{\partial r} \left(A \frac{\partial (r\psi)}{r \partial r} + B \frac{\partial \psi}{\partial z} \right) + \frac{\partial}{\partial z} \left(B \frac{\partial (r\psi)}{r \partial r} + C \frac{\partial \psi}{\partial z} \right) = \frac{g}{c_p T_0} \frac{\partial Q}{\partial r}$$

Boundary Conditions:

$$\psi = 0$$
 if $r = 0$, or $z = 0, z_T$
 $r\psi \to 0$ as $r \to \infty$

Elliptic if:

$$AC - B^2 > 0$$

Idealized Vortex

Consider an idealized vortex such that:

$$A = e^{z/H} N^2$$
 (Assume N is constant)
 $B = 0$ (Assume vortex is barotropic

Output The inertial stability can then be written:

$$C = e^{z/H} \hat{f}^2$$

where

$$\hat{f}(r) = \left[\left(f + \frac{2v}{r} \right) \left(f + \frac{\partial(rv)}{r\partial r} \right) \right]^{1/2}$$

effective Coriolis parameter

Simplified Transverse Circulation Equation

$$N^2 \frac{\partial}{\partial r} \left(\frac{\partial (r\psi)}{r \partial r} \right) + \hat{f}^2 e^{-z/H} \frac{\partial}{\partial z} \left(e^{z/H} \frac{\partial \psi}{\partial z} \right) = \frac{g e^{-z/H}}{c_p T_0} \frac{\partial Q}{\partial r}$$

Now assume separable solutions:

$$Q(r, z) = \hat{Q}(r) \exp\left(\frac{z}{2H}\right) \sin\left(\frac{\pi z}{z_T}\right)$$
$$\psi(r, z) = \hat{\psi}(r) \exp\left(-\frac{z}{2H}\right) \sin\left(\frac{\pi z}{z_T}\right)$$

The above PDE now reduces to an ODE.

Radial Structure Equation

$$r^{2}\frac{d^{2}\hat{\psi}}{dr^{2}} + r\frac{d\hat{\psi}}{dr} - \left(\mu^{2}r^{2} + 1\right)\hat{\psi} = \frac{gr^{2}}{c_{p}T_{0}N^{2}}\frac{d\hat{Q}}{dr}$$

where

$$\mu = \frac{\hat{f}}{N} \left(\frac{\pi^2}{z_T^2} + \frac{1}{4H^2} \right)^{1/2}$$
 inverse Rossby length

 We are particularly interested in the important role played by radial variations of the inverse Rossby length.

Idealized Barotropic Vortex

Now consider the following barotropic vortex:

$$rv(r) = \frac{1}{2} \begin{cases} \left(\hat{f}_0 - f\right)r^2, & \text{if } 0 \le r \le r_1 \\ \left\{\hat{f}_0^2 r_1^4 + \hat{f}_1^2 \left(r^4 - r_1^4\right)\right\}^{1/2} - fr^2, & \text{if } r_1 \le r \le r_2 \\ \left\{\hat{f}_0^2 r_1^4 + \hat{f}_1^2 \left(r_2^4 - r_1^4\right) + \hat{f}_2^2 \left(r^4 - r_2^4\right)\right\}^{1/2} - fr^2, & \text{if } r_2 \le r < \infty \end{cases}$$

 The resulting effective Coriolis parameter is piecewise constant:

$$\hat{f}(r) = \left\{ \left(f + \frac{2v}{r}\right) \left(f + \frac{\partial(rv)}{r\partial r}\right) \right\}^{1/2} = \begin{cases} \hat{f}_0, & \text{if } 0 \le r < r_1 \text{ (eye)} \\ \hat{f}_1, & \text{if } r_1 < r < r_2 \text{ (eyewall)} \\ \hat{f}_2, & \text{if } r_2 < r < \infty \text{ (far-field)} \end{cases}$$

Diabatic Heating

Now assume the heating is confined to the eyewall region:

$$\hat{Q}(r) = \begin{cases} 0 & \text{if } 0 \le r < r_1 \text{ (eye)} \\ Q_1 & \text{if } r_1 < r < r_2 \text{ (eyewall)} \\ 0 & \text{if } r_2 < r < \infty \text{ (far-field)} \end{cases}$$

 Parameter constraint (fixes the areaaveraged rainfall):

$$(Q_1/c_p)(r_2^2 - r_1^2) = 125 \,\mathrm{K} \,\mathrm{day}^{-1}(50 \,\mathrm{km})^2$$

Resulting Radial Structure Equations

Governing equation for each region:

$$r^{2}\frac{d^{2}\hat{\psi}}{dr^{2}} + r\frac{d\hat{\psi}}{dr} - (\mu_{j}^{2}r^{2} + 1)\hat{\psi} = 0$$

$$j=0 \text{ eye}$$

$$j=1 \text{ eyewal}$$

$$j=2 \text{ far-field}$$

Jump conditions at the edges of the eyewall:

$$\left[\frac{d(r\hat{\psi})}{r\,dr}\right]_{r_1-}^{r_1+} = \frac{gQ_1}{c_pT_0N^2}$$

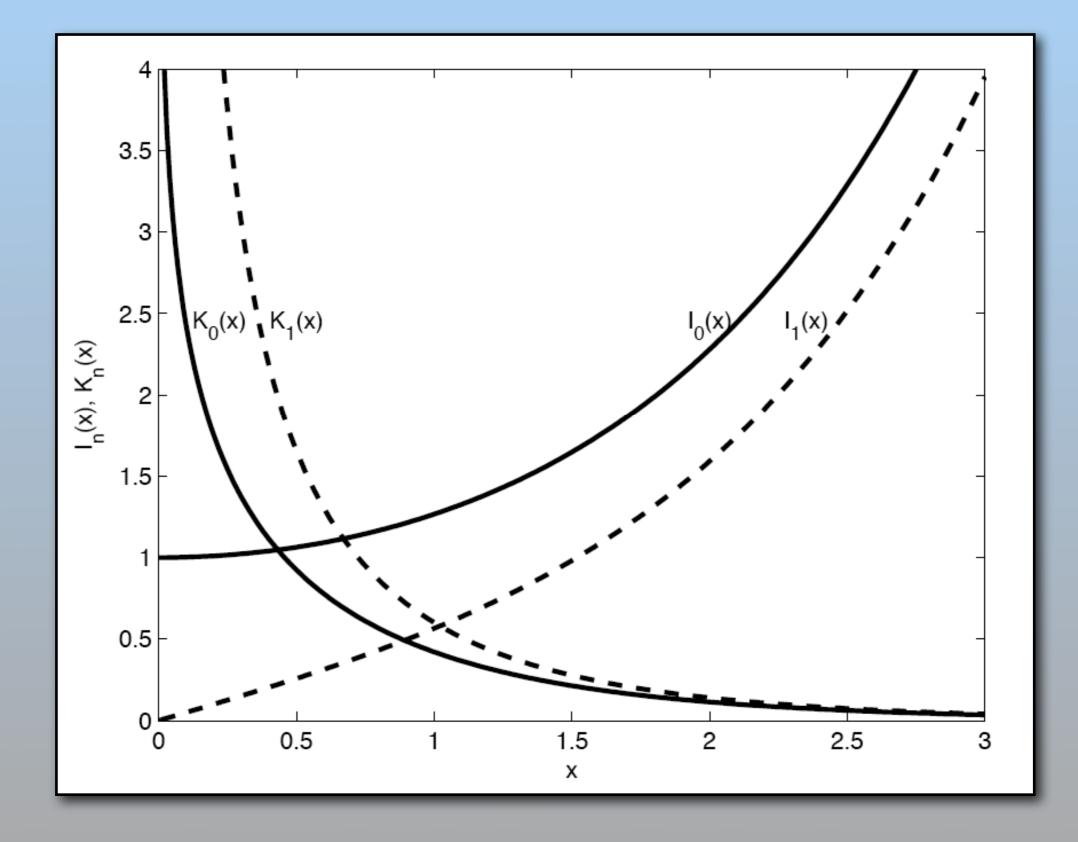
$$\left[\frac{d(r\hat{\psi})}{r\,dr}\right]_{r_2-}^{r_2+} = -\frac{gQ_1}{c_pT_0N^2}$$

(inner edge of eyewall)

(outer edge of eyewall)

With our heating profile, solutions are simply the first-order modified Bessel functions.

Modified Bessel Functions



Streamfunction Solution

$$\hat{\psi}(r) = \begin{cases} \hat{\psi}_1 I_1(\mu_0 r) / I_1(\mu_0 r_1) & \text{if } 0 \le r \le r_1 \\ \{\hat{\psi}_1 F(r, r_2) + \hat{\psi}_2 F(r_1, r)\} / F(r_1, r_2) & \text{if } r_1 \le r \le r_2 \\ \hat{\psi}_2 K_1(\mu_2 r) / K_1(\mu_2 r_2) & \text{if } r_2 \le r < \infty \end{cases}$$

where $F(x,y) = I_1(\mu_1 x)K_1(\mu_1 y) - K_1(\mu_1 x)I_1(\mu_1 y)$

- The parameters $\hat{\psi}_1$ and $\hat{\psi}_2$ are determined by the two jump conditions.
- Note that this solution is continuous at the edges of the eyewall.

Vertical Motion Solution

 Note the following derivative formulas for modified Bessel functions:

$$\frac{d[rI_1(\mu r)]}{rdr} = \mu I_0(\mu r)$$

$$\frac{d[rK_1(\mu r)]}{rdr} = -\mu K_0(\mu r)$$

Output The vertical motion solution then becomes:

$$\frac{d(r\hat{\psi})}{r\,dr} = \begin{cases} \hat{\psi}_1 \mu_0 I_0(\mu_0 r) / I_1(\mu_0 r_1) & \text{if } 0 \le r < r_1 \\ \{\hat{\psi}_1 \mu_1 G(r, r_2) - \hat{\psi}_2 \mu_1 G(r, r_1)\} / F(r_1, r_2) & \text{if } r_1 < r < r_2 \\ -\hat{\psi}_2 \mu_2 K_0(\mu_2 r) / K_1(\mu_2 r_2) & \text{if } r_2 < r < \infty \end{cases}$$

where
$$G(x, y) = I_0(\mu_1 x)K_1(\mu_1 y) + K_0(\mu_1 x)I_1(\mu_1 y)$$

Temperature Tendency

The temperature tendency for a barotropic vortex is given by:

$$\frac{\partial \hat{T}}{\partial t} = \frac{\hat{Q}}{c_p} - \frac{T_0 N^2}{g} \frac{d(r\hat{\psi})}{r dr}$$

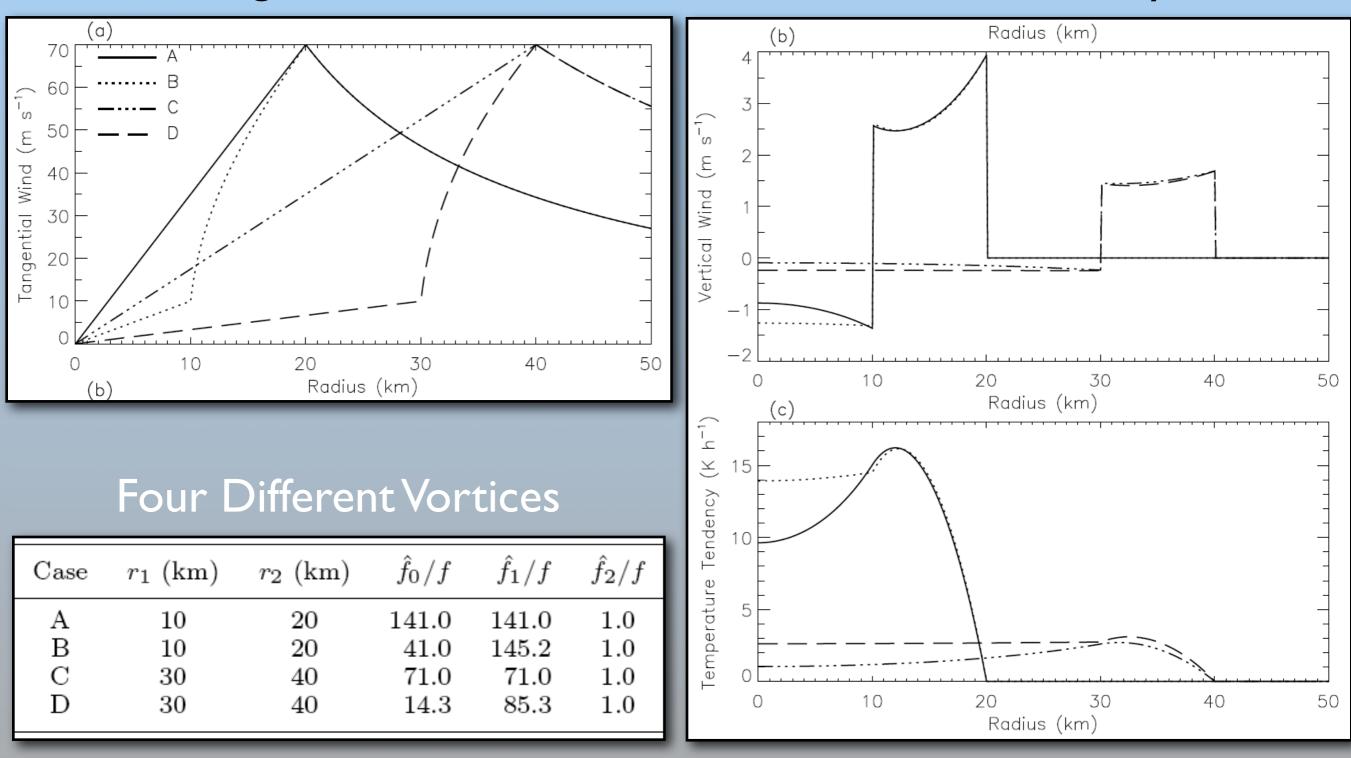
With our heating profile, this becomes:

$$\frac{\partial \hat{T}}{\partial t} = \frac{Q_1}{c_p} \begin{cases} \left\{ 1 - \left(\frac{1-\alpha}{1-\alpha\beta}\right)\mu_1 r_2 G(r_1, r_2) - \left(\frac{1-\beta}{1-\alpha\beta}\right) \right\} \frac{I_0(\mu_0 r)}{I_0(\mu_0 r_1)} & 0 \le r \le r_1 \\ 1 - \left(\frac{1-\alpha}{1-\alpha\beta}\right)\mu_1 r_2 G(r, r_2) - \left(\frac{1-\beta}{1-\alpha\beta}\right)\mu_1 r_1 G(r, r_1) & r_1 \le r \le r_2 \\ \left\{ 1 - \left(\frac{1-\alpha}{1-\alpha\beta}\right) - \left(\frac{1-\beta}{1-\alpha\beta}\right)\mu_1 r_1 G(r_2, r_1) \right\} \frac{K_0(\mu_2 r)}{K_0(\mu_2 r_2)} & r_2 \le r < \infty \end{cases}$$

Plots for Different Vortices

Tangential Wind

Vertical Velocity



Temperature Tendency

Hub Cloud of Hurricane Isabel



13 Sept 2003

Courtesy of Sim Aberson

Hurricane Isabel Flight-Level Data

Solid:Tangential WindDashed:TemperatureDotted:Dewpoint Temp.

13 Sept 2003

80 (a) 16 14m S⁻¹) 60 () () Tangential Wind (m emperature 2 40 \cap 20 6 30 10 20 40 50 60 0 Radius (km) 80 24 (b) 22 s⁻¹) 60 20 💬 E o a Temperature Tangential Wind 40 20 14 12 30 60 10 20 50 0 40 Radius (km)

m altitude

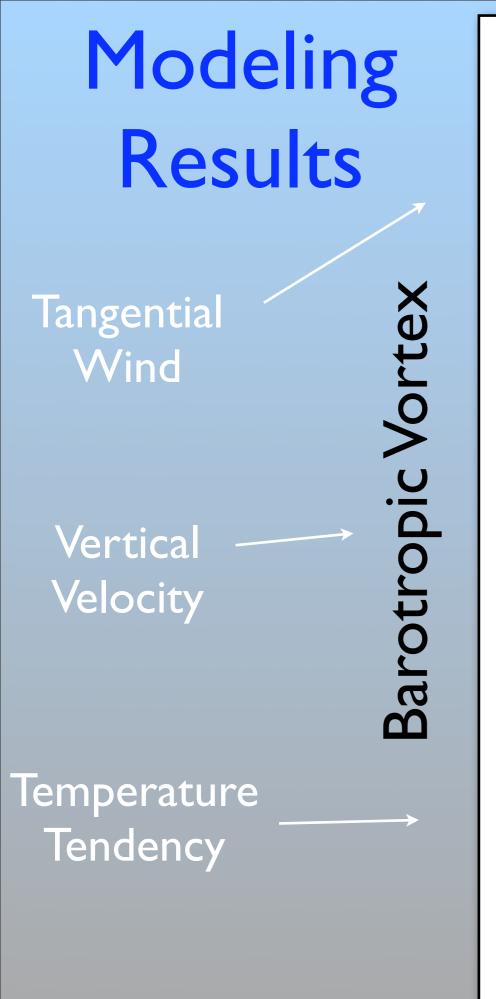
m altitude

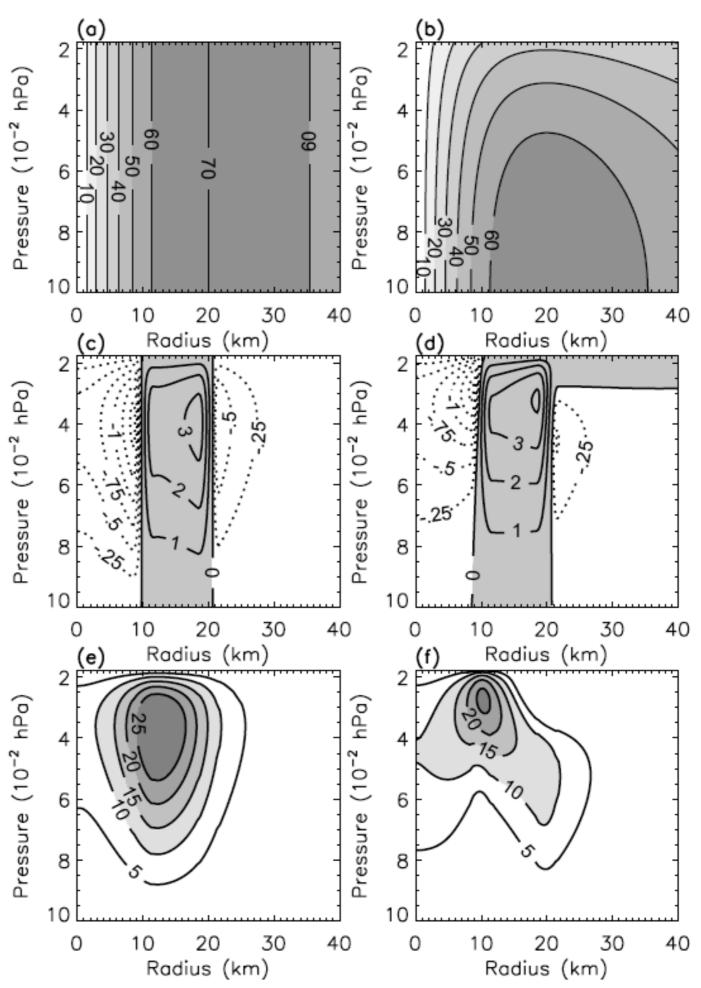
Conclusions

- There is less than 10% horizontal variation in the eye subsidence when the ratio of eye radius to Rossby length in the eye is less than 0.6
- The subsidence rate at the edge of the eye is more than twice as strong as that at the center of the eye when the ratio of eye radius to Rossby length in the eye is greater than 1.8 (i.e., large eyes and/or eyes with high inertial stability)

Conclusions (continued)

The existence of a hub cloud at the center of the eye, cascading pileus in the upper troposphere on the edge of the eye, a clear inner moat in the lower troposphere on the edge of the eye, and a warm-ring structure are all associated with strong inertial stability in the eye and a relatively large eye radius.





Baroclinic Vortex