

Enabling Adaptivity & Parallelism for Computational Relativity

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Introduction

- Dendro-GR: New framework for Computational relativity
 - High-degree of spatial adaptivity
 - High levels of parallelism
- Intermediate Mass Ratio Inspirals (IMRIs)
- Wavelet Adaptive Multi-Resolution (WAMR)
- First BBH Evolutions

IMBHs and IMRIs

- Binaries with intermediate mass ratios $10 \leq q \leq 100$
- IMBHs
 - Collapse of Pop III stars
 - Mergers in stellar clusters
 - Accretion onto stellar mass BHs
 - Collapse of gas clouds in the early universe
- LIGO has detected remnants with masses $\sim 20-60~M_{\odot}$
- Computational Challenge: Resolution

Mass ratio q = 100



• Lousto & Zlochower, PRL 106 041101 (2011)



Spherhake, Cardoso, Ott, Schnetter & Witek, PRD 84 084038 (2011).

Numerical Relativity

- Conventional AMR uses nested boxes
- Boxes don't naturally capture the geometry of binary black holes
- Numerical artefacts
- Computational Inefficiency
- Need unstructured grids
- Need supercomputers



Why Block Adaptivity is not enough



Octrees & Wavelets







Wavelet Adaptive Multiresolution



Octree-based AMR

- Axis-aligned subdivision of space
- In 2*D* each node has 4 children, 8 in 3*D*
- Provides high-levels of adaptivity while enabling simple and efficient data-structures, especially in parallel









- Wavelet adaptive multiresolution
- Unstructured Octree Grid
- High levels of fine-grained parallelism
- Automatic code-generation via symbolic interface
- Extensible
- Portable and highly-scalable on modern supercomputers



Parallelism



- Top-down algorithm for constructing octree
- Only leaf nodes are storedlinear octree
- Leaves are ordered according to Space filling Curves (SFC)
 - High spatial locality
 - Hilbert ordering
 - Morton Ordering



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SFCs for Partitioning Data

level=5, #partitions=4

level=5, # partitions=3



2:1 Balance constraint

- Simplifies mesh & neighborhood
- Does not sacrifice adaptivity
- Minimizes the need to interpolate data
- Minimizes data-dependencies



Computational Methods

- Relativistic Fluids
 - Finite difference HRSC Method
 - HLLE flux
 - MP5 reconstruction
- Einstein Equations
 - BSSN formulation
 - 4th order finite differences
 - Kreiss-Oliger dissipation

Computational Methods

- Relativistic Fluids
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- Einstein Equations
 - BSSN formulation
 - 4th order finite differences
 - Kreiss-Oliger dissipation

These are just the conventional numerical methods

Finite differences & unstructured grids

- We need a regular grid to apply FD stencils
- This is not available everywhere for octree-refined grids







BSSN equations applied at a block level.

Automatic Code Generation

BSSN Equations

Dendro Code

from dendro import *

 $\partial_t \alpha = \mathcal{L}_\beta \alpha - 2\alpha K$ a_rhs = l1*dendro.lie(b, a) - 2*a*K

 $\partial_t \beta^i = \lambda_2 \beta^j \partial_j \beta^i + \frac{3}{4} f(\alpha) B^i \qquad b_{rhs} = [3/4*f(a)*B[i] + 12*dendro.vec_j_ad_j(b, b[i]) for i in dendro.e_i]$

Automatic Code Generation

```
# declare variables
```

```
a = dendro.scalar("alpha", "[pp]")
Gt = dendro.vec3("Gt", "[pp]")
gt = dendro.sym_3x3("gt", "[pp]")
```

```
dendro.set_metric(gt)
igt = dendro.get_inverse_metric()
```

```
a_rhs = l1*dendro.lie(b, a) - 2*a*K
```

••

outs = [a_rhs, b_rhs, gt_rhs, chi_rhs, At_rhs, K_rhs, Gt_rhs, B_rhs]
vnames = ['a_rhs', 'b_rhs', 'gt_rhs', 'chi_rhs', 'At_rhs', 'K_rhs', 'Gt_rhs', 'B_rhs']
dendro.generate(outs, vnames, '[pp]')

Automatic Code Generation

```
// Dendro: {{{
    // Dendro: original ops: 678611double DENDRO_0 = 2*alpha;
    double DENDRO_1 = 0.75*alpha*lambda_f[1] + 0.75*lambda_f[0];
    double DENDRO_2 = grad(0, beta0);
    double DENDRO_3 = grad(1, beta1);
```

```
B_rhs0 = -B0*eta - DENDRO_952*lambda[3] + DENDRO_993 + lambda[2]*(beta0*agrad(0, B0) + beta1*agrad(1, B0) + beta2*agrad(2, B0));
B_rhs1 = -B1*eta + DENDRO_1003 - DENDRO_994*lambda[3] + lambda[2]*(beta0*agrad(0, B1) + beta1*agrad(1, B1) + beta2*agrad(2, B1));
B_rhs2 = -B2*eta - DENDRO_1004*lambda[3] + DENDRO_1006 + lambda[2]*(beta0*agrad(0, B2) + beta1*agrad(1, B2) + beta2*agrad(2, B2));
// Dendro: reduced ops: 4602
// Dendro: }}
```

Automatic Code Generation - vectorization

// Dendro vectorized code: {{{
 double v0 = 2.0;
 double v1 = alpha[pp];
 double v2 = dmul(v1, v0);

•

•

v14 = B2[pp]; v15 = eta; v16 = dmul(v15, v14); v17 = dmul(v16, negone); v18 = DENDRO_989; v19 = lambda[3]; v20 = dmul(v19, v18); v21 = dmul(v20, negone); v22 = dadd(v21, v17); v23 = dadd(v22, v13); v24 = dadd(v23, v0); B_rhs2[pp] = v24; // Dendro vectorized code: }}}

Automatic Code Generation - CUDA

//input vars begin

double * K = __sm_base + 0; double * gt1 = sm base + 27; double * beta1 = sm base + 54; double * gt3 = sm base + 81; double * At1 = sm base + 108; double * gt5 = sm base + 135; double * alpha = sm base + 162; double * gt4 = sm base + 189; double * gt2 = sm base + 216; double * beta2 = sm_base + 243; double * At3 = sm base + 270; double * At4 = sm base + 297; double * At0 = sm base + 324; double * At2 = sm base + 351; double * beta0 = sm base + 378; double * gt0 = sm base + 405; double * chi = sm base + 432; double * At5 = sm base + 459;

// deriv vars begin

double * grad2 0 0 gt3 = sm base + 486; double * grad2 2 2 alpha = sm base + 513; double * grad2 1 2 gt1 = sm base + 540; double * grad 2 gt3 = sm base + 567; // load data from global to shared memory cuda:: loadGlobalToShared3D<double>(& unzipInVar[cuda::VAR::U_K][of fset],(double *) K,(const unsigned int *) ijk_lm,(const unsigned int *) alignedSz,(const unsigned int *) tile_sz);

if(!((threadIdx.x>(ijk_lm[1]-ijk_lm[0])) || (threadIdx.y>(ijk_lm[3]-ijk lm[2])))){

```
double x,y,z,r coord,eta;
unsigned int pp =
0*tile_sz[0]*tile_sz[1]+threadIdx.y*tile_sz[1]+threadIdx.x;
for(unsigned int k=0;k<=(ijk lm[5]-
ijk_lm[4]);++k,pp+=tile_sz[0]*tile_sz[1]){</pre>
```

But modern clusters are heterogeneous!

Heterogeneous Architectures



- GPUs are very fast, but require SIMD (Single Instruction Multiple Data)
- CPUs handle inter-processor communication and boundary zones
- GPUs work on interior
- Computation and communication are interleaved

Experiments

Nonlinear Sigma Model

Connection to BH critical phenomena, Liebling (2004)

$$\partial_t^2 \phi - \nabla^2 \phi = -\frac{\sin 2\phi}{r^2}$$

r = symbols('r')

chi_rhs = phi



Binary Black Holes

0.75 0.50 0.25 0.00

















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Need for true Adaptivity



Performance & Scalability

- Evaluated at Stampede2 & Comet (XSEDE)
- Large-scale scalability on Titan at ORNL
 - Cray XK7 with 18,688 nodes with Nvidia K80s.

ET Comparison

DENDRO-GR dofs (along the diagonal) \uparrow per core 250K35.46 - 12.34 6.88 4.01 2.06 1.06

total

dofs



EINSTEIN TOOLKIT(strong scaling)
 DENDRO-GR(strong scaling)
 EINSTEIN TOOLKIT(weak scaling)

DENDRO-GR(weak scaling)

1536

3072

-0-





Weak Scaling



Scaling test performed on Titan with 18 levels of refinement.

Strong Scaling

54

GPU Performance

"Early Access" GPUs for Coral Sierra (LLNL)

GPUs on Comet

Open Source

- Dendro-GR is available on Github.
 - <u>https://github.com/paralab/Dendro-GR</u>
 - git clone git@github.com:paralab/Dendro-GR.git
- Dendro-GR builds with CMake. Requires MPI and GSL. CUDA optional for GPU support.
- Public version
 - Wave Equation
 - Maxwell Equations* (Baumgarte's BSSN-like formulation)
 - BSSN Equations
- Support FEM in addition to FD
 - DG support coming soon
 - Extensively used for CFD
- For more details <u>Fernando+ 1807.06128</u>

Summary

- Dendro: Octree + Wavelet Adaptive Multiresolution (WAMR)
- Scaling to 10⁵ cores *with refinement*
- Conventional finite difference/finite volume numerical methods
- Applications: Relativistic fluids and the BSSN equations
- Currently testing binary black hole simulations
- Future work
 - IMRIs
 - Neutron stars