

JARNIK'S CONVEX LATTICE

POLYGON FOR NON-SYMMETRIC NORMS

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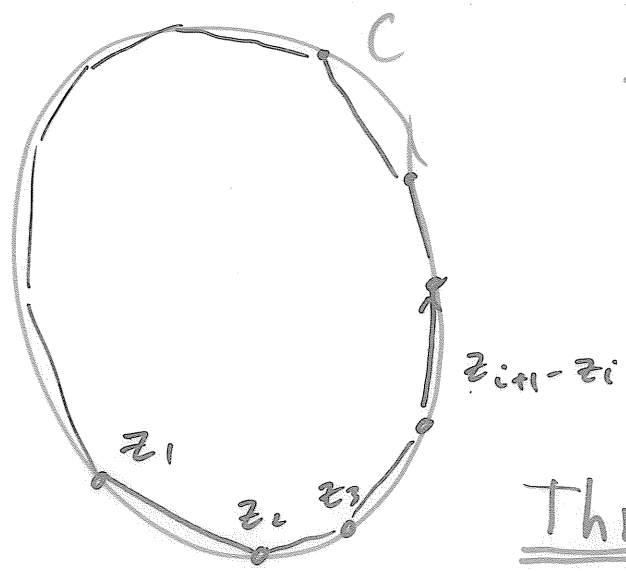
Jarnik (1926) C is a strictly convex (closed) curve in \mathbb{R}^2 , of length l . Then

$$|C \cap \mathbb{Z}^2| \leq \frac{3}{\sqrt[3]{2\pi}} l^{\frac{2}{3}} (1 + o(1)).$$

Moreover, this is best possible.

Here $\text{conv}(C \cap \mathbb{Z}^2)$ is a convex lattice n -gon with $n = |C \cap \mathbb{Z}^2|$

$$\mathcal{P}_n = \{ \text{convex lattice } n\text{-gons} \}$$



Def

$$L_n = \min \{ \text{per } P : P \in \mathcal{P}_n \}$$

Thm (Jarnik)

$$L_n = \frac{\sqrt{6\pi}}{9} n^{\frac{3}{2}} (1 + o(1))$$

Proof $z_{i+1} - z_i \in \mathcal{P}$ (primitive vectors) $\forall i$
 for the minimizer $P \in \mathcal{P}_n$.

$$E(P) = \{ z_{i+1} - z_i : i=1 \dots n \} \text{ distinct vectors } \in \mathcal{P}$$

$$L_n = \sum_1^n |z_{i+1} - z_i| \cong \boxed{\text{total length of the } n \text{ shortest distinct primitive vectors}}$$

density principle

can be computed

Take the smallest $\lambda > 0$ with

$$|\lambda B \cap P| \geq n$$

↑
unit ball (Eucl.)



⇓
 $\frac{6}{\pi^2} \lambda^2 \pi \approx n$

$$\sum_{p \in \lambda B \cap P} |p| \approx \frac{6}{\pi^2} \int_{\lambda B} |x| dx = \frac{6}{\pi^2} \cdot \frac{2}{3} \cdot \lambda^3 \pi$$

$$\Rightarrow L_n \geq \frac{\sqrt{6\pi}}{9} n^{3/2} (1 + o(1)).$$

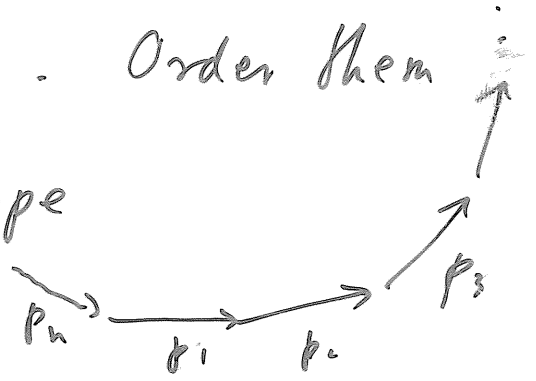
Construction for $L_n \leq$ same:

n even, take pairs $p, -p \in \lambda B \cap P$

then $\sum p = 0$. Order them

by increasing slope

n odd



$$L_n < L_{n+1}$$

increasing slope construction

FACT Assume $P_n \in \mathcal{P}_n$ is a minimizer.

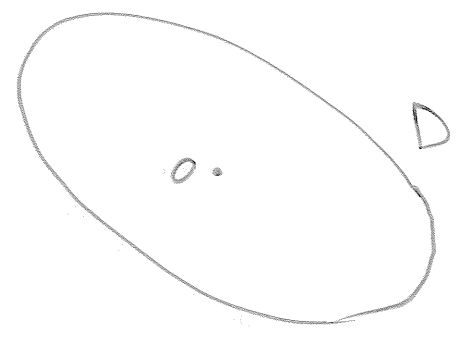
then $n^{-\frac{3}{2}} P_n$ tends to a circle

P_n has a limit shape (after scaling)

Assume D is the unit ball of another norm in \mathbb{R}^2 (symmetric)

for $P \in \mathcal{P}_n$

$$\text{per}_D P = \sum_1^n \|z_{i+1} - z_i\|_D$$



$$L_n(D) = L_n = \min \{ \text{per } P_D : P \in \mathcal{P}_n \}$$

Thm (M. PRODROMOU 2006)

$\lim n^{-\frac{3}{2}} L_n(D)$ exists and

equals
$$\frac{\pi \int_D |x| dx}{(6 \text{ Area } D)^{\frac{1}{3}}}$$

Thm (same)

the minimizers have a limit shape

Note! (after scaling)

$K \subset \mathbb{R}^2$ convex, compact. $0 \in \text{int} K$

$$d = \max \{ |x| : x \in K \}$$

$w = \text{width of } K$. ($w \geq 3 \text{ say}$)

$$\left| |K \cap \mathbb{Z}^2| - \text{Area } K \right| \ll \frac{1}{w} \text{Area } K$$

$$\left| |K \cap \mathbb{P}| - \frac{6}{\pi^2} \text{Area } K \right| \ll \frac{\log d}{w} \text{Area } K$$

$f: \mathbb{R}^2 \rightarrow \mathbb{R}$ 1-homogeneous $\left(\begin{array}{l} f(\lambda x) = \lambda f(x) \\ \forall \lambda > 0 \forall x \end{array} \right)$
then

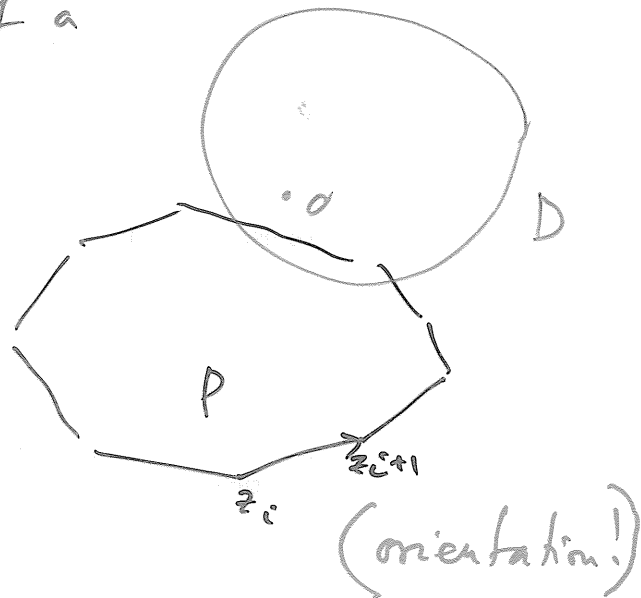
$$\left| \sum_{z \in K \cap \mathbb{P}} f(z) - \frac{6}{\pi^2} \int_K f(x) dx \right| \ll \frac{\log d}{w} \cdot \int_K |f(x)| dx$$

standard

Let D be the unit ball of a non-symmetric norm in \mathbb{R}^2

$$\text{per}_D P = \sum_1^n \|z_{i+1} - z_i\|_D$$

$$L_n(D) = \min \{ \text{per } P : P \in \mathcal{P}_n \}$$



Result (I. B., N. ENRIQUEZ, 2009)

$$L_n(D) = c(D) n^{\frac{3}{2}} (1 + o(1))$$

there is a limit shape :

$$n^{-\frac{3}{2}} P_n \rightarrow C \quad \text{a fixed convex body in } \mathbb{R}^2$$

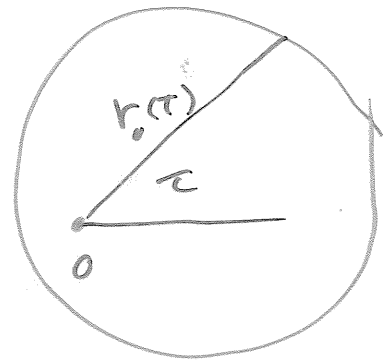
Assume D is strictly convex (convenience)

$$\Rightarrow L_n < L_{n+1} \quad \text{and}$$

$$\Rightarrow \text{every edge of a minimizer } \in \mathcal{P}$$

VARIATIONAL PROBLEM $VP(r_0)$

$r_0 : [0, 2\pi] \rightarrow \mathbb{R}^+$ is
the radial function
of D



find $r : [0, 2\pi] \rightarrow \mathbb{R}^+$, cont. $r(0) = r(2\pi)$ with

$$\int_0^{2\pi} \frac{r^3(\tau)}{r_0(\tau)} d\tau \rightarrow \min$$

subject
to

$$\int_0^{2\pi} r^3(\tau) \cos \tau d\tau = 0$$

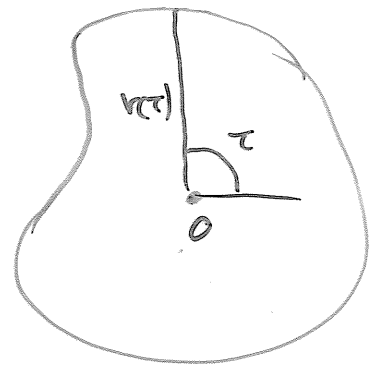
$$\int_0^{2\pi} r^3(\tau) \sin \tau d\tau = 0$$

$$\frac{1}{2} \int_0^{2\pi} r^2(\tau) d\tau = 1$$

} centre of
gravity of
 K is 0

} Area $K = 1$

r is the radial
function of a convex
(starshaped) body K



KNOWN all optimal sol's to $VP(r_0)$ satisfy

$$\frac{1}{r(\tau)} = \frac{a}{r_0(\tau)} + b \cos \tau + c \sin \tau$$

NOT KNOWN existence of opt. sol's to $VP(r_0)$
(later)

FACT If $r(\tau)$ is a feasible sol'n to $VP(r_0)$,
then there is a sequence $Q_n \in P_n$
with

$$\lim_{n \rightarrow \infty} n^{-\frac{3}{2}} \text{per } Q_n = \frac{2}{3\sqrt{6}\pi} \int_0^{2\pi} \frac{r(\tau)^3}{r_0(\tau)} d\tau$$

Proof $r(\tau)$ is the radial function of K

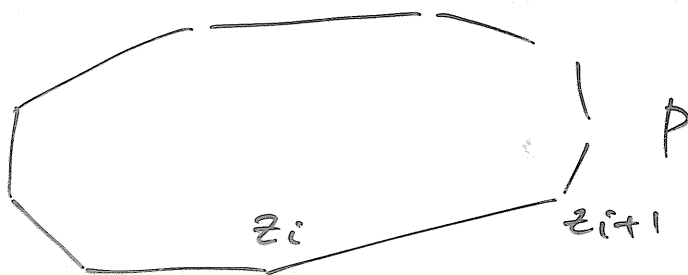
Area $K=1$, center of gravity of $K=0$

choose $\lambda > 0$ minimal with $(\lambda K \cap P) = \emptyset$

.....

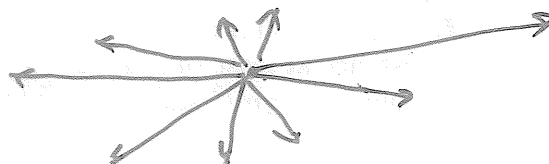


$$P \in \mathcal{P}_n$$



$$E(P) = \{ z_{i+1} - z_i : i = 1, \dots, n \}$$

$$C(P) = \text{conv } E(P)$$



Suppose $P \in \mathcal{P}_n$ is a minimizer.

FACT $E(P) \subset P$

FACT $L_n < L_{n+1}$

HOPE $E(P)$ contains all (almost all) primitive vectors from $C(P)$

and $C(P)$ is nice

OF COURSE $c_1 |x| \leq \|x\|_D \leq c_2 |x|$

OR $|x| \ll \|x\|_D \ll |x|$

OR $c_1 B \subset D \subset c_2 B$

Claim 1. $L_n \gg n^{3/2}$

Proof (simple) The n shortest (in D -norm) distinct vectors in P are p_1, \dots, p_n .

Then for $P \in \mathcal{P}_n$

$$\begin{aligned} \text{per } P &\cong \sum_1^n \|p_i\| \approx \frac{6}{\pi^2} \int_{\lambda D} \|x\| dx \\ &= \frac{6}{\pi^2} \lambda^3 \int_D \|x\| dx \end{aligned}$$

where $\lambda > 0$ is minimal with $\{p_1, \dots, p_n\} \subset \lambda D$.

$$\begin{aligned} &\updownarrow \\ n &\approx \frac{6}{\pi^2} \lambda^2 \text{Area } D \end{aligned}$$

$$\Rightarrow L_n \gg n^{3/2} \quad \square$$

Claim 2. $L_n \ll n^{3/2}$

Proof $p_0 = -\sum_1^n p_n \quad \|p_0\| \ll n^{3/2}$

increasing slope construction on p_0, p_1, \dots, p_n

gives a convex lattice n or $(n+1)$ -gon with perimeter $\ll n^{3/2} \quad \square$

Corollary $\liminf n^{-\frac{3}{2}} L_n(\mathbb{D}) = \alpha > 0$

Now $P_n \in \mathcal{P}_n$ is a minimizer.

$$E_n = E(P_n), \quad C_n = C(P_n)$$

$w = w(E_n) =$ width of E_n

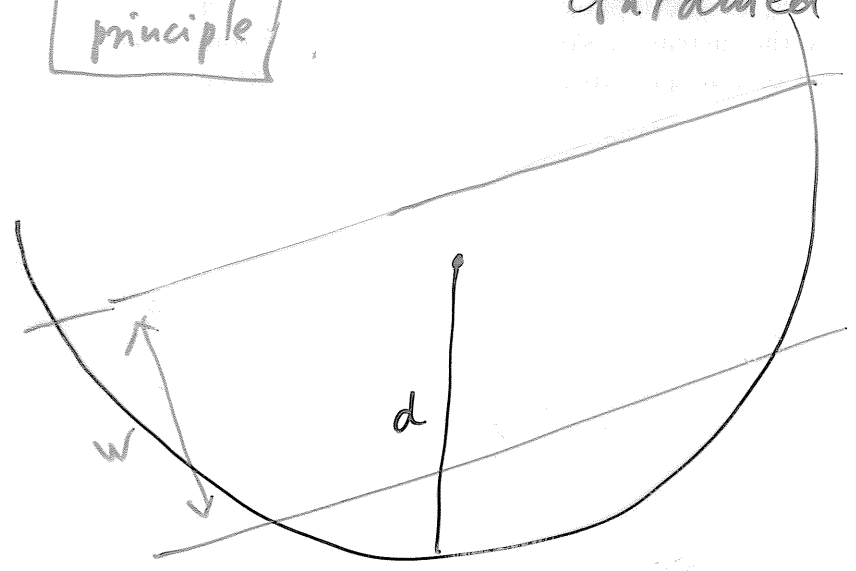
Claim 3. $w \gg n^{\frac{1}{2}}$

Proof (sketch) Set $w = \gamma n^{\frac{1}{2}}$. NEED $\gamma \gg 1$.

$$L_n = \sum_{v \in E_n} \|v\| \gg \sum_{v \in E_n} |v| \geq M_n(w)$$

where $M_n(w) =$ min num of the n shortest (Euclidean) distinct vectors contained in a strip of width w

density principle



... turns out to be $\gg \frac{1}{\gamma} n^{\frac{3}{2}}$. \square

Claim 4 $R > 0$ minimal with $E_n \subset RB$ (3)

$$\Rightarrow R \ll n^{\frac{1}{2}}$$

□

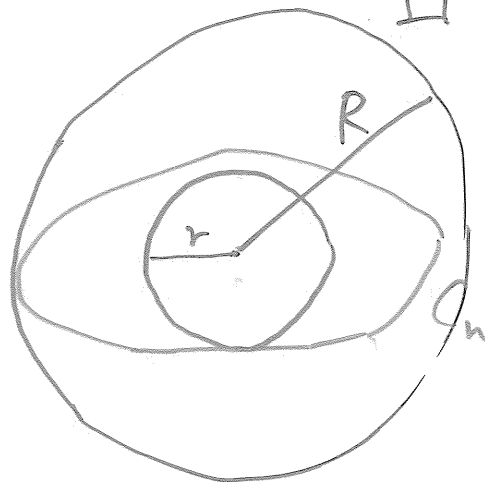
Claim 5 $r > 0$ maximal with $rB \subset C_n$

$$\Rightarrow r \gg n^{\frac{1}{2}}$$

□

Cor $\exists 0 < r < R$
such that

$$rB \subset \frac{1}{\sqrt{n}} C_n \subset RB$$



LEMMA $\forall \varepsilon > 0 \exists n_0 = n_0(\varepsilon, D)$ such that
for all $n > n_0$

$$P \cap (1 - \varepsilon) C_n \subset E_n$$

Proof is based on a special approximation
lemma (omitted here)

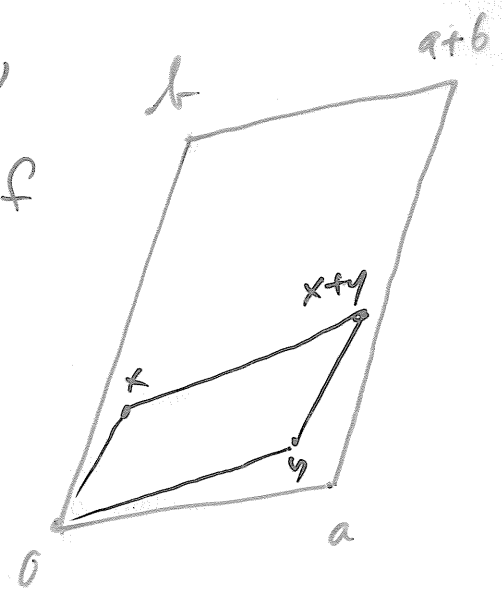
and on

Claim 6. $a, b \in E_n, a \neq \pm b,$

$T = \text{conv} \{0, a, b, a+b\}$. If

$x, y \in (P \cap T) \setminus E_n$ and $x \neq y$

then $x+y \notin T$.



Cor $E_n \approx P \cap C_n$

Now one can choose a subsequence $n_1 < n_2 < \dots$

with

$$\lim_{k \rightarrow \infty} n_k^{-\frac{3}{2}} L_{n_k} = \alpha$$

and

$$\lim_{k \rightarrow \infty} \frac{1}{\sqrt{\text{Area } C_{n_k}}} C_{n_k} = C$$

(using the Blaschke selection thm)

where C is a convex body with

$$rB \subset C \subset RB.$$

Claim 7. The radial function of C , $v(\tau)$,
is a feasible sol'n to $VP(r_0)$

Proof Area $C = 1$ and the center
of gravity is O . \square

Thm $v(\tau)$ is an optimal sol'n to $VP(r_0)$

Proof $v(\tau)$ is feasible.

If v^* is also feasible and

$$\int_0^{2\pi} \frac{v^*(\tau)^3}{v_0(\tau)} d\tau < \int_0^{2\pi} \frac{v(\tau)^3}{v_0(\tau)} d\tau,$$

then (from FACT) one gets a sequence

$Q_n \in \mathcal{P}_n$ with

$$\lim n^{-\frac{3}{2}} \text{per } Q_n < \alpha$$

\square

Claim 8 If the centre of gravity of D is 0 , then the opt. sol'n to $VP(r_0)$ is unique.

Proof If r is a feasible sol'n, then by Hölder

$$\int_0^{2\pi} r^2(\tau) d\tau = 2 \left(\int_0^{2\pi} \frac{r^3(\tau)}{r_0(\tau)} d\tau \right)^{\frac{2}{3}} \left(\int_0^{2\pi} r_0^2(\tau) d\tau \right)^{\frac{1}{3}}$$

$= 2$
 $= 2 \text{Area } D$

with equality iff $r = \gamma r_0$ ($\gamma > 0$). \square

Then the opt. sol'n to $VP(r_0)$ is unique

Proof Assume $r: [0, 2\pi] \rightarrow \mathbb{R}^+$ is an opt sol'n to $VP(r_0)$. Then

$$\frac{1}{r(\tau)} = \frac{a}{r_0(\tau)} + b \cos \tau + c \sin \tau$$

with suitable $a, b, c \in \mathbb{R}$.

(13)

Claim 9. Here $a \int_0^{2\pi} \frac{r^3}{r_0} = 2$.

\Rightarrow the coefficient a is the same for all optimal solutions.

Proof

$$\begin{aligned} 2 &= \int_0^{2\pi} r^2 = \int_0^{2\pi} r^3 \frac{1}{r} = b \int_0^{2\pi} r^3 \cos \tau - c \int_0^{2\pi} r^3 \sin \tau \\ &= \int_0^{2\pi} r^3 \left[\frac{1}{r} - b \cos \tau - c \sin \tau \right] = \\ &= a \int_0^{2\pi} \frac{r^3}{r_0} d\tau \quad \square \end{aligned}$$

Suppose now $r_i: [0, 2\pi] \rightarrow \mathbb{R}^+$ ($i=1,2$)

are opt sol's to $VP(r_0)$. Then

$$\frac{1}{r_i} = \frac{a}{r_0} + b_i \cos \tau + c_i \sin \tau$$

$$\int_0^{2\pi} \frac{r_i^3}{r_i} = \int_0^{2\pi} r_i^3 \left(\frac{a}{r_0} + b_i \cos \tau + c_i \sin \tau \right) = a \int_0^{2\pi} \frac{r_i^3}{r_0}$$

r_1 is the radial function of a starshaped body whose centre of gravity is O .

Claim 8 $\Rightarrow r_1 = r_2$. \square