

A hierarchy of models, from planetary to mesoscale, within a single switchable numerical framework

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Outline of the talk

Met Office

- Overview of Met Office's Unified Model
- Spherical Geopotential Approximation
- Relaxing that approximation
- A method of implementation
- Alternative view of Shallow-Atmosphere Approximation
- Application to departure points



Met Office's Unified Model

Unified Model (UM) in that single model for:

Operational forecasts at

➢ Mesoscale (resolution approx. 12km → 4km → 1km)

≻Global scale (resolution approx. 25km)

- Global and regional climate predictions (resolution approx. 100km, run for 10-100 years)
- + Research mode (1km 10m) and single column model
- 20 years old this year

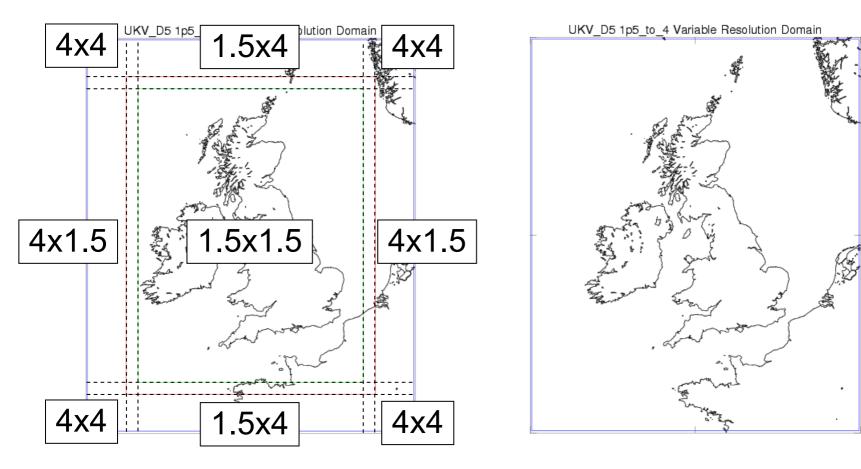


Timescales & applications

Forecast Uncertainty Past data Years Decades Weather risk Risk Months Seasons management Regulatory standards Policy and Scenario Balancing Maintenance regulation Today Week schedules planning resources New Asset Nowcast Disruption Maintenance infrastructure Event Resource management and real-time verification planning planning planning Resources Resources Mitigation/ Design Demand Balancing Market planning Financial advertising standards resources trading forecasts and property portfolio risk Investment Forecast Weather Stand-by Operations Hurricane strategies planning calibration preparations warnings forecasts management Minutes, 1–2 Weeks Hours, Days Months Seasons Years Decades

Forecast lead-time



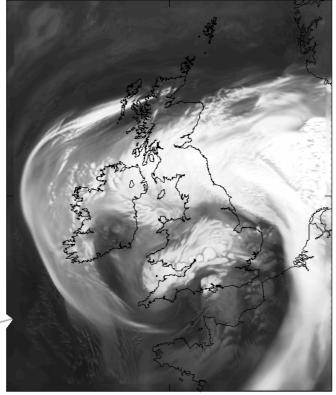


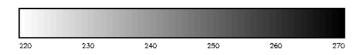


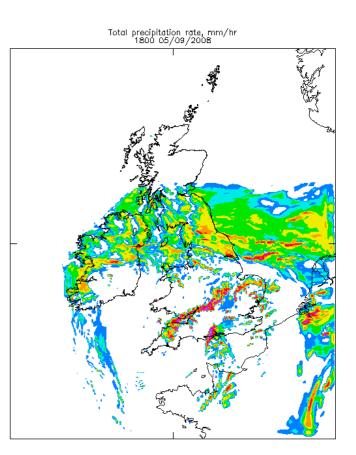
The 'Morpeth Flood', 06/09/2008

1.5 km L70 Prototype UKV From 15 UTC 05/09 12 km









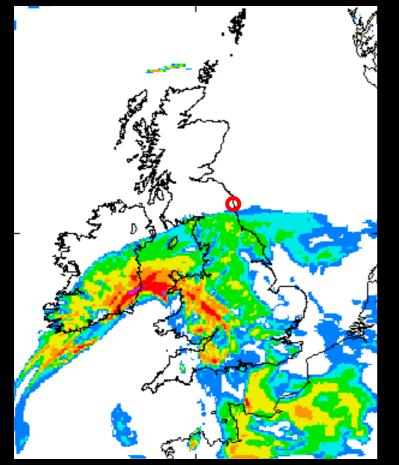


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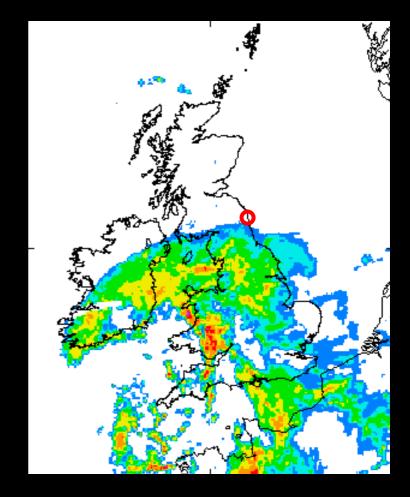
0600 UTC



2009: Clark et al, Morpeth Flood



UKV 4-20hr forecast 1.5km gridlength convection permitting model

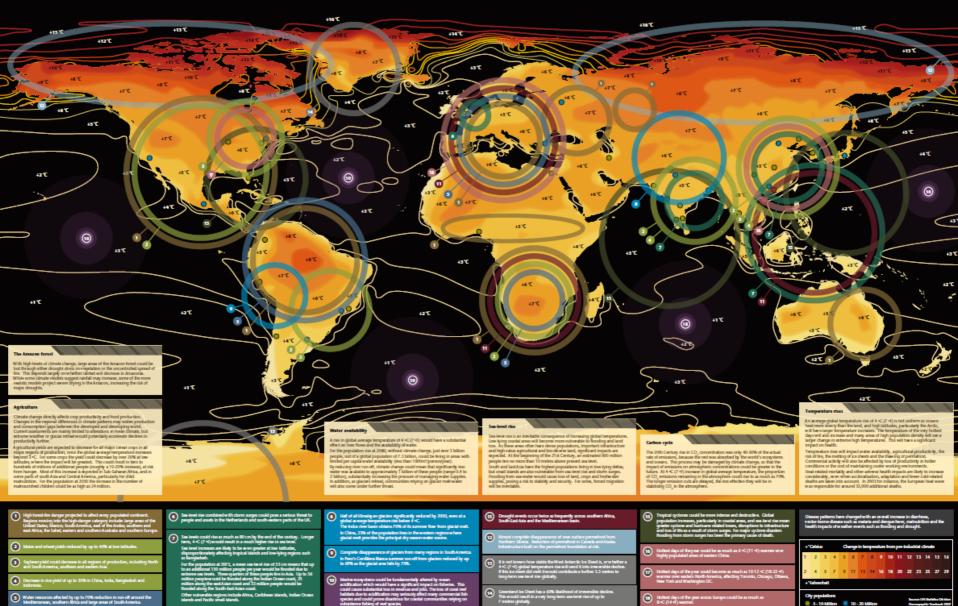


UK radar network

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The impact of a global temperature rise of 4 °C (7 °F)

HMGovernment





The underlying equations...

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Traditional Spherically Based Equations

 $\frac{D_r u}{Dt} - \frac{uv \tan \phi}{r} - (2\Omega \sin \phi) v + \frac{c_{pd}\theta_v}{r \cos \phi} \frac{\partial \Pi}{\partial \lambda} = \boxed{-\frac{uw}{r} - (2\Omega \cos \phi) w} + S^u$ $\frac{D_r v}{Dt} + \frac{u^2 \tan \phi}{r} + (2\Omega \sin \phi) u + \frac{c_{pd}\theta_v}{r} \frac{\partial \Pi}{\partial \phi} = \boxed{-\frac{vw}{r}} + S^v$ $\boxed{\frac{D_r w}{Dt}} + c_{pd}\theta_v \frac{\partial \Pi}{\partial r} + \frac{\partial \Phi}{\partial r} = \boxed{\frac{u^2 + v^2}{r} + (2\Omega \cos \phi) u + S^w}$

Red = Shallow-Atmosphere Approx (eliminate boxed terms) **Blue** = Hydrostatic Approx (eliminate boxed term) **Green** = Spherical Geopotential Approx ($\partial \Phi / \partial r = g(r)$ only)



The Spherical Geopotential Approximation

- All geopotentials including the Earth's surface are represented by concentric *spheres*
- Apparent gravity acts towards the centre of the Earth
- To prevent spurious vorticity sources/sinks g cannot vary with latitude (White et al., 2005)
- Nearly all numerical models of the global atmosphere are based on the SGA



- Good to get Earth's shape right(ish) in models!
- Good to include the small (\approx 0.5%) increase of g from Equator to Poles without spurious vorticity sources/sinks
- Most likely to be evident in long term climate simulations
- Good to quantitatively test the accuracy of the spherical geopotential approximation
- Important though that signal is not influenced by different numerics



The Shape of the Earth

- The Earth is approximately an oblate spheroid:
 - Equatorial radius

$$a = 6378 \,\mathrm{km}$$

– Polar radius

 $c=6357\,\mathrm{km}$

• Ellipticity small

$$\varepsilon \equiv \frac{a-c}{a} \approx \frac{1}{298}$$

• But a - c = 21 km, more than twice the height of Everest



Geopotential near rotating fluid spheroid

- Classical problem (Clairaut in 18th century) to account for the observed increase of g with latitude
- Re-examined by White et al. (2008) to find geopotentials
- Two small parameters:

$$\varepsilon \equiv \frac{a-c}{a} \approx \frac{1}{298}$$
$$m \equiv \frac{\mathbf{\Omega}^2 a^3}{\gamma M_E} \sim \left| \frac{\text{centrifugal force}}{\text{Newtonian gravity}} \right| \approx \frac{1}{289},$$

 \bullet To $O\left(\varepsilon,m\right)$ geopotentials are spheroids

$$x^{2} + z^{2} \left[1 + (2\epsilon - m) \frac{a^{2}}{R^{2}} + m \frac{R^{3}}{a^{3}} \right] = R^{2}$$



Variation of Gravity

• Near to the surface R = a

$$\frac{g_P}{g_{Eq}} = 1 + \frac{5}{2}m - \epsilon$$

• Earth's actual mass distribution $m = \epsilon$ (to within 3%):

$$\frac{g_P}{g_{Eq}} = 1 + \frac{3}{2}\epsilon \implies 0.5\% \text{ increase}$$

• For uniform mass distribution $m = 4\epsilon/5$ (Newton)

$$\frac{g_P}{g_{Eq}} = 1 + \epsilon$$

captures 2/3rds of the variation



Spheroidal Coordinates...

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Confocal Oblate Spheroids

• "Oblate spheroidal coordinates" of e.g. Gill (1982) and Gates (2004) are in fact *confocal* oblate spheroidal coordinates:

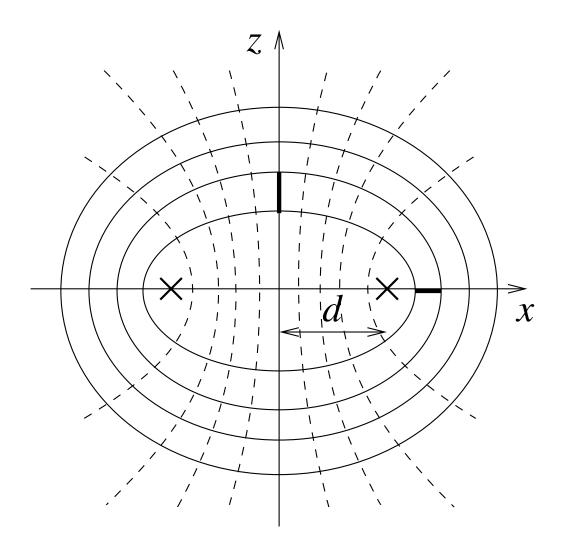
$$\frac{x^2}{\cosh^2 \eta} + \frac{z^2}{\sinh^2 \eta} = d^2$$

with d fixed, η variable

 \bullet Match to Earth's surface \Rightarrow

$$\frac{g_P}{g_{Eq}} = \tanh \eta_0 \approx 1 - \epsilon \neq 1 + \frac{3}{2}\epsilon!$$

• Separation along minor axis greater than along major axis



CONFOCAL ELLIPSES and HYPERBOLAS



Similar ellipses

Equation is:

$$x^2 + (1+\mu) \, z^2 = \mathbf{r^2}$$

Now μ fixed, **r** variable

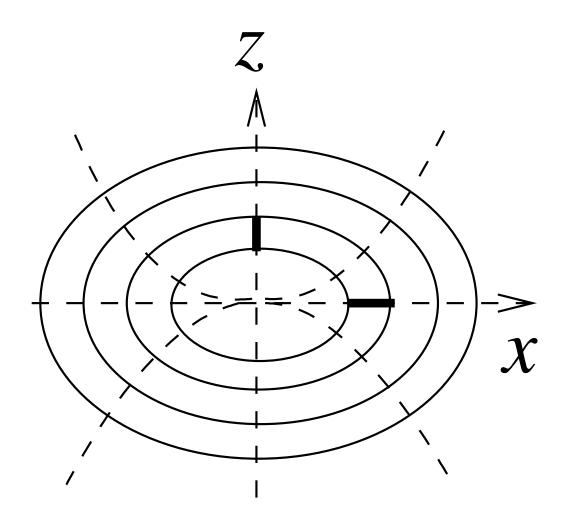
Match to Earth's surface \Rightarrow

$$\mu = 2\epsilon$$

and

$$\frac{g_P}{g_{Eq}} = (1+\mu)^{1/2} \approx 1+\epsilon$$

- Sign right!
- Captures 2/3rds actual variation for actual Earth
- Exact for an Earth with uniform mass distribution



SIMILAR ELLIPSES



Implementation...

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Semi-Lagrangian scalar advection

Consider

$$\frac{DF}{Dt} = R$$

Integrate *along trajectory*:

$$\int_{t}^{t+\Delta t} \frac{DF}{Dt} dt = \frac{F^{t+\Delta t}\left(x\right) - F^{t}\left(x - U\Delta t\right)}{\Delta t} = \int_{t}^{t+\Delta t} Rdt \approx \frac{R^{t+\Delta t}\left(x\right) + R^{t}\left(x - U\Delta t\right)}{2}$$

i.e.

$$\left(F - \frac{\Delta t}{2}R\right)_{A}^{t+\Delta t} = \left(F + \frac{\Delta t}{2}R\right)_{D}^{t}$$



Semi-Lagrangian vector advection

For the vector equation

$$\frac{D\mathbf{v}}{Dt} = \mathbf{S}$$

Integrate *along trajectory* to obtain, as before:

$$\left(\mathbf{v} - \frac{\Delta t}{2}\mathbf{S}\right)_{A}^{t+\Delta t} = \left(\mathbf{v} + \frac{\Delta t}{2}\mathbf{S}\right)_{D}^{t}$$

All well and good...

But what of components?



The Problem of Curvilinear Coordinates...

The familiar problem that

$$\mathbf{i} \cdot \frac{D\mathbf{v}}{Dt} \equiv \mathbf{i} \cdot \frac{D}{Dt} \left(u\mathbf{i} + v\mathbf{j} + w\mathbf{k} \right) \neq \frac{D\left(\mathbf{i} \cdot \mathbf{v}\right)}{Dt} \quad \left[= \frac{Du}{Dt} \right]$$

translates into the SL equivalent that

$$\mathbf{i}_A.(\mathbf{v}_D) \neq (\mathbf{i}_A.\mathbf{v}_A)_D \ [= u_D]$$

In fact

$$\mathbf{v}_D = u_\mathcal{D} \mathbf{i}_D + v_\mathcal{D} \mathbf{j}_D + w_\mathcal{D} \mathbf{k}_D$$

so that, e.g.,

$$\mathbf{i}_{A} \cdot (\mathbf{v}_{D}) = u_{\mathcal{D}} \left(\mathbf{i}_{A} \cdot \mathbf{i}_{D} \right) + v_{\mathcal{D}} \left(\mathbf{i}_{A} \cdot \mathbf{j}_{D} \right) + w_{\mathcal{D}} \left(\mathbf{i}_{A} \cdot \mathbf{k}_{D} \right)$$



• Extending this to all three directions:

$$\begin{pmatrix} u \\ v \\ w \end{pmatrix}_{D} = \mathbf{M} \begin{pmatrix} u_{\mathcal{D}} \\ v_{\mathcal{D}} \\ w_{\mathcal{D}} \end{pmatrix}$$

where

$$\mathbf{M} \equiv \begin{pmatrix} \mathbf{i}_A \cdot \mathbf{i}_D & \mathbf{i}_A \cdot \mathbf{j}_D & \mathbf{i}_A \cdot \mathbf{k}_D \\ \mathbf{j}_A \cdot \mathbf{i}_D & \mathbf{j}_A \cdot \mathbf{j}_D & \mathbf{j}_A \cdot \mathbf{k}_D \\ \mathbf{k}_A \cdot \mathbf{i}_D & \mathbf{k}_A \cdot \mathbf{j}_D & \mathbf{k}_A \cdot \mathbf{k}_D \end{pmatrix}$$

$\bullet \ \mathbf{M} \ \text{transforms}$

- from: vector components in the *departure-point* frame
- to: vectors in the *arrival-point* frame



The Consequence...

Component form of momentum equations can be written as:

$$\left(\mathbf{v} - \frac{\Delta t}{2}\mathbf{S}\right)_{\mathcal{A}}^{t+\Delta t} = \mathbf{M}\left(\mathbf{v} + \frac{\Delta t}{2}\mathbf{S}\right)_{\mathcal{D}}^{t}$$

where

$$\mathbf{X}_{\mathcal{A}} \equiv \left(X_{\mathcal{A}}, Y_{\mathcal{A}}, Z_{\mathcal{A}} \right)^{T}$$
$$\mathbf{X}_{\mathcal{D}} \equiv \left(X_{\mathcal{D}}, Y_{\mathcal{D}}, Z_{\mathcal{D}} \right)^{T}$$

- No explicit metric terms
- No singularity at the pole



M for Spherical Polar Coordinates

 $M_{11} = \cos\Delta\lambda, \quad M_{12} = \sin\phi_D \sin\Delta\lambda, \quad M_{13} = -\cos\phi_D \sin\Delta\lambda$ $M_{21} = -\sin\phi_A \sin\Delta\lambda$ $M_{22} = \cos\phi_A \cos\phi_D + \sin\phi_A \sin\phi_D \cos\Delta\lambda$ $M_{23} = \cos\phi_A \sin\phi_D - \sin\phi_A \cos\phi_D \cos\Delta\lambda$ $M_{31} = \cos\phi_A \sin\Delta\lambda$ $M_{32} = \sin\phi_A \cos\phi_D - \cos\phi_A \sin\phi_D \cos\Delta\lambda$ $M_{33} = \sin\phi_A \sin\phi_D + \cos\phi_A \cos\phi_D \cos\Delta\lambda$



Where have the Metric Terms Gone?

- SL form holds for finite displacements
- Reduces to Eulerian form as displacement ($\Delta \lambda \equiv \lambda_A \lambda_D$) and $\Delta t \rightarrow 0$:
- In this limit:

$$\sin \Delta \lambda \to \Delta \lambda, \ \frac{\Delta \lambda}{\Delta t} \to \frac{u_{\mathcal{A}}}{r_A \cos \phi_A}, \ u_{\mathcal{D}} \to u_{\mathcal{A}}, \ v_{\mathcal{D}} \to v_{\mathcal{A}}$$

and so:

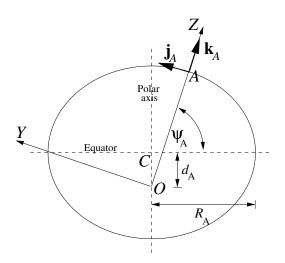
$$M_{12}v_{\mathcal{D}} = v_{\mathcal{D}}\sin\phi_D\sin\Delta\lambda \to \frac{u_{\mathcal{A}}v_{\mathcal{A}}\tan\phi_A}{r_A}\Delta t$$

 \bullet Each off-diagonal element of ${\bf M}$ generates one metric term



Similar Oblate Spheroidal Form

- Basic derivation is the same
- "Trick": use secondary meridional coordinate, ψ , geographic latitude (not orthogonal to vertical coordinate)



- Significantly M has *exactly* the same form but with $\phi \rightarrow \psi$
- Contrast Eulerian formulation where (for COS at least) two extra metric terms appear



Shallow-atmosphere...

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Shallow-Atmosphere Approximation

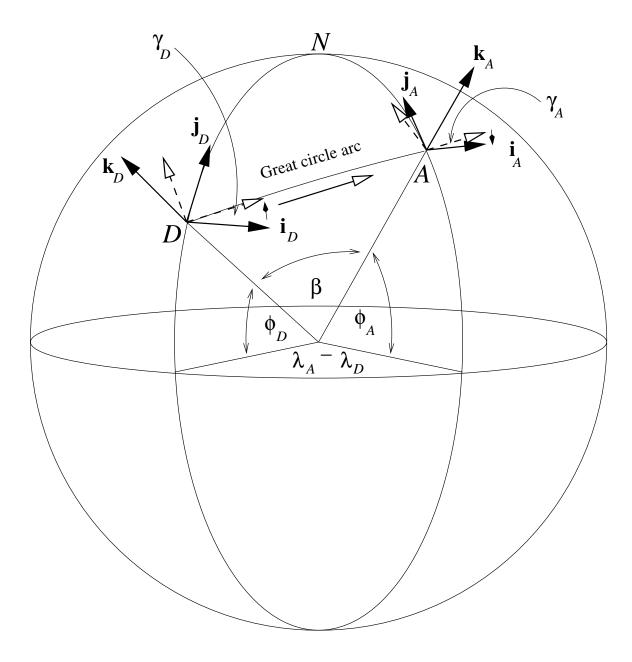
 $\frac{D_r u}{Dt} - \frac{uv \tan \phi}{r} - (2\Omega \sin \phi) v + \frac{c_{pd}\theta_v}{r \cos \phi} \frac{\partial \Pi}{\partial \lambda} = \left[-\frac{uw}{r} - (2\Omega \cos \phi) w \right] + S^u$ $\frac{D_r v}{Dt} - \frac{u^2 \tan \phi}{r} + (2\Omega \sin \phi) u + \frac{c_{pd}\theta_v}{r} \frac{\partial \Pi}{\partial \phi} = \left[-\frac{vw}{r} \right] + S^v$ $\left[\frac{D_r w}{Dt} + c_{pd}\theta_v \frac{\partial \Pi}{\partial r} + \frac{\partial \Phi}{\partial r} \right] = \left[\frac{u^2 + v^2}{r} + (2\Omega \cos \phi) u + S^w$

- Traditional approach:
 - Eliminate all boxed red terms
 - Replace r with a in all algebraic expressions



- ${\ensuremath{\,\bullet\,}} \mathbf{M}$ is a rotation matrix
- Can be factored in infinite number of ways
- One is of particular interest:
- **1.** Rotate departure-point UVT about radial direction to line up i with Great Circle connecting r_D with r_A
- 2. Rotate UVT along Great Circle arc about new \boldsymbol{j} direction

3. Rotate UVT about new radial direction to line up i direction



A PARTICULAR DECOMPOSITION OF ${\rm M}$



Shallow-atmosphere - a different perspective

\bullet Result is $\mathbf{M}=\mathbf{A}\mathbf{C}\mathbf{D}$

- \bullet Shallow-atmosphere approximation is obtained by $\mathbf{C} \rightarrow \mathbf{I}$
- $\mathbf{M} \rightarrow \mathbf{AD} \equiv \mathbf{Q}$:

$$\mathbf{Q} = \begin{pmatrix} p & q & 0 \\ -q & p & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

with

$$p = \frac{M_{11} + M_{22}}{1 + M_{33}}$$

and

$$q = \frac{M_{12} - M_{21}}{1 + M_{33}}$$

giving the same result as in Temperton et al. (2001)



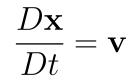
Departure points...

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Rotation Matrix for Departure Points?

- Departure points now take on greater significance
- And they are also governed by a vector equation



- So can we apply consistent approach (and avoid polar singularity issues)?
- Discrete vector form is

$$\mathbf{x}_D = \mathbf{x}_A - \frac{\Delta t}{2} \left(\mathbf{v}_A + \mathbf{v}_D \right)$$



Rotation Matrix for Departure Points II

As before component form can be written as

$$\mathbf{M}\mathbf{x}_{\mathcal{D}} = \mathbf{x}_{\mathcal{A}} - \frac{\Delta t}{2} \left(\mathbf{v}_{\mathcal{A}} + \mathbf{M}\mathbf{v}_{\mathcal{D}} \right)$$

Or, since
$$\mathbf{x}_{\mathcal{A}} = (0, 0, r_A)^T$$
 and $\mathbf{x}_{\mathcal{D}} = (0, 0, r_D)^T$:

$$M_{13}r_{\mathcal{D}} = -\frac{\Delta t}{2} \left[u_{\mathcal{A}} + (M_{11}u_{\mathcal{D}} + M_{12}v_{\mathcal{D}} + M_{13}w_{\mathcal{D}}) \right]$$
$$M_{23}r_{\mathcal{D}} = -\frac{\Delta t}{2} \left[v_{\mathcal{A}} + (M_{21}u_{\mathcal{D}} + M_{22}v_{\mathcal{D}} + M_{23}w_{\mathcal{D}}) \right]$$
$$M_{33}r_{\mathcal{D}} = r_{A} - \frac{\Delta t}{2} \left[w_{\mathcal{A}} + (M_{31}u_{\mathcal{D}} + M_{32}v_{\mathcal{D}} + M_{33}w_{\mathcal{D}}) \right]$$



A Local Cartesian Transform Approach

The result can be cast into a local Cartesian form:

$$X_{DA} = -\frac{\Delta t}{2} \left(U_{\mathcal{A}} + U_{DA} \right)$$
$$Y_{DA} = -\frac{\Delta t}{2} \left(V_{\mathcal{A}} + V_{DA} \right)$$
$$Z_{DA} = r_A - \frac{\Delta t}{2} \left(W_{\mathcal{A}} + W_{DA} \right)$$

where

$$\mathbf{X}_{DA} \equiv \mathbf{M}\mathbf{x}_{\mathcal{D}},$$

and

$$\mathbf{V}_{DA} \equiv \mathbf{M} \mathbf{v}_{\mathcal{D}}$$

are Departure point *coordinates* and *velocities* as seen in the Arrival-point Cartesian system



The Inverse Transformation

From (X_{DA}, Y_{DA}, Z_{DA}) obtain departure point as:

$$\tan\left(\lambda_A - \lambda_D\right) = \frac{-X_{DA}}{Z_{DA}\cos\phi_A - Y_{DA}\sin\phi_A}$$

$$r_D^2 = X_{DA}^2 + Y_{DA}^2 + Z_{DA}^2$$

$$\sin \phi_D = \frac{Y_{DA} \cos \phi_A + Z_{DA} \sin \phi_A}{r_D}$$

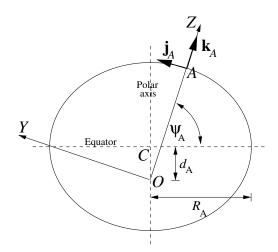


and

Spheroidal Case

Spheroidal case follows spherical case closely except:

$$\mathbf{x}_{\mathcal{A}} = (0, 0, r_A)^T - d_A (0, \cos \phi_A, \sin \phi_A)^T$$
$$\mathbf{x}_{\mathcal{D}} = (0, 0, r_D)^T - d_D (0, \cos \phi_D, \sin \phi_D)^T$$





Spheroidal Case

 \bullet Applying ${\bf M}$ as before, but to modified position vectors, results in

$$X_{DA} = M_{13}r_{\mathcal{D}}$$
$$Y_{DA} = M_{23}r_{\mathcal{D}} + (d_A - d_D)\cos\phi_A$$
$$X_{DA} = M_{33}r_{\mathcal{D}} + (d_A - d_D)\sin\phi_A$$

• Inversion is a little more complicated



Shallow-atmosphere case

Form obtained by constraining departure point to remain on arrival sphere:

$$\mathbf{x}_{D} = \mathbf{x}_{A} - \frac{\Delta t}{2} \left(\mathbf{v}_{A} + \mathbf{v}_{D} \right) + \mathcal{B} \left(\mathbf{x}_{A} + \mathbf{x}_{D} \right)$$

1. Cartesian transformation as before

$$\mathbf{X}_{DA} \equiv \mathbf{M}\mathbf{x}_{\mathcal{D}},$$

2. but velocities transform using momentum shallow-atmosphere rotation matrix

$$\mathbf{V}_{DA} \equiv \mathbf{Q} \mathbf{v}_{\mathcal{D}},$$

3. and

$$\mathbf{X}_{\mathcal{D}}^{H} = -\left(\frac{1+M_{33}}{2}\right)\frac{\Delta t}{2}\left(\mathbf{V}_{\mathcal{A}}^{H} + \mathbf{V}_{\mathcal{D}}^{H}\right)$$



What does it all mean?

Single set of equations:

$$\left[(u_i - \alpha \Delta t \Psi_i)^{n+1} = M_{ij} \left(u_j + \beta \Delta t \Psi_j \right)_D^n \right]$$

with similar ones for departure point equations Need only to specify:

- 1. ∇ -operator in terms of the coordinate metric factors
- 2. Components of gravity vector (only vertical is non-zero)
- **3.** Matrix elements M_{ij}
- 4. Components of Earth's rotation vector in chosen coordinate system

Similar Oblate Spheroidal and Spherical Polar Coordinates

	${\bf Similaroblatespheroids(SOS)}$	Spherical polars
	$(\varepsilon \equiv \text{common eccentricity})$	$(\varepsilon = 0)$
	$(\lambda,arphi,{ t r})$	(λ,ϕ,r)
Coords	$\lambda = ext{longitude}$	$\lambda = $ longitude
	$\varphi = meridional coordinate$	$\phi = $ latitude
	$\mathbf{r} = \text{semi} - \text{major} \text{ axis of spheroid}$	r = radius of sphere
Metric factors	$h_{\lambda} = \left(\frac{\mathbf{r}}{h_{\mathbf{r}}}\right) \cos \phi$ $h_{\varphi} = \left(\frac{\mathbf{r}}{h_{\mathbf{r}}}\right) \left(\frac{\sin \phi \cos \phi}{\sin \varphi \cos \varphi}\right)$ $h_{\mathbf{r}} = \left(1 - \varepsilon^2 \sin^2 \phi\right)^{1/2}$	$h_{\lambda} = r \cos \phi$ $h_{\phi} = r$ $h_{r} = 1$
Gravity	$g = \frac{g_a}{\left(1 - \epsilon^2 \sin^2 \phi\right)^{1/2}} \left(\frac{a}{r}\right)^2$	$g = g_a \left(\frac{a}{r}\right)^2$
Notes	$\phi = \text{geographic latitude} \\ \varphi \to \phi \& \mathbf{r} \to r \text{ as } \varepsilon \to 0$	Geographic latitude = spherical polar latitude when $\varepsilon = 0$



Identical numerical system encompassing hierarchy of model approximations:

- 1. Similar Oblate Spheroidal geopotential approximation
- 2. Deep-atmosphere Spherical geopotential approximation (no latitudinal variation of g)
- 3. Shallow-atmosphere approximation (non-Euclidean geometry)
- 4. Cartesian

Hydrostatic versions apply to (2)-(4)



- Elements of approach form basis of deep-atmosphere, nonhydrostatic, spherical New Dynamics in MetUM
- Next version of dynamical core (ENDGame) will incorporate full functionality
- Allow clean assessment of impact of various approximations



Summary

- Can relax Spherical Geopotential Approximation
 - Retain 2/3rds latitudinal variation of gravity
- Semi-Lagrangian method can be straightforwardly applied to non-spherical coordinate systems
 - Especially when use geographic latitude
- Eulerian component equations can be derived from the rotation matrix forms



Summary

 Matrix view gives insight into the non-Euclidean nature of the shallow-atmosphere approximation

– neglect of great circle curvature

- Therefore desirable to avoid shallow-atmosphere approximation:
 - avoid loss of some (significant) Coriolis terms
 - and avoid non-Euclidean, space-distorting character



Questions?

Further details can be found in:

- Spheroidal coordinate systems for modelling global atmospheres White, Staniforth and Wood 2008 *Q.J.R.Meteorol.Soc.* 134 pp 261-270
- Rotation matrix treatment of vector equations in semi-Lagrangian models of the atmosphere I: Momentum equation Staniforth, White and Wood. 2010 Q.J.R.Meteorol.Soc. 136 pp 497-506
- Rotation matrix treatment of vector equations in semi-Lagrangian models of the atmosphere II: Kinematic equation Staniforth, White and Wood. 2010 Q.J.R.Meteorol.Soc. 136 pp 507-516